

# Going to 2.1 μm for Space Quantum Key Distribution

A. Groulard<sup>1\*</sup>, S. Chaabani<sup>1\*</sup>, M. Zajmulina<sup>2</sup>, J.-B. Lecourt<sup>2</sup>, Y. Hernandez<sup>2</sup> and S. Habraken<sup>1</sup>

<sup>1</sup>Centre Spatial de Liège, Université de Liège, Avenue du Pré-Ailly 19, 4031 Angleur, Belgium

<sup>2</sup>Multitel, Rue Pierre Et Marie Curie 2, 7000 Mons, Belgium

\*Both authors contributed equally to this work. Email: [agroulard@uliege.be](mailto:agroulard@uliege.be), [selim.chaabani@uliege.be](mailto:selim.chaabani@uliege.be).

## Introduction

Quantum Key Distribution (QKD) is a secure cryptographic information transmission method that relies on quantum properties of photons. Space QKD imposes a huge challenge: efficiently crossing the atmosphere. Motivated by a comprehensive model to simulate a satellite-to-ground link, we address this problem with a tuneable heralded single-photon source with signal photons ranging from 1064 to 2300 nm in free-space. The photons are produced via spontaneous parametric down-conversion (SPDC) in a seven-grating periodically poled lithium niobate (PPLN) crystal.

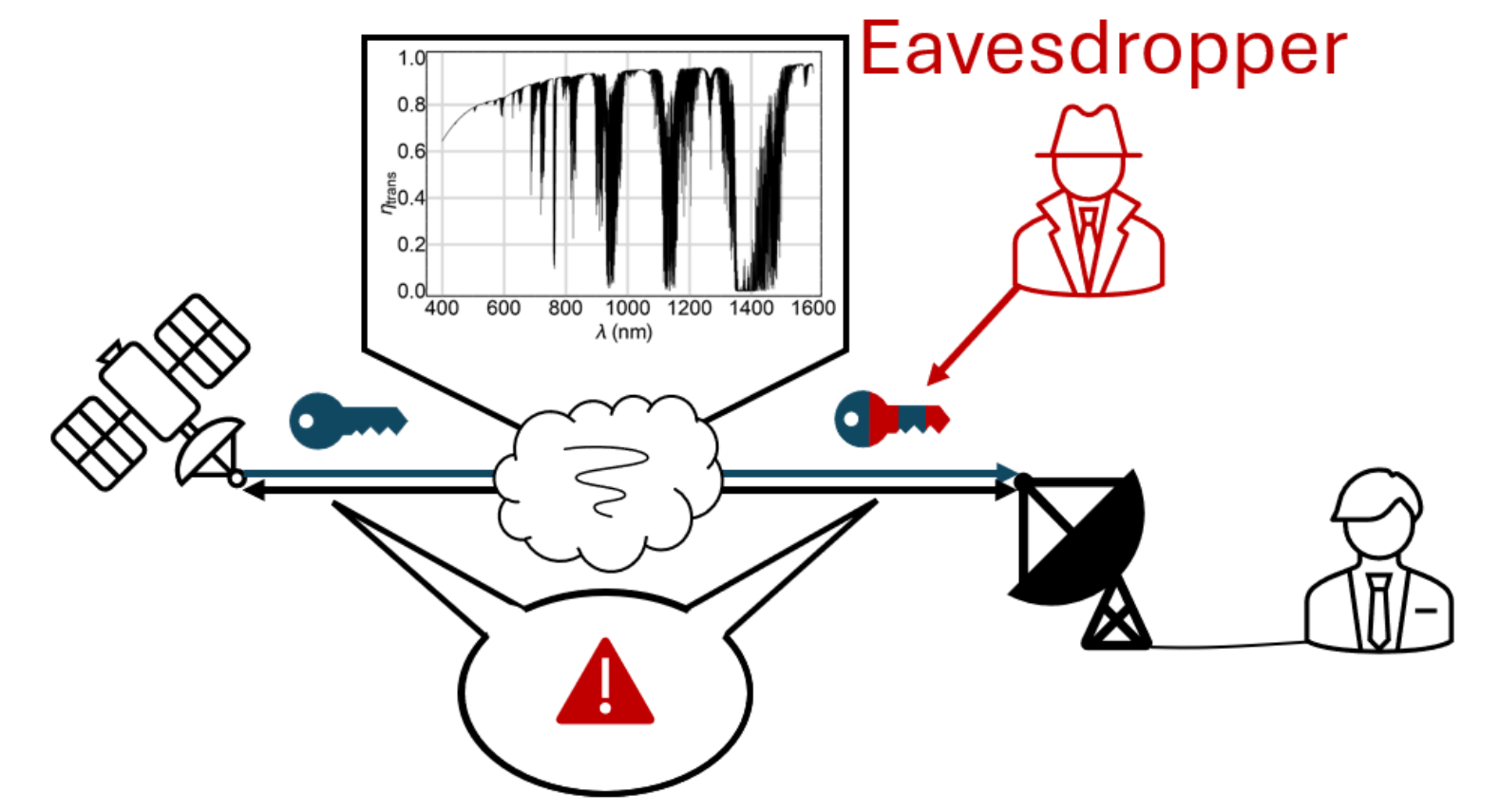


Fig. 1: A QKD scheme for space with the atmospheric transmittance over the wavelength. Adapted from [1].

## Broadband atmospheric model

Including multiple effects, the model predicts the best wavelength in terms of signal-to-noise ratio (SNR) for a given affordable loss (Fig. 2).

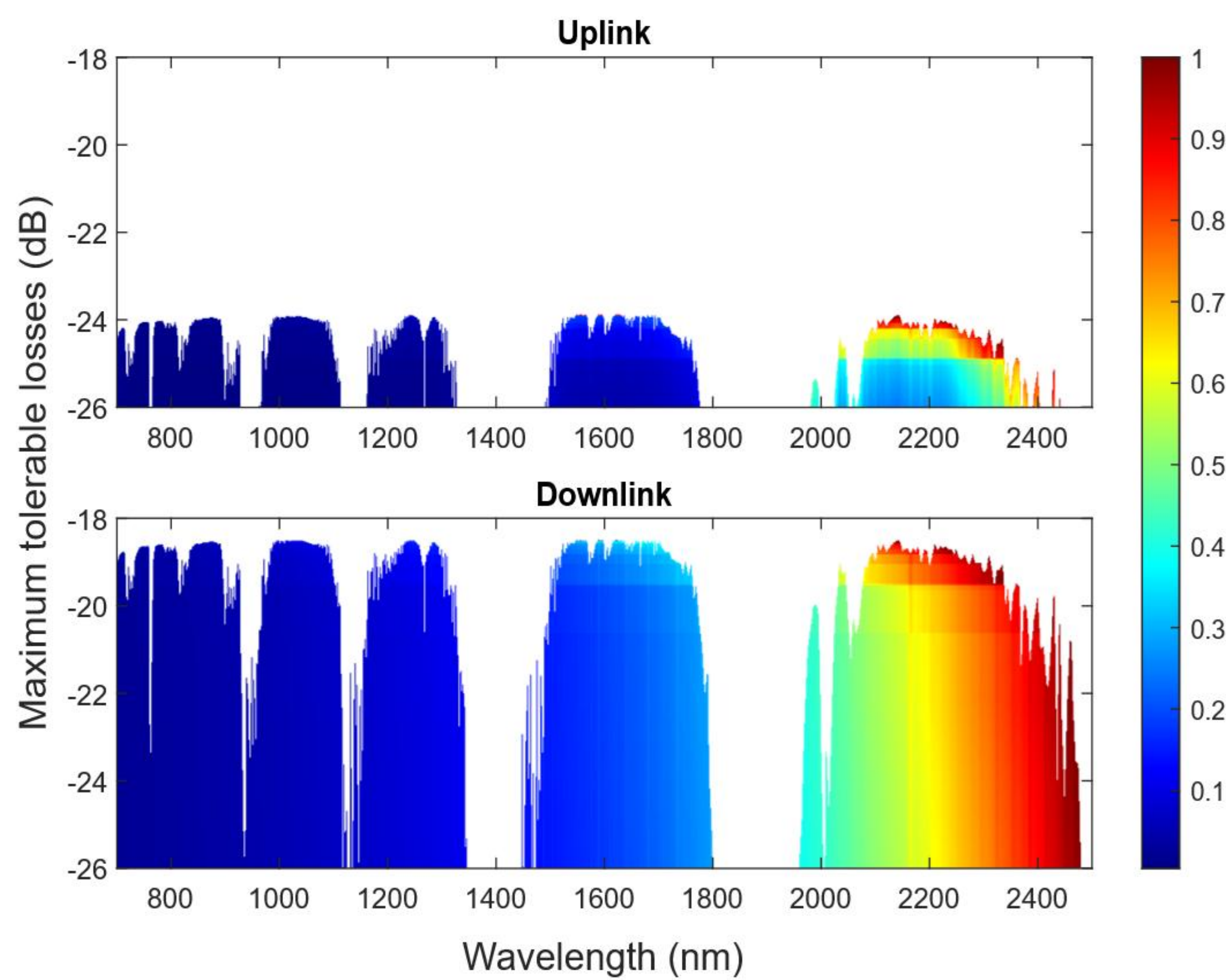


Fig. 2: The outline of the curves is the intrinsic atmospheric loss specific to each wavelength. The colour map indicates the most interesting wavelength provided a given level of loss in the optical link.

**Best atmospheric window: over 2 μm**

## Heralded photon source

With our heralded photon source, we can study many signal wavelengths thanks to the idler photons [2].

Pump laser: mode-locked, 1064 nm, 7,9 ps pulses, 32.45 MHz, 300 pm linewidth

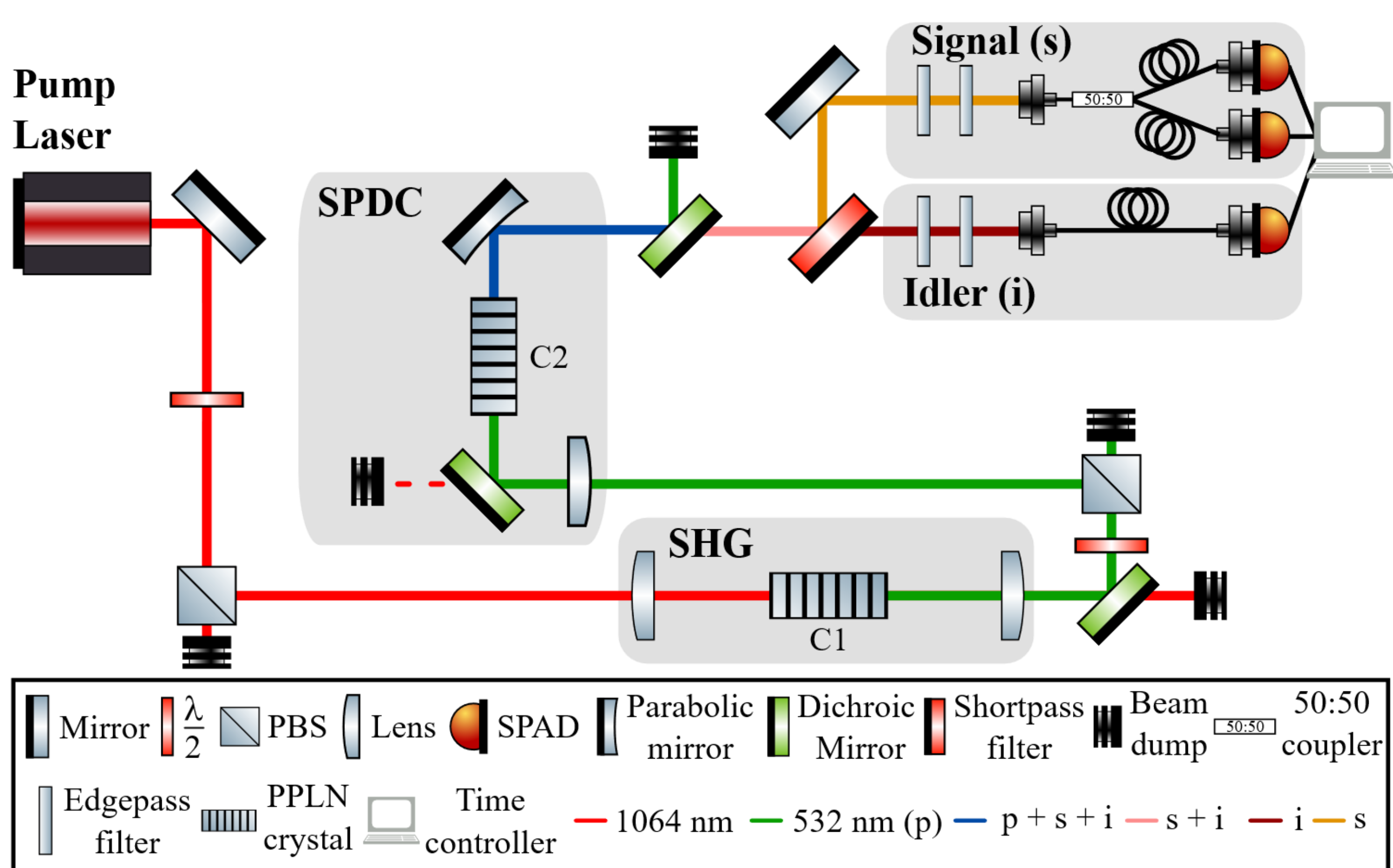


Fig. 3: Optical set-up. PBS: polarisation beamsplitter, SHG: second harmonic generation, C2: multi-grating PPLN crystal, SPDC: (type-0) spontaneous parametric down-conversion.

## Spectral output

$$\frac{1}{\lambda_p} = \frac{1}{\lambda_i} + \frac{1}{\lambda_s}$$

e.g.  $\frac{1}{532} = \frac{1}{805} + \frac{1}{1565}$

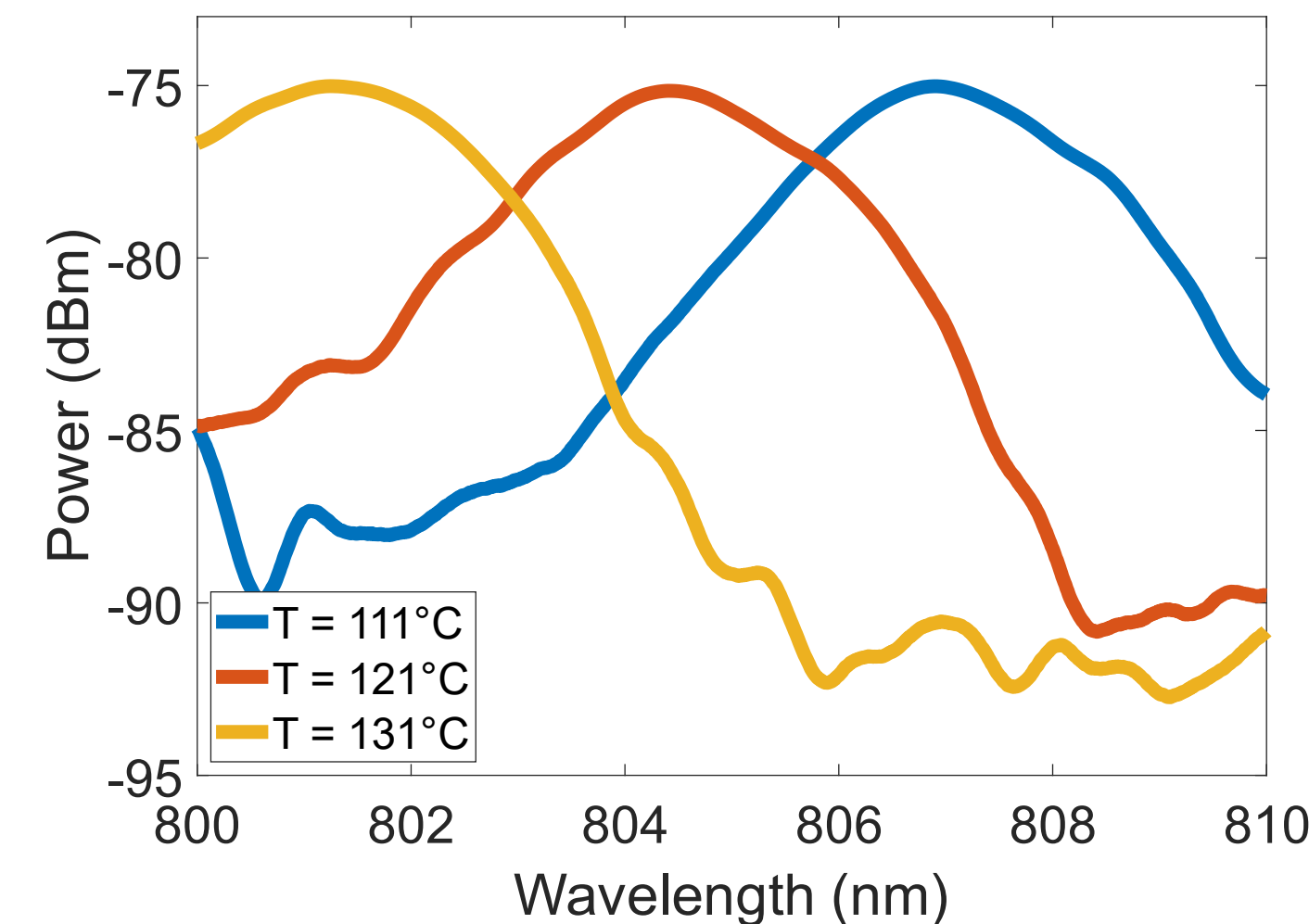


Fig. 5: Spectrum of the idler photon out of the fourth grating of C2 at different temperatures.

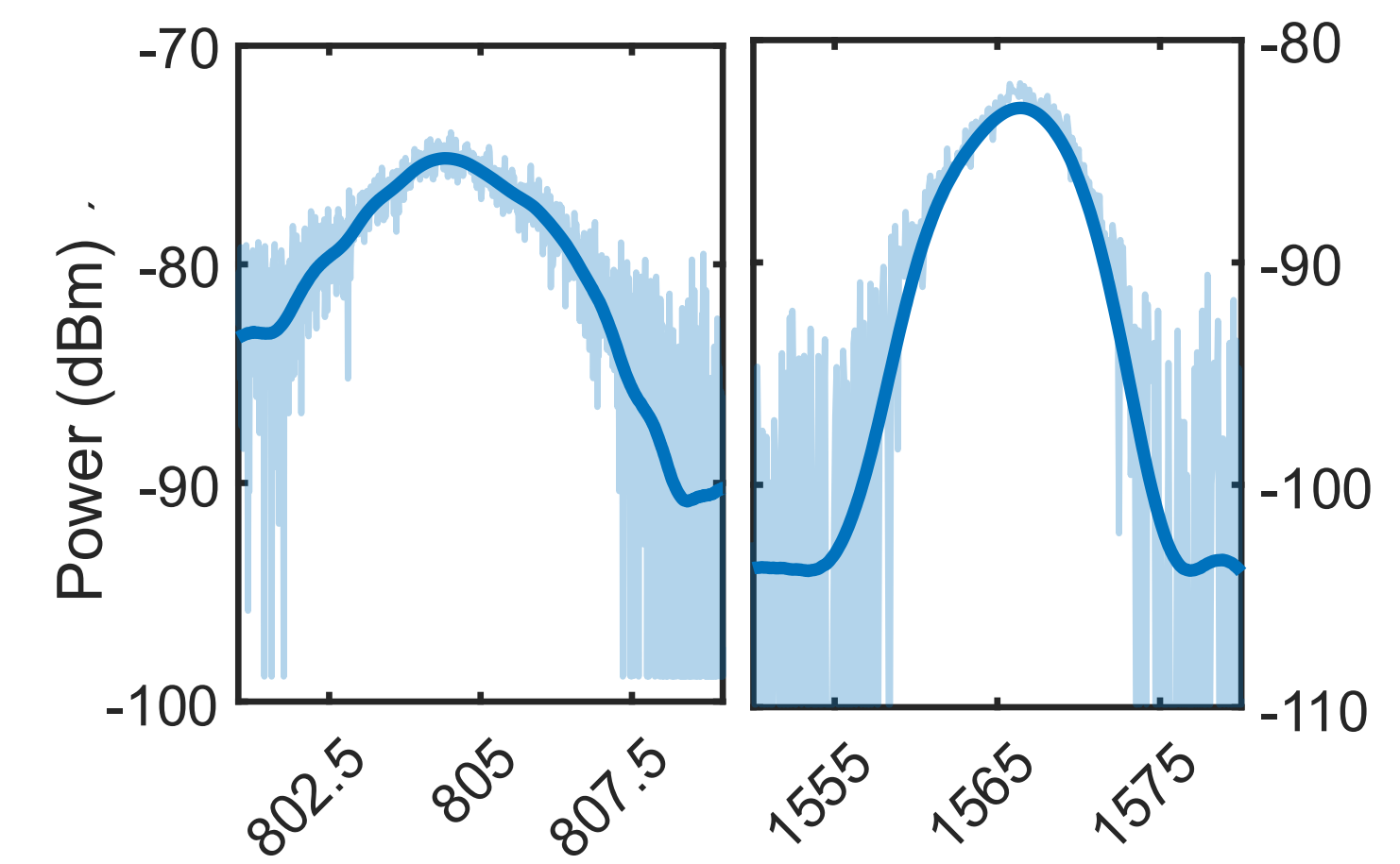


Fig. 4: Complementary idler and signal peaks.

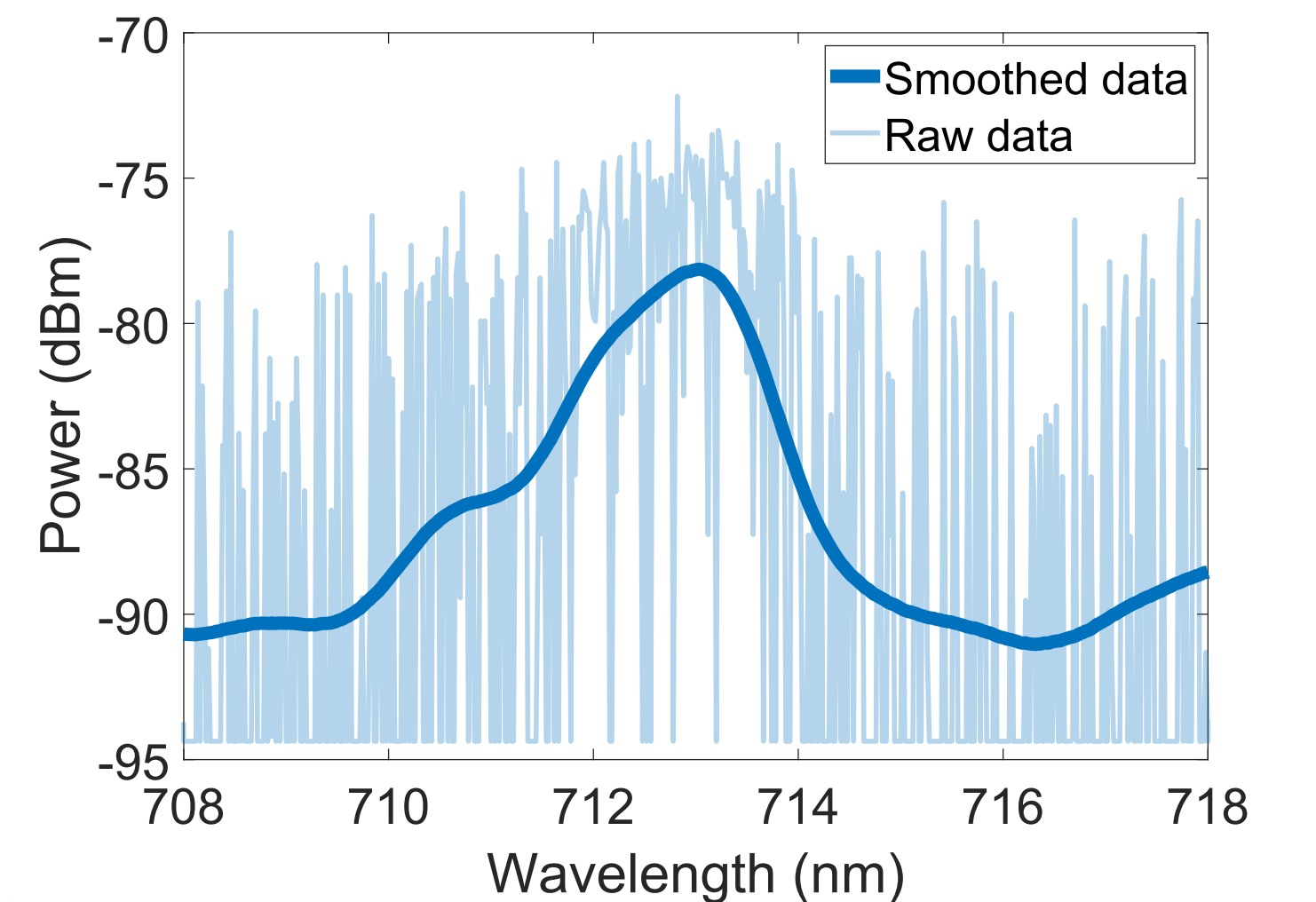


Fig. 6: Spectrum of the idler photon out of the seventh grating of C2 at ambient temperature.

$$\Delta T + \#grating \rightarrow \lambda_s \in [1064 : 2300 \text{ nm}]$$

However, no access to the whole range with our current detectors

## Temporal output

**Heralding of signal photons by idler**

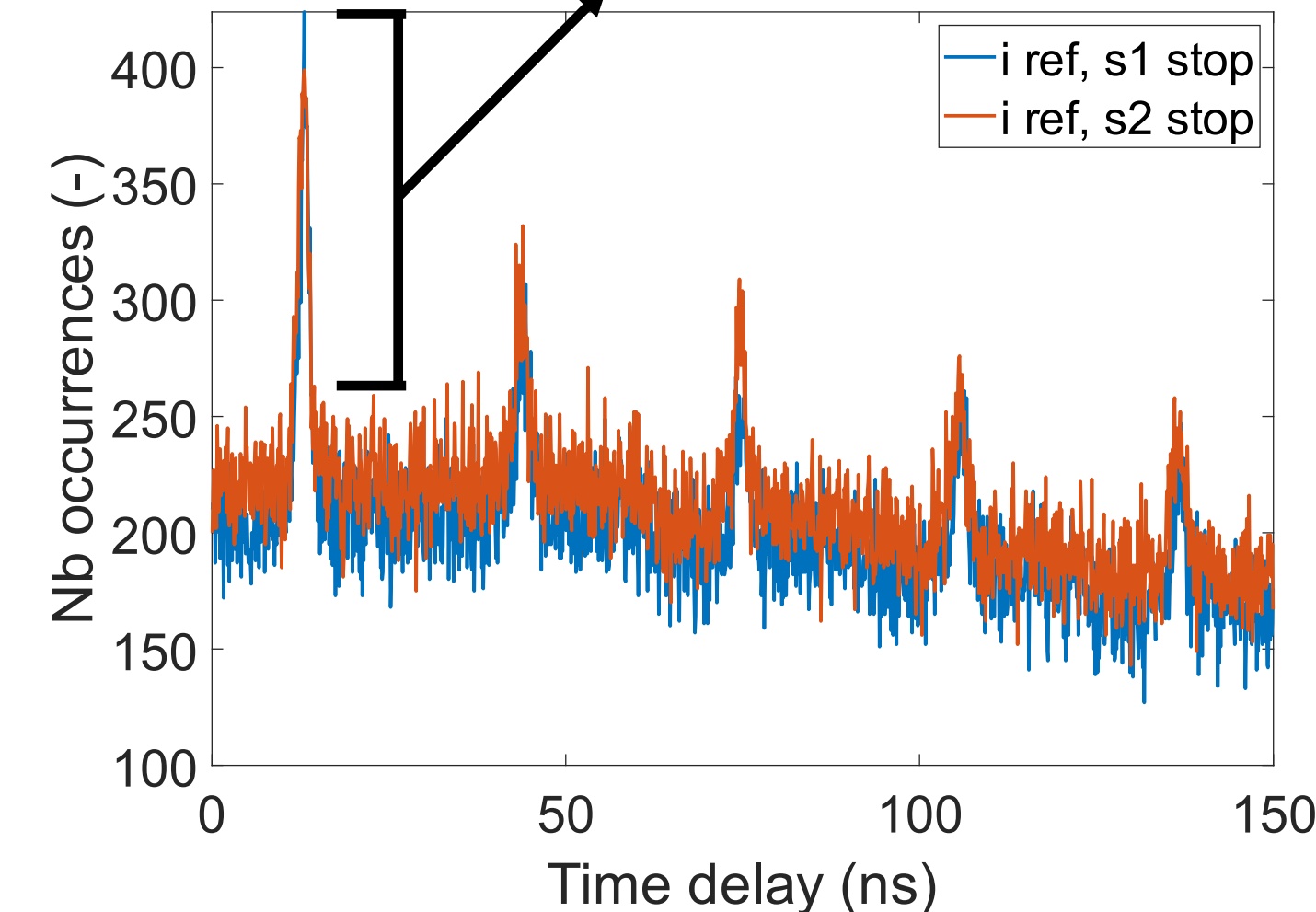


Fig. 7: Time delay between consecutive detections in the idler detector and signal detectors.

**No triple coincidences → single photons**

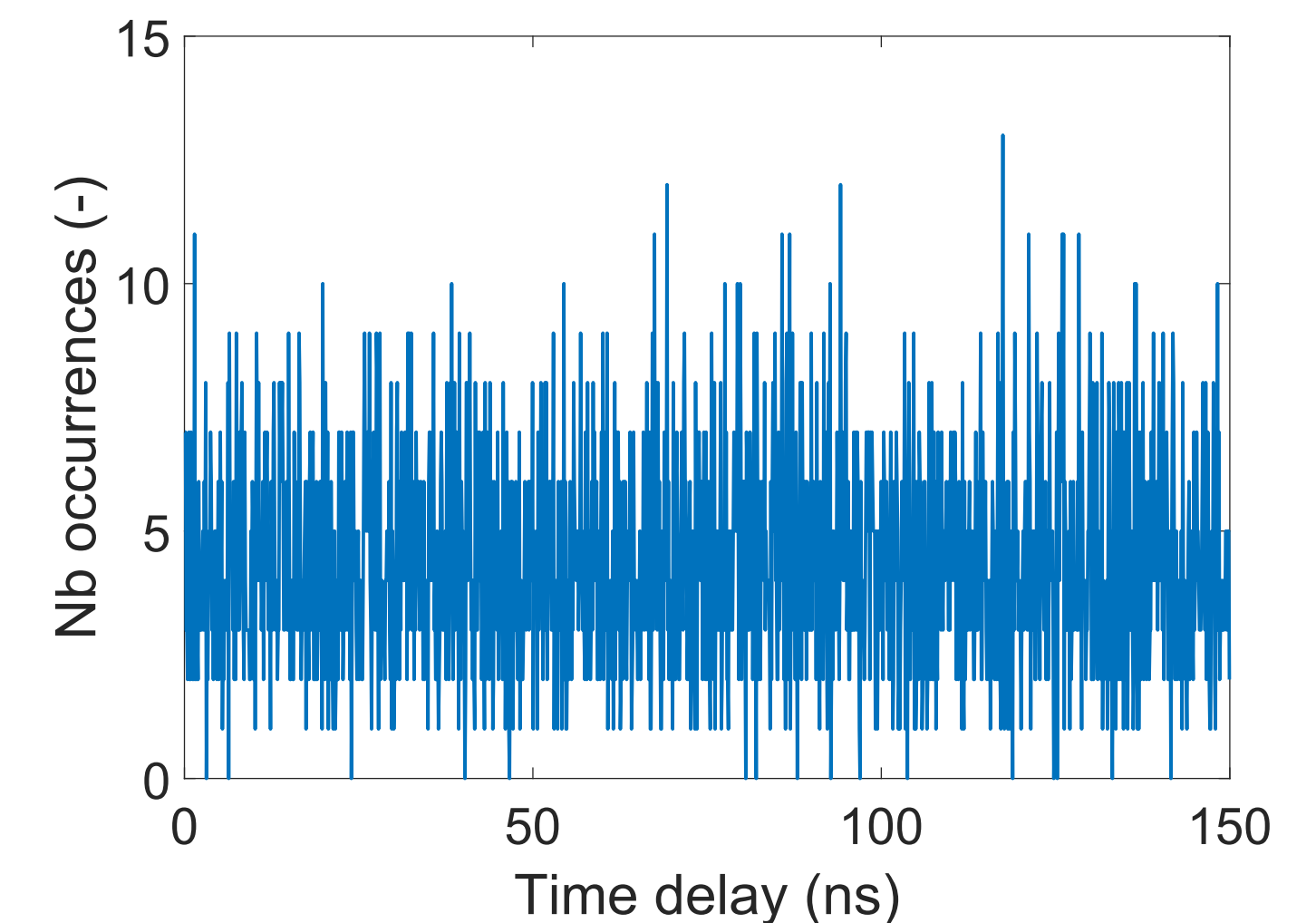


Fig. 8: Time delay between consecutive detections in the signal detectors.

## Conclusions

- ☁ 2 μm atmospheric window is advantageous
- ✓ Can reach it with our source

## Perspectives

- 📷 Include detectors in the model
- 🔗 Optimisation of heralding and single photon rate
- 🎯 Study further wavelengths and focus on 2.1 μm

## References

- [1] R. N. Lanning et al., Phys. Rev. Applied 16, 044027 (2021), doi: 10.1103/PhysRevApplied.16.044027
- [2] S. A. Cappelletto and R. E. Scholten, Eur. Phys. J. Appl. Phys. 41, 181-194 (2008), doi: 10.1051/epjap:2008029.