# 2 thermometer: Application to the melting conditions and mantle

# 3 lithologies of large igneous provinces

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# 27 ABSTRACT

28	The application of the olivine-spinel aluminum exchange thermometer to natural samples is
29	limited by the restricted experimental dataset on which it was calibrated. Here, we present a
30	new dataset of 46 high-temperature crystallization experiments and 21 reanalyzed published
31	experiments, which we used to extend the calibration to higher and lower temperatures. The
32	final calibration dataset spans a range of conditions relevant to crustal and upper mantle
33	processes: 1174–1606 °C, 0.1–1350 MPa, QFM–2.5 to QFM+7.2 (oxygen fugacity, fO <sub>2</sub> ,
34	reported in log units relative to the quartz-fayalite-magnetite buffer, QFM), and 0-7.4 wt.%
35	$H_2O_{melt}$ . We propose three new models. The first is thermodynamically self-consistent,
36	based on spinel Fe, Mg, Al, and Cr compositions and Al exchange between olivine and
37	spinel. The second and third are empirical models that consider fewer elemental exchanges:
38	the second uses only Al exchange and spinel compositions, whereas the third considers
39	olivine-spinel Al and Cr exchange. All models include the modest effect of pressure on
40	olivine-spinel equilibrium chemistry, whereas $fO_2$ and water content have negligible effects.
41	In general, as fewer elements are considered in the olivine-spinel exchange, the fit to
42	experimental data worsens. Conversely, the associated decrease in model complexity
43	improves their robustness against systematic errors when applied to natural crystal pairs: the
44	thermodynamic model may underestimate crystallization temperatures in natural samples
45	due to spinel subsolidus re-equilibration, whereas the empirical models (independent of Fe
46	and Mg in spinel) are less sensitive to re-equilibration but yield temperatures with larger
47	uncertainties. We applied a statistical test to select the most appropriate model for
48	application to natural samples. When applied to lavas from mid-ocean ridges, Iceland, Skye,
49	Emeishan, Etendeka, and Tortugal, our new temperature estimates are 30–100 °C lower than
50	previously proposed. The lower temperature estimates cause a lower mantle melting
51	temperature and significant impacts on the mantle lithology constraints.
52	Keywords: Cr-spinel; Large igneous province; Mantle melting; Olivine; Thermometry

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#### 54 DATA AVAILABILITY

- 55 Additional data is available as supplementary electronic files. All calculations of each model,
- the protocol to choose the results are written in a python script and Excel spreadsheet
- 57 (<u>https://github.com/eazzzon/olspthermo</u>).

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RIGIT

# 68 INTRODUCTION

69	In the mantle, convection and upwelling plumes generate substantial magma volumes in
70	mid-ocean ridge (MOR) systems and can produce oceanic or continental large igneous
71	provinces (LIPs; e.g., Morgan, 1971; White & McKenzie, 1989; Campbell & Griffiths,
72	1990). Buoyantly ascending plume materials melt at relatively shallow depths (Sleep, 1992,
73	1996), eventually producing large volumes of magma through decompression melting of the
74	plume head. Mantle plumes are sustained by temperature-induced density differences (e.g.,
75	Rayleigh-Taylor instability) compared to the ambient mantle (White & McKenzie, 1995),
76	and large thermal anomalies (ca. 100–200 °C; White & McKenzie, 1989) are generally
77	considered requisite for LIP activity. This thermal anomaly is commonly expressed as the
78	excess temperature ( $T_{\rm ex}$ ), which is the difference between the mantle potential temperature
79	$(T_{\rm p},$ the mantle temperature extrapolated along the adiabat at 1 bar without melting;
80	McKenzie & Bickle, 1988) at hotspots and of the ambient mantle (i.e., average MORs, $T_{\rm p} \approx$
81	1330 °C; e.g., Green & Falloon, 2005; Falloon et al., 2007), where melts are generated by
82	near-adiabatic upwelling of the upper mantle (McKenzie & Bickle, 1988).
83	The temperatures of LIP mantle sources have often been estimated by applying
84	olivine-liquid geothermometers (Beattie, 1993; Putirka, 2005, 2008) to picritic basalts
85	containing forsteritic olivine (forsterite content expressed as $Fo = 100 \times molar Mg/[Mg + 1$
86	Fe]) to calculate the olivine-melt equilibrium temperature and then $T_p$ (Putirka <i>et al.</i> , 2007).
87	Other approaches involve modeling mantle melting as a function of pressure and
88	temperature and comparing the modeled primary magmas with those determined from
89	natural (near-) primary melts (Ghiorso & Sack 1995: Ghiorso <i>et al.</i> 2002: Herzberg &
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90	Asimow, 2015; Brown Krein <i>et al.</i> , 2021). However, these models require assumptions of
90 91	Asimow, 2015; Brown Krein <i>et al.</i> , 2021). However, these models require assumptions of primary melt compositions, even though near-primary melts rarely erupt (O'Hara, 1968;
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90 91 92 93	Asimow, 2015; Brown Krein <i>et al.</i> , 2021). However, these models require assumptions of primary melt compositions, even though near-primary melts rarely erupt (O'Hara, 1968; Neave <i>et al.</i> , 2019; Neave & Namur, 2022). Therefore, primary melt compositions are usually estimated by adding or subtracting an olivine component iteratively based on

95	leading to large uncertainties in the FeO and MgO contents of the calculated primary melt
96	(e.g., Herzberg <i>et al.</i> , 2007).

97	The olivine-spinel aluminum exchange thermometer (hereafter OSAT; Wan et al.,
98	2008; Coogan <i>et al.</i> , 2014) is based on the exchange of Al between olivine and spinel,
99	written as $K_{DAl} = [Al_2O_3]_{Ol}/[Al_2O_3]_{Spl}$ (Al_2O_3 concentrations in wt.%). This thermometer
100	obviates the need to assume a melt composition because it determines the crystallization
101	temperature directly from the equilibrium compositions of olivine and coexisting spinel.
102	However, the OSAT was only calibrated at low pressure (0.1 MPa) and a restricted range of
103	dry melt compositions (SiO <sub>2</sub> : 40.5–45.9 wt.%, MgO: 17.5–23.5 wt.%, FeO: 10.6–12.2 wt.%,
104	Al <sub>2</sub> O <sub>3</sub> : 9.0–15.4 wt.%), olivine compositions (Fo <sub>86–100</sub> ), spinel compositions (Cr# = 0–0.69,
105	where $Cr = molar Cr/[Cr + Al]$ , and oxygen fugacities ( $fO_2 = QFM-1.5$ to QFM+0.5, with
106	70% of the experiments performed at QFM-1.5; QFM indicates the quartz-fayalite-
107	magnetite equilibrium). Therefore, its applicability in diverse geological settings,
108	particularly those with Cr-rich spinel in equilibrium with Fo-rich olivine, is questionable; it
109	is unknown how intensive parameters such as pressure, $fO_2$ , and volatile contents affect the
110	OSAT calibration.
111	In this study, we performed and analyzed new experiments, reanalyzed prior
112	experiments, and combined our results with the high-quality measurements of Wan et al.
113	(2008) and Coogan et al. (2014) to extend the OSAT calibration. Our final experimental
114	database spans 0.1–1350 MPa, 1174–1606 °C, and QFM–2.5 to QFM+7.2 (i.e., air), and
115	includes three hydrous experiments containing $3.7-7.4$ wt.% $H_2O_{melt}$ .
116	We also developed a thermodynamic formalism to better understand olivine–spinel
117	Al exchange. Our model suggests that, besides $K_{D_{Al}}$ and spinel Cr#, other components
118	involving Fe and Mg in spinel have a significant impact on the calculated olivine-spinel
119	equilibrium temperature and therefore on the previous OSAT calibration. This model is
120	expected to be the most accurate parameterization for olivine and spinel compositions that
121	quenched rapidly, inhibiting Mg-Fe exchange during cooling. However, this comprehensive

122 thermodynamic model may retrieve systematically low temperature estimates when applied 123 to natural rocks due to subsolidus olivine-spinel Fe-Mg re-equilibration. Therefore, we also 124 report two empirical expressions that may be less prone to underestimating olivine-spinel 125 equilibration temperatures in variably re-equilibrated natural samples, but are less precise 126 and accurate than the thermodynamic model. The first empirical model follows a similar formalism as Coogan *et al.* (2014), only involving  $K_{D_{Al}}$  and spinel Cr#. The second 127 incorporates the exchange of Cr between olivine and spinel  $(K_{D_{Cr}} = [Cr_2O_3]_{Ol}/[Cr_2O_3]_{Spl})$ 128 129 Cr<sub>2</sub>O<sub>3</sub> concentrations in wt.%) which improves the accuracy of this model. We find that pressure has a moderate impact on the temperature calculations in all models, whereas  $fO_2$ 130 131 and water have negligible effects. 132 We applied our new models to re-constrain the crystallization temperature for 133 published olivine-spinel pairs in natural rocks. A multi-component mantle melting model 134 (Phipps Morgan, 2001; Shorttle et al., 2014) is then applied to match crystallization temperature with geochemical constraints to estimate mantle temperatures and mantle 135 136 lithology components (Matthews et al., 2016; 2021). Our results suggest that previous 137 studies using the model of Coogan et al. (2014) may have overestimated the crystallization temperature of olivine by as much as  $\sim$ 30–100 °C for intra-plate LIPs, significantly 138 overestimating  $T_p$  and impacting constraints on the mantle lithology. Based on our results, 139 we provide new constraints on mantle melting conditions and mantle lithologies for MORs, 140 141 Iceland, Skye, and intra-plate LIPs including Emeishan, Etendeka, and Tortugal. 142 143 The olivine-spinel aluminum exchange thermometer (OSAT) 144 The OSAT (Wan et al., 2008; Coogan et al., 2014) determines the equilibrium temperature 145 of olivine-spinel pairs, which corresponds to the minimum saturation temperature of these two phases near the liquidus of their primitive parental melts (e.g., Jennings et al., 2019). 146 147 The crystallization temperature is calculated based on the concentration (in wt.%) of  $Al_2O_3$ 148 in olivine and those of  $Al_2O_3$  and  $Cr_2O_3$  in spinel as:

$$T(K) = \frac{10,000}{0.575(0.162) + 0.884(0.043) \operatorname{Cr}\# - 0.897(0.025) \ln(K_{D_{Al}})}$$
(1)

149	with the standard errors on the determined coefficients reported in parentheses. The slow
150	diffusion rate of $Al_2O_3$ in olivine minimizes the effect of re-equilibration at low
151	temperatures (Spandler & O'Neil, 2010). The OSAT has been widely used to investigate the
152	temperature of formation of terrestrial and extraterrestrial magmas (Gavrilenko et al., 2016;
153	Prissel et al., 2017; Trela et al., 2017; Goltz et al., 2020; Jennings et al., 2020; Matthews et
154	al., 2021; Ramsey et al., 2021; ), including those of LIPs. However, the model has faced
155	criticism for the limited experimental database used for its calibration, which has limited its
156	application. The main issues with the previous calibration are: (1) the range of experimental
157	spinel compositions is quite narrow compared to those observed in natural rocks (see next
158	subsection); (2) the highest experimental temperature in the database (1450 °C) is
159	significantly lower than magmatic temperatures calculated in many natural settings
160	(>1500 °C; e.g., komatiite, Trela et al., 2017; LIPs, Matthews et al., 2021; Martian
161	shergottites, Ramsey et al., 2021), implying that the currently available expression must be
162	extrapolated beyond its calibration range; and (3) high-pressure and hydrous experiments
163	are absent in the calibration database.
164	$\checkmark$
165	Comparison of the original calibration experiments with natural rocks
166	To demonstrate the limitations of the existing olivine-spinel aluminum exchange calibration,

160 To demonstrate the initiations of the existing onvine-spiner administration experiments to those
167 we compare the olivine and spinel compositions used in the calibration experiments to those
168 most common in LIPs.

169To compare spinel compositions, we used the ternary projection of the spinel prism, Al–Cr–170 $Fe^{3+}$  (Fig. 1a). Spinel  $Fe^{3+}$  contents were calculated by charge balance on a four-oxygen171basis (Droop, 1987). Both experimental and natural spinel compositions show a dominant172trend from the Cr apex towards the Al apex. However, spinel in primitive basalts commonly173have compositions plotting far beyond the calibration range (Cr# = 0.23–0.85 in natural174spinel vs. 0–0.69 in the calibration experiments; Fig. 1b). Natural spinel also has higher  $Fe^{3+}$ 

175	contents and occasionally higher $TiO_2$ contents (>5 wt.%; Fig. 1c) compared to the
176	calibration dataset. Therefore, the experimental spinel compositions are only directly
177	comparable to MORB compositions and partially overlap those from Iceland. The chemical
178	composition of igneous spinel is known to be a complex function of intensive parameters
179	(pressure, temperature, $fO_2$ ) and melt composition (Katsura & Ito, 1989; Ballhaus <i>et al.</i> ,
180	1991; Ariskin & Nikolaev, 1996; Kamenetseky et al., 2001). The chemical differences
181	between experimental spinel compositions in the calibration dataset and natural spinel may
182	be due to the narrow $fO_2$ range and/or restricted melt compositions used in the experiments;
183	many of the experiments were performed in the Cr-free simplified CaO-MgO-Al <sub>2</sub> O <sub>3</sub> -SiO <sub>2</sub>
184	(CMAS) system, with pure spinel in equilibrium with forsterite. However, the observed
185	differences could also be partly related to subsolidus Fe-Mg re-equilibration between olivine
186	and spinel in natural rocks.
187	Figure 2 compares olivine $Al_2O_3$ and $Cr_2O_3$ concentrations as a function of Fo
188	content. The restricted experimental conditions also limited the range of experimental
189	olivine compositions to Fo <sub>86-100</sub> (Fo <sub>100</sub> crystalized in CMAS experiments), compared to the
190	much wider range observed in natural olivine (Fo <sub>68–94</sub> , Fig. 2). The Al <sub>2</sub> O <sub>3</sub> (190–1,519 $\mu$ g/g)
191	and $Cr_2O_3$ (165–3,420 µg/g) contents of natural olivine crystals also commonly fall outside
192	the calibration interval (Al <sub>2</sub> O <sub>3</sub> , 450–1,686 $\mu$ g/g; Cr <sub>2</sub> O <sub>3</sub> , 45–2,990 $\mu$ g/g).
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104	EVERIMENTAL OTDATECY AND ANALYTICAL METHODS

### 194 EXPERIMENTAL STRATEGY AND ANALYTICAL METHODS

**195** Selection and preparation of starting materials

To co-saturate olivine and spinel at different temperatures from a wide range of melt
compositions, we used a variety of natural and synthetic starting compositions with 10.1–
34.9 wt.% MgO, 38.0–50.2 wt.% SiO<sub>2</sub>, 3.8–16.3 wt.% Al<sub>2</sub>O<sub>3</sub>, 0.2–6.0 wt.% total alkalis
(Na<sub>2</sub>O + K<sub>2</sub>O) and 0.1–2.0 wt.% Cr<sub>2</sub>O<sub>3</sub>. Details of the different starting compositions are
reported in Table 1 and summarized here.
We used two komatiitic compositions from the Tortugal lava suite, Costa Rica: TO-

202 080514-1 and TO-080514-2 (Trela *et al.*, 2017), here abbreviated as TO1 and TO2,

20	3	respectively. These lavas have >33 wt.% MgO, different $TiO_2$ contents (TO1: 0.79
20	4	wt.%; TO2: 0.89 wt.%) and preserve olivine with high-Cr# spinel inclusions
20	5	(Cr# >0.85).
20	6 2)	Five basaltic compositions cover a range of melt SiO <sub>2</sub> (38.0–50.2 wt.%,) and $Al_2O_3$
20	7	contents (3.8–16.3 wt.%), which may affect the exchange of Al between the liquid,
20	8	olivine, and spinel (Evans et al., 2008; Hanson & Jones, 1998). These five
20	9	compositions include: primitive basalts from Pico, Azores (PI052, van Gerve et al.,
21	0	2021) and Osorno volcano, Chile (OS082; Bechon et al., 2022); two alkali-rich (>5
21	1	wt.%), SiO2-undersaturated basalts from Nyiragongo volcano, DR Congo (Ny17-
21	2	135 and Ny17-161, Molendijk et al., 2023a); and a porphyritic basalt from the
21	3	Nyamuragira volcano, DR Congo (NYA2012-09-018, here abbreviated as NYAM).
21	4 3)	We prepared three synthetic compositions as analogues to TO1, TO2, and Ny17-135
21	5	(Table 1). The objective here was to test the potential effect of relict olivine and
21	6	spinel grains in natural materials with liquidus temperatures too high to completely
21	7	melt in a standard muffle furnace prior to experimental runs in the vertical gas
21	8	mixing furnace (see Experimental methods, below).
21	9 4)	Finally, we synthesized the starting composition of Wan et al. (2008; here named
22	0	2Cr'+1Cr) to check interlaboratory experimental reproducibility and test analytical
22	1	uncertainties under different analytical settings.
22	2 🔨	Natural samples were crushed to finer than 1 $\mu$ m in a ball mill and melted three
22	3 times a	at >1500 °C in a Pt crucible in a muffle furnace: the first melting batch was used to
22	4 saturat	e the Pt crucible and was not kept; the two additional batches were used to ensure the
22	5 compo	sitional homogeneity of the glass. The synthetic compositions were prepared from
22	6 mixtur	es of high-purity oxides and silicates (SiO <sub>2</sub> , TiO <sub>2</sub> , Al <sub>2</sub> O <sub>3</sub> , MnO, MgO, Fe <sub>2</sub> O <sub>3</sub> , CaSiO <sub>3</sub> ,
22	7 AlPO <sub>4</sub>	, Na <sub>2</sub> SiO <sub>3</sub> , K <sub>2</sub> Si <sub>4</sub> O <sub>9</sub> , Cr <sub>2</sub> O <sub>3</sub> , NiO) in appropriate proportions. The silicates (CaSiO <sub>3</sub> ,
22	8 Na <sub>2</sub> SiC	$O_3$ , and $K_2Si_4O_9$ ) were prepared following Zhang <i>et al.</i> (2023).
22	9	Due to the dramatic increase of $Cr_2O_3$ solubility in the melt (>1 wt.%) at high
23	0 temper	catures (>1450 °C; Borisova <i>et al.</i> , 2020) and the potential increase of the $Cr_2O_3$

evaporation rate with increasing temperature (Sossi *et al.*, 2019) in the gas mixing furnace,

most natural and all synthetic compositions were doped with 1-2 wt.% Cr<sub>2</sub>O<sub>3</sub> (compositions

with 'Cr1' suffixes for 1 wt.% and 'Cr2' for 2 wt.% dopant, respectively; Table 1). In some

234 experiments, chromite and forsterite powders were added to change the liquidi of olivine

and spinel (Table 1). All starting powders were mixed and ground in ethanol for one hour

then dried at 120 °C before use in experiments.

237

#### 238 Experimental methods

239 Experiments were conducted in a Nabertherm HTRV 50/150/17 vertical tube furnace at KU Leuven, Belgium, or a GERO HTRV 70-250/18 vertical tube furnace at the University of 240 Liège, Belgium. Approximately 50 mg of starting material was suspended on a 0.2-mm-241 diameter Pt loop. To avoid potential Fe loss, the Pt loops were pre-saturated with the 242 243 experimental material for ~24 h at the same temperature and  $fO_2$  to be used in the experiment, then cleaned with HF. Prior to experiments, all sample pellets were briefly 244 heated in a muffle furnace at 1550 °C to ensure sintering. Experimental samples were then 245 246 suspended for 7-120 h in the hot spot of the vertical furnace (~5 cm in height, where temperature is stable to within  $\pm 3$  and  $\pm 1$  °C in the Nabertherm HTRV 50/150/17 and 247 GERO HTRV 70-250/18 furnaces, respectively). Temperature was measured using a S-type 248 249 (Pt–Pt<sub>90</sub>Rh<sub>10</sub>) thermocouple calibrated against the melting points of Ag and Au. Oxygen 250 fugacity was controlled using mixtures of high-purity CO and CO<sub>2</sub> flushed upward from the 251 bottom of the alumina ceramic furnace tube. Bronkhorst gas-flow controllers were used to 252 set the flow rate at 0.12 cm/s. The accuracy on the  $fO_2$  was checked with a zirconia oxygen 253 sensor. Because it is challenging to produce large spinel and olivine crystals above 1500 °C, 254 in experiments at such high temperatures, the temperature was first oscillated by  $\pm 10$  °C for 255  $\sim$ 1–2 h to increase the sizes of olivine and spinel crystals before holding the experiments 256 isothermally for the remainder of the experiment. All experiments were drop-quenched in 257 water. Experimental run conditions are detailed in Table 2.

#### 258 Selection of previous experiments

259 To expand the experimental database and test the effects of pressure and water content on

- the Al exchange between olivine and spinel, we also selected and reanalyzed 21 published
- 261 experiments, including crystallization experiments on Munro komatiite, Canada (Thy, 1995),
- 262 lunar magma ocean (Charlier *et al.*, 2018), harzburgite melting experiments (Parman &
- 263 Grove, 2004), hydrous melting experiments on komatiite from Commondale Ultramafic
- 264 Suite, South Africa (Barr et al., 2009), and melting experiments on fertile mantle component
- 265 (Mitchell & Grove, 2015). These experiments span a range of temperature between 1215

and 1350 °C, pressure between 0.1 and 1350 MPa,  $fO_2$  from QFM-2.5 to QEM, and between

- **267** 0 and 7.4 wt.%  $H_2O_{melt}$ .
- 268

#### 269 Analytical methods

Analyses were performed using a JEOL JXA-8530F Wavelength Dispersive Spectrometer
(WDS) Field Emission Gun Electron probe micro-analyzer (EPMA) at the Department of
Material Engineering, KU Leuven, Belgium, and a Cameca SX-100 Electron Microprobe at
the Department of Earth Sciences, University of Cambridge. Major and trace elements were
measured with different calibrations and analytical setups detailed in the Supplementary
data and summarized here. Major element concentrations in olivine and spinel were
generally within 98% accuracy of the reference values.

277 Experimental products were mounted in high-purity epoxy and polished for analysis. 278 Electron probe micro-analyzer measurements of trace elements in olivine, especially Cr and 279 Al, are known to suffer from secondary fluorescence from surrounding spinel or melt 280 (Llovet et al., 2012, Gavrilenko et al., 2023; Llovet et al., 2023). We used FANAL (Llovet 281 et al., 2012) and PENEPMA (Llovet & Salvat, 2017), Monte Carlo simulation tools to 282 model coupled electron-photon transportation, and simulated analyses of San Carlos olivine 283 and VG-2 glass standards. The results show that measuring olivine compositions at least 5 284  $\mu$ m (based on FANAL) and 15  $\mu$ m (based on PENEPMA) inward from the phase margin 285 dramatically reduces the influence of secondary fluorescence (Fig. S1). This analytical



- 313 X-ray compositional maps were acquired following the procedure of Shea et al. (2019), but
- using a 20 kV and 300 nA electron beam. 314

315	At the University of Cambridge, we followed a similar procedure using a focused
316	beam operating at 15 kV and 100 nA beam to measure Al, Cr, and P in olivine. Detection
317	limits were 28 $\mu$ g/g for Al <sub>2</sub> O <sub>3</sub> , 117 $\mu$ g/g for P <sub>2</sub> O <sub>5</sub> , and 87 $\mu$ g/g for Cr <sub>2</sub> O <sub>3</sub> . We used the
318	MongOl sh11-2 olivine ( $n = 8$ ) as a secondary standard and obtained $487 \pm 10 \ \mu g/g \ Al_2O_3$
319	(1 $\sigma$ ), 228 ± 40 µg/g P <sub>2</sub> O <sub>5</sub> (1 $\sigma$ ), and 209 ± 20 µg/g Cr <sub>2</sub> O <sub>3</sub> (1 $\sigma$ ), again consistent with the
320	reference values (463 $\pm$ 18 µg/g Al <sub>2</sub> O <sub>3</sub> , 152 $\pm$ 8 µg/g P <sub>2</sub> O <sub>5</sub> , 182 $\pm$ 6 µg/g Ct <sub>2</sub> O <sub>3</sub> , uncertainties
321	in $2\sigma$ ; Batanova <i>et al.</i> , 2019).
322	
323	Interlaboratory experimental and analytical uncertainties
324	To calibrate our extended OSAT with as many relevant experiments as possible, we chose to
325	include the experimental results of Wan et al. (2008) and Coogan et al. (2014), which we
326	could not reanalyze for this study. It is therefore important to critically assess whether we
327	can reproduce their phase equilibria observations and reproduce their chemical
328	measurements, including minor elements such as Al in olivine. Therefore, we reproduced
329	two experiments from Wan et al. (2008): w83 (~QFM-1.6, 1250 °C), reproduced here in
330	experiment NAB89U, and w71 (~QFM-1.6, 1300 °C), reproduced here in experiment
331	NAB90U. We chose these two relatively low-temperature experiments because it is more
332	challenging to attain equilibrium under those conditions.
333	Experiments NAB89U and NAB90U produced phase equilibria (liquid, olivine, and
334	spinel) and compositions similar to those reported by Wan et al. (2008). Importantly,
335	$\ln(K_{D_{Al}})$ values were almost identical: $-6.16 \pm 0.04$ (1 $\sigma$ , uncertainty calculated from error
336	propagation with a Monte Carlo algorithm) in NAB89U vs. $-6.14 \pm 0.05$ (1 $\sigma$ ) in w83, and
337	$-5.80 \pm 0.03$ (1 $\sigma$ ) in NAB90U vs. $-5.90 \pm 0.05$ (1 $\sigma$ ) in w71. Spinel compositions were also
338	identical: $\text{Fe}^{3+}/\text{Fe}^{2+}$ in spinel was $0.16 \pm 0.05 (1\sigma)$ in NAB89-U vs. $0.17 \pm 0.09 (1\sigma)$ in w83,
339	and $0.21 \pm 0.03$ (1 $\sigma$ ) in NAB90-U vs. $0.21 \pm 0.07$ (1 $\sigma$ ) in w71. This suggests that

340 interlaboratory analytical and experimental uncertainties are minor, and further that we can

include the published experimental results from Wan *et al.* (2008) and Coogan *et al.* (2014)

- 342 in our new OSAT calibration database. This also suggests that our experiments attained
- 343 equilibrium (see next subsection).
- 344

#### 345 Attainment of equilibrium

346 The attainment of equilibrium between melt, olivine, and spinel, especially for slowly diffusing minor elements, is critical to our study. The considerably long durations of our 347 348 experimental runs (7-120 h), most much longer than in high-temperature experiments in the 349 previous calibration (Wan et al., 2008) or other high-temperature experimental studies (e.g., Tuff et al., 2005; Matzen et al., 2011; Krasheninnikov et al., 2017; Koshlyakova et al., 350 2022), were chosen to ensure that equilibrium was attained. We note that a few high-351 352 temperature experiments (>1500 °C, Table 2) lasted only a few hours (≥7 h), but that the fast elemental diffusion at such high temperatures and the low crystallinity of those 353 354 experiments suggest that equilibrium was attained even in those shorter runs. 355 We also evaluated the attainment of equilibrium based on compositional and 356 textural characteristics. Olivine and spinel were generally euhedral, suggesting they formed 357 at equilibrium. The experimental charges also show homogeneous backscattered electron intensities for melts and minerals (Fig. 3). Fe-Mg partition coefficients between olivine and 358 melt  $(K_{\text{DFe}_{7}\text{Mg}}^{\text{Ol-melt}})$  range from 0.26 to 0.34 (see in the Supplementary data), consistent with 359 360 previous experiments (e.g., Roeder & Emslie, 1970; Toplis & Carroll, 1995) and thermodynamic models (e.g., Toplis, 2005; Blundy *et al.*, 2020). In general,  $K_{\text{DFe-Mg}}^{\text{Ol-melt}}$  is 361 362 positively correlated with the melt  $SiO_2$  content, which supports the impact of melt silica 363 activity on Fe-Mg exchange between olivine and melt (Gee & Sack, 1988; Toplis, 2005; 364 Koshlyakova et al., 2022). In particular, experiments using silica-undersaturated and alkalirich compositions have  $K_{\text{DFe-Mg}}^{\text{Ol-melt}} < 0.27$  (0.22–0.26), consistent with experimental results 365 366 on similar compositions (Gee & Sack, 1988; Molendijk et al., 2023b). In contrast, high-

367	temperature experiments with quench textures occasionally have $K_{\text{DFe-Mg}}^{\text{Ol-melt}} = 0.34-0.35$ ,
368	beyond the range $0.30 \pm 0.03$ ; such high values may have resulted from analytical
369	heterogeneity due to the glass analyses involving mixtures of quench crystals and interstitial
370	melt (Mitchell & Grove, 2015).
371	Trace element (Al, P, Cr) partitioning in olivine may be affected due to the
372	formation of a boundary layer enrichment during initially rapid crystal growth (e.g., Shea et
373	al., 2019; Lang et al., 2022). This means that the above indicators of equilibrium (Fe-Mg
374	exchange and textural constraints) may not reflect trace element equilibration. Therefore, we
375	measured rim-to-core profiles comprising $6-7$ spot analyses (ca. $5-6 \mu m$ spacing) per
376	profile in one olivine crystal from each experiment to check the homogeneity of their trace
377	element distributions. The profiles generally show very small deviations (~50 $\mu$ g/g, 1 $\sigma$ ) for
378	Al <sub>2</sub> O <sub>3</sub> and even less variability for other trace elements (see Supplementary material,
379	Section 2). We also acquired EPMA X-ray compositional maps of experiments treated with
380	thermal oscillation and experiments that were run at more than 200 °C below their liquidi.
381	Chemical maps do not show any apparent P, Al, or Cr zoning (see Supplementary material,
382	Section 2). Relict phases were occasionally observed in the cores of olivine crystals in
383	experiments performed on compositions with high liquidus temperatures. The uniform
384	compositions of the crystal rims, however, suggest that the early disequilibrium did not
385	preclude the attainment of equilibrium between the outer part of the crystals and the adjacent
386	melt. Experiments conducted at more than 300 °C below the liquidus typically produced
387	small olivine crystals (~10 $\mu$ m in length) with dendritic P and Al zoning patterns. These
388	experiments also produce olivine populations with distinct high and low Al <sub>2</sub> O <sub>3</sub>
389	concentrations; we used the low-concentration population in this study because it may
390	represent relaxation and equilibration during equilibrium crystal growth (Shea et al., 2019).
391	
392	RESULTS

Experimental results

more than 300 °C below the liquidus also contain clinopyroxene. Compositions of the

396 phases are given in supplementary dataset (Table S1).

397

- 398 Melt
- 399 Melt compositions are plotted as functions of MgO (in wt.%) and temperature in Figs. 4 and
- 400 S3, respectively. The experimental melts span a wide range of compositions, with 7.0-31.5
- 401 wt.% MgO, 38.9–54.6 wt.% SiO<sub>2</sub>, 6.0–12.8 wt.% FeO, 3.9–17.3 wt.% Al<sub>2</sub>O<sub>3</sub>, 0.0–0.9 wt.%
- 402  $Cr_2O_3$ , and 0.1–4.6 wt.% Na<sub>2</sub>O + K<sub>2</sub>O. Melt Al<sub>2</sub>O<sub>3</sub> contents are anticorrelated with melt

403 MgO content, whereas melt  $Cr_2O_3$  contents are positively correlated with melt MgO content

- 404 and temperature, consistent with the higher solubility of  $Cr_2O_3$  in the melt at higher
- 405 temperature (Borisova *et al.*, 2020).

406

407 *Olivine* 

408 Olivine crystals are euhedral and ranged in size from  $\sim 10 \mu m$  to  $>100 \mu m$  (Fig. 3). Olivine Fo contents range from Fo<sub>84</sub> to Fo<sub>96</sub> (Fig. 2). Minor and trace elements span a wide 409 compositional range: 0.1-1.3 wt.% CaO, 362-2118 µg/g Al<sub>2</sub>O<sub>3</sub>, 411-3706 µg/g Cr<sub>2</sub>O<sub>3</sub>, and 410 411 below the detection limit to 511  $\mu$ g/g P<sub>2</sub>O<sub>5</sub>. Al<sub>2</sub>O<sub>3</sub> and Cr<sub>2</sub>O<sub>3</sub> contents are weakly correlated 412 with Fo content (Fig. 2). Olivine  $Cr_2O_3$  content is positively correlated with melt  $Cr_2O_3$ 413 content (Fig. S4b). A broad positive correlation is also observed between olivine and melt 414 CaO contents (Fig. S4c), and the highest olivine CaO contents occur in experiments on 415 SiO<sub>2</sub>-undersaturated compositions, consistent with previous findings (Gee & Sack, 1988; 416 Molendijk et al., 2023b). No correlation is apparent between olivine and melt Al<sub>2</sub>O<sub>3</sub> or P<sub>2</sub>O<sub>5</sub> 417 contents (Fig. S4a, d, respectively).

418

419 Spinel

420 Spinel crystals range in size from  $\sim 1 \mu m$  to  $\sim 40 \mu m$  (Fig. 3). Spinel compositions are shown 421 in Figs. 1 and in Fig. S5 as functions of melt composition. Spinel Cr# ranges from 0.40 to

422	0.88, Mg# = molai	$Mg/(Mg + Fe^{2\tau})$	)] from 0.59 to 0.91	, Fe <sup>3+</sup> /Fe <sup>2+</sup> fro	m 0.08 to 2.48, the

- 423 ratio of  $\text{Fe}^{3+}$  to the sum of trivalent cations,  $\text{Fe}^{3+}/(\text{Fe}^{3+} + \text{Cr} + \text{Al})$ , from 0.02 to 0.22, and
- 424 TiO<sub>2</sub> content from 0.5 to 5.5 wt.%. Experiments NAB01-F and NAB02-F, performed on
- 425 SiO<sub>2</sub>-undersaturated and alkali-rich compositions, show the highest spinel TiO<sub>2</sub> contents (>5
- 426 wt.%), which correspond to the highest melt  $TiO_2$  contents (>3 wt.%; Fig. S5). Spinel FeO
- 427 and  $Cr_2O_3$  contents are only weakly correlated with melt FeO and  $Cr_2O_3$  contents (Fig. S5).
- 428 A more significant correlation is observed between spinel and melt  $Al_2O_3$  contents (Fig. S5),
- 429 consistent with previous observations (Borisova et al., 2020; Sack & Ghiorso, 1991b;
- 430 Ariskin & Nikolaev, 1996; Poustovetov & Roeder, 2001).
- 431

#### 432 Summary of remeasured prior experiments

- 433 Melt compositions in the reanalyzed published experiments and calibrations contain 8.4–
- 434 25.0 wt.% MgO, 42.0–55.6 wt.% SiO<sub>2</sub>, 0–17,5 wt.% FeO, 8.3–21.6 wt.% Al<sub>2</sub>O<sub>3</sub>, and 0–0.54
- 435 wt.%  $Cr_2O_3$ . Olivine has Fo<sub>75-100</sub> and contain 0–1.4 wt.% CaO, 183–1,686 µg/g Al<sub>2</sub>O<sub>3</sub>, 0–
- 436 2,990  $\mu$ g/g Cr<sub>2</sub>O<sub>3</sub>, and 0–2,338  $\mu$ g/g P<sub>2</sub>O<sub>5</sub>. Spinel has Cr# = 0–0.84, Mg# = 0.49–1,
- 437  $Fe^{3+}/Fe^{2+} = 0.08-0.71$ ,  $Fe^{3+}/(Fe^{3+} + Cr + Al) = 0-0.17$ , and contains 0-1.38 wt.% TiO<sub>2</sub>.
- 438

#### 439 Evaluating the model of Coogan *et al.* (2014)

440 We assessed the QSAT of Coogan et al. (2014), which incorporates the earlier experiments 441 of Wan et al. (2008), by applying it to our updated experimental dataset (Fig. 5a). The 442 results show that the model of Coogan et al. (2014) significantly overestimated temperatures, 446 with a maximum deviation ( $\Delta T$  [°C] = predicted temperature – measured temperature) of 444 197 °C, an average deviation of 44.7  $\pm$  46.6 °C (1 $\sigma$ ), and a root-mean-square error (RMSE) 445 of  $\pm 64.3$  °C. These errors further demonstrate the need for a new calibration covering a 446 broader temperature range and accounting for additional factors affecting olivine-spinel Al 447 exchange, including spinel composition, fO<sub>2</sub>, pressure, and melt H<sub>2</sub>O content. 448

#### 449 **REGRESSION AND THERMODYNAMIC MODELS**

- 450 To improve the reliability of the OSAT for natural samples, we here develop a
- 451 thermodynamic formalism for Al exchange between olivine and spinel and two empirical
- 452 regression models. We applied a Monte Carlo cross-validation (MCCV; e.g., Xu & Liang,
- 453 2001) algorithm to estimate the uncertainties of regression models.
- 454 **Regression strategy and models**
- 455 In our regression models, we first split the experimental dataset into a training dataset (n
- 456 95) used for regression and a test dataset including 21 experiments not included in the
- 457 calibration to test the accuracy of the regression (details of training and test dataset are given
- 458 in the supplementary dataset, Tables S2, 3). The standard error estimation (SEE) on the
- 459 training dataset and RMSE on the test dataset were calculated accordingly. Then, to assess
- 460 model uncertainties and eliminate sampling bias, we used a MCCV algorithm. In each
- 461 calculation, we randomly divided the experimental dataset into a 5:1 ratio of training and
  462 test datasets, and then performed 10-fold cross-validation on the training dataset to estimate
  463 SEE<sub>MCCV</sub> on the training dataset and then RMSE<sub>MCCV</sub> on the test dataset, which is also the
  464 recommended error on the thermometric model. Detailed information on the training and
- 465 test datasets, as well as the MCCV results, are given in the Supplementary data (Tables S2466 and S3; Section 3).
- 467

#### 468 Thermodynamic formalism

To better understand the Al exchange between olivine and spinel, we here follow a
thermodynamic formalism. Coogan *et al.* (2014) argued that the most plausible mechanism
for incorporating Al into olivine in a Cr-spinel-bearing system is:

$$(MgAl^{TET})Al^{OCT}O_4^{Spl} = (MgAl^{TET})Al^{OCT}O_4^{Ol}$$
(2)

which we follow in our thermodynamic framework (see details in Appendix A). For
simplicity, in the following we assume thermodynamic ideality in the incorporation of Al in
olivine, i.e., the excess energy caused from site ordering of Al is negligible. This assumption
is based on the following considerations: 1) trace element incorporation is usually

- 476 considered as a dilution so the internal energy can be calculated as a simplified pure
- 477 component (Ganguly, 2008); 2) to the best of the authors' knowledge, there are currently no
- 478 established models to accurately differentiate or calculate the coordination of different Al
- 479 species in olivine and no calibration for the interaction parameters between sites. Al may
- 480 also coordinate with a vacant site (Jollands et al., 2018; Shea et al., 2019; Lang et al., 2022);
- 481 3) established thermodynamic models for other trace elements in olivine (e.g., Ni) with
- 482 similar assumption of ideality have shown a minimal influence on the internal energy (e.g.,
- 483 Matzen et al., 2013). We therefore believe our thermodynamic framework remains valid

484 with this assumption. Al substitution in spinel is however more complicated due to its high

- 485 Al abundance, site ordering, and the known non-ideal mixing behavior of spinel components.
- 486 We apply the thermodynamic framework of Sack (1982) and Sack & Ghiorso (1991a,
- 487 1991b), and express the internal energy of spinel by five independent compositional
- 488 variables  $(X_i)$  (see details in Appendix A; Sack & Ghiorso, 1991b).
- 489 After organizing the thermodynamic framework, olivine–spinel Al exchange can be
- 490 described as (see Appendix A):

$$\ln\left(\frac{X_{\text{Al}_2\text{O}_3^{\text{Ol}}}}{X_{\text{Mg}^{2+}}^{\text{TET}}X_{\text{Al}^{3+}}^{\text{OCT}}}\right) + b = \frac{\varphi^{\text{Spl}}}{c_0 \times T}$$
(3)

491 which can be rearranged as:



$$\varphi^{\text{Spl}} = c_1 X_2 + c_2 X_2^2 + c_3 X_2^1$$

$$+ c_4 (1 - X_2) (1 + X_4 - X_2)$$

$$+ c_5 (1 - X_2) (X_3) + c_6 (1 - X_2) (X_4) + c_7 (1 - X_2) (X_5)$$

$$+ c_8 X_3 (X_3 + X_4 + X_5)$$

$$+ c_9 X_4 (X_3 + X_4 + X_5)$$

$$+ c_9 X_4 (X_3 + X_4 + X_5)$$

$$- c_{11} (X_3) (X_4) - c_{12} (X_3) (X_5) - c_{13} (X_3) (X_5) + c_{14}$$
494 Here, *R* is the gas constant (8.314 J mol<sup>-1</sup> K<sup>-1</sup>),  $X_{Mq^{3-1}}^{\text{ET}}$  is the mole fraction of Mg in the  
spinel tetrahedral site,  $X_{Al^{1+}}^{\text{OCT}}$  is the mole fraction of Al in the spinel octahedral site, and *T* is  
temperature in Kelvin. In Eq. 4,  $X_i$  are independent compositional variables used to describe  
spinel composition (Si, Ti, Fe, Mg, Cr, Al, and Mn; Sack & Ghiorso, 1991b); they are listed  
in Table A1. Other parameters ( $b_1 c_{0-1} c_{0+1}$ ) are regression coefficients fitted to the  
experimental dataset through a polynomial curve-fitting approach. In Eq. 3, because the  
term  $X_{Mg^{1+1}}^{\text{ETT}} X_{Al^{1+1}}^{\text{OT}}$  is proportional to spinel Al<sub>3</sub>O<sub>3</sub> content, we can simplify the equation by  
Tr (K) =  $\frac{10,000 \times \varphi^{\text{Spl}}}{-0.168 \times (\ln K_{DAI} + 0.654)}$  (5)  
Tr (K) =  $\frac{10,000 \times \varphi^{\text{Spl}}}{-0.168 \times (\ln K_{DAI} + 0.654)}$ 

R

 $+ 0.009(1 - X_2)(X_3) - 2.492(1 - X_2)(X_4) + 0.065(1 - X_2)(X_5)$ 

- $-0.031X_3(X_3 + X_4 + X_5)$
- $-4.141X_4(X_3 + X_4 + X_5)$

$$-0.428X_5(X_3 + X_4 + X_5)$$

$$+ 4.637(X_3)(X_4) + 0.054(X_3)(X_5) + 10.803(X_4)(X_5) + 2.714$$

The regression result for Eq. 5 is shown in Fig. 5b and Table A2: the SEE and  $r^2$  values for the training dataset are ±20.2 °C and 0.97, respectively, RMSE on the test dataset is 29.0 °C, and the *p*-value of the fit, where a no correlation alternative null hypothesis being tested against, is 1.11e–16. The MCCV result shows that the median SEE<sub>MCCV</sub> and  $r^2_{MCCV}$  values for the training dataset are 20.2 °C and 0.97, respectively, and the median RMSE<sub>MCCV</sub> value for the test dataset is ~23.9 °C.

## 510 Empirical expressions

By using our expanded experimental dataset, we first recalibrated an empirical model
similar to the regression format of Coogan *et al.* (2014) but with unweighted multiple linear
regression (see supplementary material for a comparison with a weighted regression method,
Section 4):

$$T (K) = \frac{10,000}{0.740(0.317) + 1.144(0.136)Cr\# - 0.865(0.046)lnK_{DAl}}$$
(7)

The regression result of this new thermometer is shown in Fig. 5c and Table A2: the SEE and  $r^2$  values for the training dataset are ±44.8 °C and 0.83, respectively, RMSE on the test dataset is 34.4 °C, and the *p*-value of the fit is 2.88e-36. The MCCV result shows that the

(6)

518 median SEE<sub>MCCV</sub> and  $r^{2}_{MCCV}$  values for the training dataset are 42.1 °C and 0.92,

respectively, and the median  $RMSE_{MCCV}$  value for the test dataset is 43.3 °C.

- 520 Considering that Cr may have a coupled substitution with Al when entering olivine
- 521 to form a (MgAl)Cr<sub>2</sub>O<sub>4</sub> structure (Hanson & Jones, 1998; Milman-Barris *et al.*, 2008;
- 522 Jollands *et al.*, 2018), we explored whether the Cr exchange between olivine and spinel
- 523  $(K_{D_{Cr}})$  would improve the empirical thermometric equation:
  - *T* (K)

$$=\frac{10,000}{0.049(0.241)-0.657(0.040)\ln K_{\rm D_{Al}}-0.389(0.041)\ln K_{\rm D_{Cr}}+0.543(0.121)\rm Cr\#}$$

524

536

537

Cr-free experiments were excluded during regression (Fig. 5d and Table A2). The SEE and 525 526  $r^2$  values for the training dataset are  $\pm 31.7$  °C and 0.91, respectively, RMSE on the test dataset is 39.1 °C, and the p-value of the fit is 3.79e-47. The MCCV result shows that the 527 SEE<sub>MCCV</sub> and  $r^2_{MCCV}$  values for the training dataset are 35.8 °C and 0.94, respectively, and 528 the median RMSE<sub>MCCV</sub> value for the test dataset is 34.2 °C. We are aware that Eq. 8 includes 529 all the regression variables from Eq. 7. To determine if the improved performance of Eq. 8 530 is due to more regression terms or statistical significance, we performed a F-test (see 531 Appendix B for further details on the F-test model). The results of the F-test show an F 532 533 score of 88.4 and a *p*-value of 5.4e–15, which is statistically significant (p < 0.05) at 95% confidence. Therefore, the incorporation of Cr exchange as  $K_{D_{Cr}}$  into Eq. 8 produced a 534 statistically superior fit compared to Eq. 7. 535

#### Factors affecting the performance of the models

538 Dependence of the models on spinel composition and olivine–spinel Cr exchange
539 Previous studies have noted limitations of the original thermometric calibration due

- 540 to the restricted range of Cr# values used in the calibration (e.g., Heinonen *et al.*, 2015; Xu
- 541 & Liu, 2016; Trela *et al.*, 2017). In Fig. 6, we compare spinel Cr# with the  $\Delta T$  calculated
- 542 using our three new models and that of Coogan *et al.* (2014). There is no apparent

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543	correlation between $\Delta T$ and spinel Cr# calculated using Eqs. 5, 7, and 8, whereas the model
544	of Coogan <i>et al.</i> (2014) shows a moderate correlation ( $r^2 = 0.41$ ). Spinel Fe <sup>3+</sup> /Fe <sup>2+</sup> has a
545	moderate impact on $\Delta T$ using the model of Coogan <i>et al.</i> ( $r^2 = 0.47$ ), a limited impact using
546	Eq. 7 ( $r^2 = 0.24$ ), and negligible impacts on Eqs. 5 and 8 (Fig. S6). The effect of spinel TiO <sub>2</sub>
547	content was negligible for all models (Fig. S7).
548	Spinel Mg# and Fe <sup>3+</sup> /(Fe <sup>3+</sup> + Cr + Al) affect the performance of Eq. 7 ( $r^2 = 0.28$ and
549	0.66, respectively) and the model of Coogan <i>et al.</i> (2014; $r^2 = 0.13$ and 0.76, respectively),
550	but have less effect on Eqs. 5 and 8 (Figs. 7, 8). Similarly, $K_{DCr}$ has a moderate effect on Eq.
551	7 and the model of Coogan <i>et al.</i> (2014), but not on Eqs. 5 and 8, suggesting that $K_{D_{Cr}}$ is
552	indeed an important parameter for improving the accuracy of the empirical models (Fig. S8).
553	The thermodynamic model Eq. 5 including the spinel composition, may also mimic the
554	effect of $K_{D_{Cr}}$ . The correlation of $\Delta T$ with the spinel compositional parameters and $K_{D_{Cr}}$
555	indicates that using only spinel Cr# and $K_{DA1}$ is not sufficient to accurately predict
556	temperature in all circumstances, particularly when spinel has a low Mg# or high $Fe^{3+}/(Fe^{3+})$
557	+ Cr + Al) as is often observed in natural LIP spinels (Fig. 1). In such cases, the model of
558	Coogan et al. (2014) and Eq. 7 may have larger uncertainties and may strongly overestimate
559	temperature.
560	
561	Dependence of the thermometers on $fO_2$ , $H_2O$ , and pressure
562	$fO_2$ does not strongly affect the accuracy of the temperature calibrations of any of
563	the models, but a weak negative correlation was observed for Eq. 8 (Fig. S9). We further
564	tested the effect of melt water content and pressure with our reanalyzed experimental dataset

567 effect on the olivine liquidus (Médard & Grove, 2008), this appears to be of negligible

565

566

568 consequence on the accuracy of any of the thermometric models (Fig. S10). A potential

(three hydrous experiments with 3.7–7.4 wt.% H<sub>2</sub>O in the experimental glass; seven dry

high-pressure experiments at 500-1350 MPa). Although water is known to have a profound

569 explanation of this is that the effect of water is compensated by changes to the olivine and

- 572 this correlation is due to sample bias related to the sparse high pressure experiments, or if
- 573 pressure may affect on the molar volume of spinel and site occupancy (Hamecher *et al.*,
- 574 2013). Nonetheless, D'Souza et al. (2020) proposed that pressure has no effect on the model
- 575 of Coogan *et al.* (2014).

576

570

571

#### 577 DISCUSSION

578 In this section, we discuss the limits of the thermometric models and their sensitivities to 579 subsolidus re-equilibration. We then establish a protocol for selecting the appropriate model 580 for applications to natural samples. Finally, we apply our models to constrain mantle 581 melting conditions and lithological components in a range of geological settings and use 582 thermodynamic calculations to assess potential re-equilibration in natural olivine-spinel 583 pairs.

584

### 585 Concerns, limits, and sensitivities of the models

Our thermodynamic model (Eq. 5) is accurate but involves the exchange of Fe and Mg, 586 which diffuse rapidly in spinel (Vogt et al., 2015). This means that temperature estimates in 587 588 natural systems could be affected by low-temperature re-equilibration due to olivine-spinel 589 subsolidus Fe-Mg exchange during cooling. We note that the system itself doesn't have be 590 subsolidus, i.e., the olivine-spinel pair could crystallize at high temperature but experienced 591 A Fe-Mg reset after being carried along in an evolved melt, or set in a partially molten mush 592 before re-entrainment and eruption. The empirical model Eq. 8 does not involve Fe and Mg 593 but incorporates Cr exchange between olivine and spinel. Subsolidus Cr re-equilibration in 594 olivine, potentially caused by Cr reset between the carrier liquid, could also result in Eq. 8 595 potentially yielding anomalously low temperature estimates, we however note that the 596 diffusivity of Cr in olivine is poorly constrained and may vary with fO<sub>2</sub> and Cr 597 concentrations in olivine (Ito & Ganguly 2006; Jollands et al., 2018). In contrast, the

598 empirical models of Eq. 7 and Coogan et al. (2014) only involve slow diffusive elements in 599 spinel and olivine (Vogt et al., 2015; Spandler & O'Neill, 2010). Therefore they are not 600 impacted by low-temperature re-equilibration, but tend to overestimate temperature when spinel has low Mg# or high  $Fe^{3+}/(Fe^{3+} + Al^{3+} + Cr^{3+})$  (Figs. 7 and 8). 601 602 To quantitively investigate the sensitivities of Eq. 5 to Fe-Mg re-equilibration and 603 Eq. 8 to Cr re-equilibration in olivine, we performed two numerical simulations: 1) we 604 modeled Fe-Mg re-equilibration in a stoichiometric Cr-spinel with the formula (Mg,  $Fe^{2+}$ )(Al, Cr,  $Fe^{3+}$ )<sub>2</sub>O<sub>4</sub>, and 2) we progressively added  $Cr_2O_3$  to olivine to model Cr re-605 606 equilibration. In the Fe-Mg re-equilibration model, Al diffusivity in olivine is slow (Spandler & O'Neill, 2010), so we assumed that spinel exchanges only Fe<sup>2+</sup> and Mg with 607 olivine and, therefore, the spinel Cr# remains constant at a fixed  $fO_2$ , which is consistent 608 with natural observations (Guo et al., 2009; Hu et al., 2022). Spinel Mg# were set to values 609 610 vary from 0.3 to 1 with an increment of 0.01 at each spinel Cr#, which we varied from 0.4 to 611 0.9 with the same increment to cover the whole range observed in nature. We held spinel 612  $Fe^{3+}/Fe^{2+}$  constant at 0.5, corresponding to the median of natural spinel compositions, Olivine  $Al_2O_3$  content was set to 500 µg/g. Temperatures were then calculated using Eq. 5. 613 Because the mechanism of Cr re-equilibration in olivine has not been well determined 614 (Milman-Barris et al., 2008; Shea et al., 2019; Lang et al., 2022), we modeled two scenarios. 615 616 In the first scenario we assumed a decoupled substitution between Cr and Al, i.e., the Cr 617 content changes but the Al content remains constant, consistent with our experimental 618 observation. In the second scenario we assumed coupled substitution of Cr and Al with a 619 molar ratio of 1, similar to the observations of Milman-Barris et al. (2008) in a fast growth 620 regime. Both scenarios olivine  $Cr_2O_3$  start from 200 to 1500 µg/g. Olivine Al<sub>2</sub>O<sub>3</sub> content 621 was set to 500  $\mu$ g/g in the first scenario and range from 134 to 1006  $\mu$ g/g in the second 622 scenario, holding molar Cr/Al = 1. Temperatures were then calculated using Eq. 8. 623 Our model results are shown in Fig. 10. Decreasing spinel Mg# and decreasing 624 olivine Cr<sub>2</sub>O<sub>3</sub> content (both decoupled and coupled substitutions) return lower temperatures 625 when using Eqs. 5 and 8, respectively. For a given degree of Fe-Mg or Cr re-equilibration,

the decrease in apparent temperature after a given degree of re-equilibration is greater for

spinel with higher Cr#, whereas Eq. 5 tends to underestimate temperature for spinel with

628 lower Mg#.

629

#### 630 Selecting the appropriate model for application to natural samples

**631** The reliability of the empirical model in Eq. 7 is limited to a specific range of spinel

632 compositions and has larger errors on the temperature estimation compared to the

thermodynamic model Eq. 5 and the empirical model Eq. 8. In contrast, Eqs. 5 and 8 are

634 more precise but may be affected by chemical re-equilibration, as discussed above. Given

635 the greater uncertainties in the empirical expressions (Eqs. 7 and 8), for a fully equilibrated,

636 primitive olivine-spinel pair, we expect them to return temperatures scattering above and

637 below that returned by the thermodynamic expression (Eq. 5). In turn, we expect Eq. 5 to

return lower temperatures than the Eqs. 7 and 8 if diffusive re-equilibration occurs.

Therefore, we consider that Eqs. 7 or 8 should be used when Eq. 5 has a more than 50%

640 chance of underestimating the true temperature (i.e., when the mean temperatures estimated

641 using Eq. 5 and either Eq. 7 or 8 differ sufficiently that a *p*-value test on the distributions

having the same mean returns p < 0.5). Model selection is then determined by the likelihood that Eq. 5 yields an underestimation. If the temperature estimated using Eq. 5 is higher than that estimated using Eq. 7 or 8, the result from the thermodynamic expression should be

645 chosen. If not, a statistical Z-test is introduced to decide whether to choose the

646 thermodynamic expression or empirical expression (e.g., Snedecor & Cochran, 1989):

$$Z = \frac{\left|T_i - T_j\right|}{\sqrt{\sigma_{X_i}^2 + \sigma_{X_j}^2}} \tag{9}$$

647 where  $\sigma_{X_i}$  and  $\sigma_{X_j}$  are the RMSEs of the two models being compared (Eqs. 5 vs. Eq. 7 648 and/or Eq. 8) and  $T_i$  and  $T_j$  are the respective temperature estimates. If Z > 1.35, there is a 50% 649 chance that the two temperatures differ by a value larger than that attributed to their 650 combined uncertainties, and that Eqs. 7 or 8 return a higher temperature which is closer to

651	the true value. If both Eqs. 7 and 8 are compared with Eq. 5, the pair that yields with the
652	higher $Z$ score should be chosen to obtain a larger probability that the temperature estimates
653	are different from each other. The Z value of 1.35 represents a width of 1.35 standard
654	deviations, which is required for 50% of normal distributions with standard deviations given
655	by each model to lie between the mean $-\frac{1}{2} T_i - T_j $ and the mean $+\frac{1}{2} T_i - T_j $ ( <i>i</i> and <i>j</i>
656	represent each thermometric model); i.e., when $Z < 1.35$ , 50% of the estimated temperatures
657	should be more similar if they are measuring the same temperature. A python script and an
658	B Excel spreadsheet to perform all relevant calculations can be found at
659	https://github.com/eazzzon/olspthermo). Table 3 lists all scenarios for selecting the
660	appropriate model.
661	To illustrate the use of the protocol outlined above, we provide two geological
662	examples. In the first example, we take an ocean island basalt from Pico Island (Azores,
663	B Portugal) containing a $Fo_{83}$ olivine crystal, which has two spinel inclusions (van Gerve <i>et al.</i> ,
664	2021). X-ray chemical maps show no Al, Cr, or P zoning in the olivine and stable Al
665	concentration profiles were acquired from the olivine till approaching the spinel-olivine
666	boundaries (see in the Supplementary material, Section 5). The spinel inclusions have Mg#
667	$\approx 0.51$ and Fe <sup>3+</sup> /(Fe <sup>3+</sup> + Cr + Al) $\approx 0.17$ . The temperatures calculated using Eqs. 5, 7, and 8
668	are $1091 \pm 23$ °C, $1171 \pm 43$ °C, and $1184 \pm 34$ °C, respectively, with errors propagated
669	from both analytical and thermometric uncertainties. Because the thermodynamic expression
670	returns a lower temperature than the empirical expressions, we applied the Z-test model. The
671	results show $Z = 1.62$ between Eqs. 5, and 7 and $Z = 0.76$ between Eqs. 5 and 8. Because of
672	the larger Z-score, the result from Eq. 7 is statistically more robust, which is consistent with
673	the temperature estimated from a melt inclusion within the same crystal using liquid
674	thermometry (~1194 $\pm$ 46 °C; van Gerve <i>et al.</i> , 2021) and previous experimental studies
675	(e.g., Toplis & Carroll, 1995; Grove et al., 1992). According to these results, Eq. 5
676	significantly underestimates the temperature by $\sim 80$ °C.

- 677 For our second example, we take four published analyses of a MORB sample (A25-
- 678 D20-8) from Coogan *et al.* (2014), which contains an  $Fo_{91}$  olivine and a spinel with Mg# =
- 679 0.81. Eqs. 5, 7, and 8 return temperatures of  $1265 \pm 23$  °C,  $1235 \pm 43$  °C, and  $1198 \pm 34$  °C,
- 680 respectively. Here, the thermodynamic expression is chosen because it returns a higher
- temperature than the empirical expressions, hence the recommended temperature is 1265 ± 681
- 682 23 °C.
- 683

689

#### Application of the extended OSAT to natural basalts 684

- 685 Here, we apply our new thermometers to basalts from various settings: the Siqueiros MOR,
- Iceland and Skye in the North Atlantic igneous province (NAIP), and the Emeishan, 686
- Etendeka, and Caribbean LIPs. The results of different models are shown in Table 4 and Fig. 687
- 11. In each case, and for each analysis, we chose results according to the Z-test protocol 688
- described above. For comparison, all temperature estimates from each model are shown in
- Fig. S11. Given the similarity of spinel compositions from MORB, Iceland and Skye with 690
- 691 those from experiments used in the previous calibration of Coogan et al. (2014) and Wan et
- al. (2008), the new temperature estimates from our models are comparable or fall within 692
- mutual uncertainties. Intra-plate igneous provinces generally have higher  $Fe^{3+}/(Fe^{3+} + Cr + Cr)$ 693
- Al) and lower Mg# (Fig. 1), therefore the model of Coogan et al. (2014) significantly 694
- 695 overestimate the temperature by 30–100 °C (median) and 60–120 °C at the maximum.
- 696
- 697 MORB (Siqueiros)

698 MORB lavas erupted at Siqueiros are generated by near-adiabatic upwelling of the ambient 699 upper mantle (e.g., Gregg *et al.*, 2009). Siqueiros spinel has Cr = 0.24-0.44, and potentially 700 experienced less Fe-Mg subsolidius re-equilibration, with generally high Mg# values (0.70-0.81) and Fe<sup>3+</sup>/(Fe<sup>3+</sup> + Cr + Al) values (0.03–0.07). Olivine contains 335–827  $\mu$ g/g Al<sub>2</sub>O<sub>3</sub> 701 702 and 508–2,750 µg/g Cr<sub>2</sub>O<sub>3</sub>. Coogan et al. (2014) and Matthews et al. (2021) calculated the co-crystallization temperature of Siqueiros olivine and spinel to be 1230<sup>+59</sup><sub>-91</sub> °C (expressed 703

as  $median^{+[95^{th} percentile - median]}_{-[median - 5^{th} percentile]}$ , being consistent with the description of distribution 704 used in Matthews et al. [2021]). Our new estimated temperatures are 1258<sup>+37</sup><sub>-76</sub> °C. These 705 706 new estimates generally agree with the previous studies, which we attribute to the spinel 707 composition in Sigueiros lavas being comparable to those used in the calibration 708 experiments of Coogan et al. (2014) and Wan et al. (2008). 709 710 NAIP (Iceland and Skye) Iceland and Skye are within the NAIP, which is accepted as having a mantle-plume origin. 711 Spinel in Iceland has Cr# 0.24–0.60, Mg# 0.54–0.79 and  $\text{Fe}^{3+}/(\text{Fe}^{3+} + \text{Cr} + \text{Al})$  0–0.13. 712 Olivine ins Iceland contain Al<sub>2</sub>O<sub>3</sub> 396–1,020  $\mu$ g/g and Cr<sub>2</sub>O<sub>3</sub> 345–2,686  $\mu$ g/g. Spinel in 713 Skye is slightly more primitive, containing Cr# 0.41–0.56, Mg# 0.70–0.77 and  $Fe^{3+}/(Fe^{3+} +$ 714 Cr + Al) 0.03–0.06. Olivine in Skye has more  $Al_2O_3$  (794–1,519 µg/g) and less  $Cr_2O_3$ 715 (1,030–1,890 µg/g). Previous crystallization temperature estimates (Spice *et al.*, 2016; 716 717 Matthews et al., 2016) using the model of Coogan et al. (2014) gave crystallization temperatures of 1283<sup>+84</sup><sub>-66</sub>  $^{\circ}$ C for Iceland and 1409<sup>+60</sup><sub>-65</sub>  $^{\circ}$ C for Skye. Our new results are 718  $1275_{-100}^{+52}$  °C for Iceland and  $1388_{-53}^{+41}$  °C for Skye, which are similar on the median values 719 720 but ~ 40 °C lower for the high temperature populations (on the  $95^{\text{th}}$ ) 721 722 Emeishan large igneous province The ~260 Ma Emeishan LIP in southwest China is considered to be of mantle-plume origin 723 (Chung & Jahn, 1995; Xu et al., 2001; Xiao et al., 2004). Emeishan spinel compositions are 724 725 much more varied compared to MORB, with Cr = 0.42-0.72, Mg = 0.48-0.72, and  $Fe^{3+}/(Fe^{3+} + Cr + Al) = 0.04-0.15$ . Olivine contains 340–990 µg/g Al<sub>2</sub>O<sub>3</sub> and 165–1,420 726 727 µg/g Cr<sub>2</sub>O<sub>3</sub>. Previous temperature estimates give a crystallization temperature of

- 728  $1271_{-55}^{+108}$  °C (Xu & Liu, 2016). Our new estimate is  $1224_{-54}^{+109}$  °C, with both the median
- and 95<sup>th</sup> value being ~ 47  $^{\circ}$ C cooler. Here, the compositional deviation of the natural olivine

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and spinel from those used in the previous calibration results in the overestimated

### 731 crystallization temperatures.

- 732
- 733 Etendeka large igneous province
- 734 The Etendeka LIP is the southern part of Paraná-Etendeka LIP, mainly outcropping in
- 735 Namibia and southern Angola (Thompson *et al.*, 2001; Gibson, 2002). The high
- 736 crystallization temperatures and high  ${}^{3}\text{He}/{}^{4}\text{He}$  values (>26  $R_{A}$ , reported relative to that of the
- 737 present-day atmosphere) of the lavas suggest a mantle plume origin (Stroncik *et al.*, 2017).
- **738** Etendeka spinel compositions are quite variable, with Cr = 0.40-0.70, Mg = 0.24-0.76,
- 739 and  $\text{Fe}^{3+}/(\text{Fe}^{3+} + \text{Al}^{3+} + \text{Cr}^{3+}) = 0.04-0.31$ . Olivine contains 300-1,200 µg/g Al<sub>2</sub>O<sub>3</sub> and 200-
- 740 2,200  $\mu$ g/g Cr<sub>2</sub>O<sub>3</sub>. Jennings *et al.* (2019) calculated the olivine crystallization temperature
- vising the model of Coogan *et al.* (2014) to be  $1323_{-101}^{+151}$  °C. Our new results are
- 742  $1307_{-88}^{+110}$  °C, which are similar at the median value but 57 °C lower for the high
- temperature populations (on the 95<sup>th</sup>).
- 744

#### 745 Caribbean large igneous province (Tortugal suite)

- 746 The Tortugal suite in Costa Rica is considered to be the product of the initial melting of the 747 Galapagos plume in the Caribbean large igneous province (Alvarado et al., 1997; Trela et al., 748 2017). Particularly, the Tortugal suite hosts high-Fo olivine (up to  $Fo_{94}$ ) with compositions 749 overlapping those of olivine in Archaean komatiites and containing 321-1,114 µg/g Al<sub>2</sub>O<sub>3</sub> 750 and  $807-2,195 \ \mu g/g \ Cr_2O_3$ . Tortugal suite spinel span a wide range of compositions, with Cr# = 0.66-0.85, Mg# = 0.33-0.72, and Fe<sup>3+</sup>/(Fe<sup>3+</sup> + Cr + Al) = 0.0-0.14. Trela *et al.* (2017) 751 752 used the model of Coogan et al. (2014) to estimate the crystallization temperature to be 753  $1492_{-173}^{+89}$  °C. Our new estimates, however, are much lower, at  $1425_{-149}^{+77}$  °C, which are 754 67 °C lower on the median and 79 °C lower for the high temperature populations (on the 95<sup>th</sup>). 755
- 756

#### 757 Implications for mantle melting, mantle lithologies, and subsolidus re-equilibration 758 Significant overestimation of temperature is often observed when the model of Coogan et al. 759 (2014) is applied to specific geological settings (e.g., intra-plate igneous provinces). This 760 model was however abundantly used to assess mantle melting conditions and lithologies in 761 those cases (e.g., Matthews et al., 2016; 2021) which therefore necessitate a revision. 762 Olivine crystallization temperatures can be converted to mantle $T_p$ when a correction for the 763 latent heat of melting is considered (Putirka et al., 2007), which is directly related to the 764 total melt fraction. The total melt fraction can be constrained from melting of a 765 homogeneous (e.g., Putirka et al., 2007) or a heterogenous mantle source (Matthews et al., 766 2016; 2021), and through geophysical observations of magmatic productivity, i.e., crustal thickness at spreading centers or magma flux at ocean islands (McKenzie & Bickle, 1988; 767 Shorttle et al., 2014; Matthews et al., 2016; 2021). 768 769 To be self-consistent with previous $T_p$ and lithology estimates, we applied the protocol of Matthews et al. (2016; 2021) to estimate the mantle melting conditions and mantle 770 771 lithologies for the localities investigated in this study. In this melting model, crustal 772 thickness at oceanic spreading centers and magma flux at ocean islands are used as 773 observable proxies for magma productivity. A multi-lithology mantle melting model during 774 adiabatic decompression following Phipps Morgan (2001) and Shorttle et al. (2014) is used 775 to calculate melting behavior. The source lithologies are assumed to be in thermal 776 equilibrium. A full description of the mathematical and computational formulation is 777 available in Phipps Morgan (2001) and Shorttle et al. (2014). An extra constraint was added 778 to prevent negatively buoyant solution in intra-plate magmatism, i.e., the multi-lithology 779 mantle should be buoyant with respect to the ambient mantle during the plume-driven 780 upwelling. We used the python interface forward model pyMelt (version 1.96; Matthews et 781 al., 2022) to calculate the melting behavior of mantle compositions comprising multiple 782 lithologies. The model considers three lithologies: lherzolite (matthews.klb1 lithology class 783 in pyMelt) and silica-deficient pyroxenite (matthews.kg1 lithology class) from Matthews et

*al.* (2021), and non-melting harzburgite (*shorttle.harzburgite* lithology class) from Shorttle

<b>785</b> <i>et al.</i> (2014)	).
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786 To convert an observation of primary crystallization temperatures  $(T_{crv})$  to the value 787 of  $T_p$ , the fractions of mantle lithologies ( $\phi_{Hz}$ , the harzburgite fraction;  $\phi_{Px}$ , the pyroxenite 788 fraction; and  $\phi_{Lz}$ , the peridotite/lherzolite fraction), we used an inversion model of 789 Matthews et al. (2021), a Bayesian Monte Carlo inference method (Feroz & Hobson, 2008 790 Feroz et al., 2009, 2013; Buchner et al., 2014) to find the set of solutions which can reproduce  $T_{cry}$  and mantle lithologies with applicable constraints (e.g., crustal thickness, 791 792 magma flux). The inversion model parameters (lithospheric thickness, crustal thickness, the 793 fraction of pyroxenite derived melt, magma flux) were kept identical to Table 1 in Matthews et al. (2021), also given in supplementary dataset (Table S4), except for the  $T_{cry}$ , which we 794 795 adapted to our new estimates above. We note that these parameters may not be the most 796 appropriate for all our tested localities, but the focus here is on the influence of changing  $T_{\rm crv}$ on mantle  $T_p$  and lithology estimates. The inversion model requires the selection of an 797 798 appropriate olivine composition and crystallization temperature as the starting point of the 799 calculation. Matthews et al. (2021) assumed Fo<sub>91</sub> olivine as the composition equilibrated with primitive mantle-derived melts based on the most primitive olivine observed in nature. 800 801 However, both experimental and natural observations (e.g., the Tortugal suite) have shown 802 olivine of higher Fo (>Fo<sub>94</sub>), indicating that the source diversity controls primary olivine 803 compositions atdifferent localities (see further discussion below). Here, we chose the 804 maximum of our calculated  $T_{\rm crv}$  estimates (the corresponding olivine Fo are close to or 805 higher than 91) at each locality as the primary olivine temperature to simplify the constraints 806 on the equilibrated olivine composition.

807 Our inversion results for  $T_p$  and mantle lithological fractions are summarized in 808 Table 5 and shown in Fig. 12, details are given in the supplementary dataset. Because 809 different model parameters are used to constrain mantle  $T_p$  in different methods (e.g., 810 Putirka, 2005; 2016), we only compare our new  $T_p$  values with those of Matthews *et al.* 811 (2021) in Fig. 13 to be self-consistent.

812	Our new estimated $T_p$ and lithology fractions ( $\phi_{Hz}$ and $\phi_{Px}$ ) for MORB (1355 <sup>+24</sup> °C, -19)
813	$0.34_{-0.30}^{+0.25}$ , and $0.02_{-0.02}^{+0.02}$ , respectively), Iceland ( $1518_{-17}^{+21}$ °C, $0.23_{-0.17}^{+0.18}$ , $0.08_{-0.04}^{+0.04}$ ), Skye
814	$(1550^{+80}_{-82} {}^{\circ}\text{C}, 0.51^{+0.32}_{-0.39}, 0.12^{+0.16}_{-0.10})$ , and Etendeka $(1577^{+106}_{-98} {}^{\circ}\text{C}, 0.53^{+0.31}_{-0.38}, 0.11^{+0.19}_{-0.10})$ are
815	comparable to those of Matthews et al. (2021). However, due to previous overestimation of
816	crystallization temperatures for intra-plate LIPs, our newly estimated $T_p$ and lithological
817	fractions differ significantly from Matthews <i>et al.</i> (2021). Our respective results for
818	Emeishan $(1481_{-78}^{+87} \circ C, 0.53_{-0.36}^{+0.29}, 0.10_{-0.09}^{+0.14})$ return a median $T_p$ that is 74 °C cooler than
819	previous inversions (Table 4), and our results for the Tortugal suite $(1648^{+193}_{-91} \circ C, 0.50^{+0.31}_{-0.38}, 0.50^{+0.31}_{-0.38})$
820	$0.15^{+0.29}_{-0.13}$ ) return a median $T_p$ that is 165 °C cooler, which leads to a remarkable difference
821	in the estimated lithological fractions (Table 4).
822	To verify the reliability of the mantle melting and lithological estimates, we
823	performed forward thermodynamic modeling based on the estimated mantle $T_p$ and
824	lithological fractions to calculate the equilibrated olivine composition in composition-
825	temperature space. We used the Matlab/Julia Mineral Assemblages Gibbs Energy
826	Minimization package (MAGEMin v1.3.0; Riel et al., 2022) to calculate the equilibrium
827	melt composition of a given mantle lithology under mantle melting conditions calculated
828	from inversion. For the modeling, we use the thermodynamic database of Holland et al.
829	(2018). The KLB-1 peridotite (Wasylenki et al., 2003) and MIX1G pyroxenite (Hirschmann
830	et al., 2003) were mixed with the median ratio of the estimated lithological fractions to
831	generate an equilibrated melt with the mantle source for each locality. The equilibrated melt
832	then crystallizes mineral phases within a given temperature interval at the base of the crust
833	(same as the inversion model, see Table S3 in the supplementary dataset and Table 1 in
834	Matthews et al. [2021]) at each locality. The model results are compared to our estimated
835	crystallization temperatures at each locality in Fig. 11. The equilibrated olivine usually
836	includes sub-populations containing higher-Fo (>Fo <sub>91</sub> ) than those in the erupted lavas, but

are generally consistent with the crystallization temperatures calculated using our extended

838 OSAT, whereas the results of Coogan *et al.* (2014) mostly overestimate the temperature.

839

#### 840 CONCLUSIONS

- 841 We revised the formalism of the olivine-spinel aluminum exchange thermometer by
- 842 performing new experiments and reanalyzing published experiments to extend the
- calibration P, T,  $fO_2$ , and  $H_2O$  conditions. Three models were regressed on the extended
- dataset: a thermodynamic-based model (Eq. 5) and two empirical models (Eqs. 7 and 8).
- 845 The thermodynamic model of Eq. 5 shows the best performance with the lowest
- 846 uncertainties. The exchange of Al and Cr between olivine and spinel significantly effects the
- results when using the empirical models, whereas  $fO_2$ , water, and pressure have little to
- 848 moderate effects on all models. The empirical models may only be accurate within a
- 849 restricted spinel composition, leading to temperature overestimations by ~30–100 °C in
- some intra-plate LIPs, but only minor differences in Iceland and MOR. These improved  $T_{\rm cry}$
- estimates translate to a significant difference in the mantle  $T_{p}$  and lithological fractions
- 852 calculated for the investigated intra-plate LIPs.
- 853

#### 854 APPENDIX A. THERMODYNAMIC MODEL OF OLIVINE-SPINEL AL

855 PARTITIONING

Here, we develop our thermodynamic formalism to understand the Al exchange between
olivine and spinel. The most plausible solution mechanism for the incorporation of Al into
olivine and Cr-spinel is (Coogan *et al.*, 2014):

$$(MgAl^{TET})Al^{OCT}O_4^{Spl} = (MgAl^{TET})Al^{OCT}O_4^{Ol}$$
(A.1)

#### 859 Thus, the equilibrium of the reaction is written as:

$$\mu_{\text{MgAl}_2\text{O}_4^{\text{Ol}}} = \mu_{\text{MgAl}_2\text{O}_4^{\text{Spl}}} \tag{A.2}$$

860 where  $\mu_{ab}$  is the chemical potential of component *a* in phase *b*. For a given phase:

$$\mu_a = \mu_a^0 + RT \ln \alpha^a \tag{A.3}$$

861 where  $\mu_a^0$  is the chemical potential of component *a* at standard state,  $\alpha^a$  is the activity of

862 component *a*, *R* is the universal gas constant [8.314 J mol<sup>-1</sup> K<sup>-1</sup>], and *T* is the absolute

- temperature in Kelvin.
- For olivine, we can write:

$$\alpha_{\mathrm{MgAl_2O_4^{Ol}}} = \gamma \times X_{\mathrm{MgAl_2O_4^{O}}}$$

$$X_{\text{MgAl}_{2}\text{O}_{4}^{\text{Ol}}} = (\text{Al}_{2}\text{O}_{3}^{\text{Ol}} / M_{\text{Al}_{2}\text{O}_{3}}) \times M_{\text{MgAl}_{2}\text{O}_{4}}$$
(A.5)

865 where  $\gamma$  is the activity coefficient and *M* the molar weight of the component, and *X* is the

866 mole fraction. Combining Eqs. A.3–A.5, we obtain:

$$\mu_{\text{MgAl}_2\text{O}_4} = \mu_{\text{MgAl}_2\text{O}_4}^0 + RT \ln \left[ \gamma \times \left( A l_2 O_3^{Ol}_{wt.\%} / M_{A l_2 O_3} \right) \times M_{MgA l_2 O_4} \right]$$
(A.6)

We apply a thermodynamic model involving cation site ordering (Sack, 1982; Sack & Ghiorso, 1991a, 1991b) to describe the thermodynamic properties of the components in Crspinel. In this model, spinel is assumed to be stoichiometrically perfect ( $R_3O_4$ ), with *Fd3m* space group symmetry. Five 'fictive' independent compositional variables ( $X_i$ ) and ordering parameters ( $s_i$ ) are needed to describe the compositional change in Cr-spinel, which can be calculated from mole fractions of cations in the spinel.

873 The independent compositional and ordering variables as well as the definitions of 874 site mole fractions are from Sack & Ghiorso (1991a, 1991b) and reported in Table. A1. The 875 molar Gibbs energy  $\overline{G}$  is calculated from the vibrational Gibbs energy  $\overline{G}^*$  and the ideal 876 molar configurational entropy  $\overline{S}^{IC}$  as:

$$\overline{G} = \overline{G}^* - T\overline{S}^{\rm IC} \tag{A.7}$$

877 A second-degree Taylor expansion of the compositional and ordering variables is used to878 describe the molar vibrational Gibbs energy as (Thompson Jr, 1969):

$$\overline{G}^{*} = g_{0} + \sum_{i} \left( g_{i}X_{i} + g_{ii}X_{i}^{2} + \sum_{j < i} g_{ij}X_{i}X_{j} \right)$$

$$+ \sum_{i} \sum_{k} \left( g_{ik}X_{i}s_{k} \right)$$

$$+ \sum_{k} \left( g_{k}s_{k} + g_{kk}s_{k}^{2} + \sum_{l < k} g_{kl}s_{k}s_{l} \right)$$
(A.8)

where g terms are Taylor coefficients from Table 3 in Sack and Ghiorso (1991b). The

881 configurational entropy is calculated from compositional variables and site fractions as:

$$\overline{S}^{\rm IC} = -R \sum_{r} \sum_{c} \overline{r} X_{c,r} \ln X_{c,r} \qquad (A.9)$$

- 882 where  $X_{c,r}$  is the fraction of cation *c* in site *r* in terms of  $X_i$  and  $s_i$ , and  $\bar{r}$  is the number of *r*
- sites per formula unit. A Darken equation is then used to manipulate and express the
- chemical potential of the spinel endmembers (Darken & Gurry, 1953; Sack, 1982; Ghiorso,
- 885 1990; Sack & Ghiorso, 1991a):

$$\mu_{j} = \overline{G} + \sum_{i} n_{i,j} (1 - X_{i}) \left( \frac{\partial \overline{G}}{\partial X_{i}} \right)_{\frac{X_{k}}{X_{l}}, s_{m}} + \sum_{i} \left( q_{ij} - s_{i} \right) \left( \frac{\partial \overline{G}}{\partial s_{i}} \right)_{X_{k}, s_{l}}$$
(A.10)

- where  $n_{i,j}$  and  $q_{i,j}$  represent coefficients of  $X_i$  and  $s_i$ , respectively, in 1 mol of spinel
- component *j*. The chemical potential of component  $MgAl_2O_4$  in a spinel solution can be
- 888 written as:

1G

# $\mu_{MgAl_{2}O_{4}}^{Spl} = \overline{G}_{2}^{*} + RT \ln[(X_{2})/(1+X_{4})(1-X_{3}-X_{4}-X_{5})^{2}]$ $+ W_{Fe-Mg}^{TET}(1-X_{2})(1+X_{4}-X_{2})$ $+ \Delta \overline{G}_{23}^{0}(1-X_{2})(X_{3}) + \Delta \overline{G}_{24}^{0}(1-X_{2})(X_{4}) + \Delta \overline{G}_{25}^{0}(1-X_{2})(X_{5})$ $+ W_{1'3'}X_{3}(X_{3}+X_{4}+X_{5})$ $+ W_{1'4'}X_{4}(X_{3}+X_{4}+X_{5})$ $+ W_{1'5'}X_{5}(X_{3}+X_{4}+X_{5})$ $- W_{3'4'}(X_{3})(X_{4}) - W_{3'5'}(X_{3})(X_{5})$ $- W_{4'5'}(X_{4})(X_{5})$

889 where  $\overline{G}_{2}^{*}$  is the vibrational Gibbs energy of the MgAl<sub>2</sub>O<sub>4</sub> endmember. The Gibbs energy 890 along joins between vertices differing in composition (i.e.,  $\Delta \overline{G}_{23}^{0}, \Delta \overline{G}_{24}^{0}$ , and  $\Delta \overline{G}_{25}^{0}$ ) are 891 standard state Gibbs free energies of the Mg-Fe exchange reactions between aluminate (*G*<sub>2</sub>), 892 chromate (*G*<sub>3</sub>), titanate (*G*<sub>4</sub>), and ferrite spinel (*G*<sub>5</sub>).  $W_{\text{Fe}-Mg}^{\text{TET}}$ ,  $W_{ij}$ , and  $W_{ij}$ <sup>\*</sup> are symmetric 893 regular solution parameters describing deviations of the Gibbs energy from ideal mixing. To 894 focus on Al partitioning between olivine and spinel, we combine Eqs. A.2, A.6, and A.10 895 with independent parameters  $X_i$  (Table A1) as:



897 On the left side of Eq. A.13,  $(X_2)/(1+X_4)(1-X_3-X_4-X_5)^2$  is equal to  $X_{Mg^{2+}}^{TET}X_{Al^{3+}}^{OCT^2}$ ; on the right 898 side of Eq. A.13, all parameters are constants or barely change with temperature (see Sack & 899 Ghiorso, 1991a), except  $\mu_{MgAl_2O_4^{Ol}}^0$ ,  $\overline{G}_2^*$  and the independent composition  $X_i$ .  $\overline{G}_2^*$  can be

900 rewritten as 
$$\left(c_1X_2+c_2X_2^2+c_3X_2^{\frac{1}{2}}\right)$$
 following Eq. A.8, where  $c_1$ ,  $c_2$  and  $c_3$  are constants

- 901 derived from the Taylor expansion coefficients. However, given that the standard state
- 902 MgAl<sub>2</sub>O<sub>4</sub> (i.e.,  $\mu_{MgAl_2O_4^{Ol}}^0$ ) in olivine is not well constrained,  $\mu_{MgAl_2O_4^{Ol}}^0$ , by definition, is
- relevant to temperature at any given pressure and composition. We thus simplified
- 904  $\mu_{MgAl_2O_4^{Ol}}^0$  to be a function of temperature and independent of olivine composition given the 905 trace concentration of Al in olivine (Ganguly, 2008), and  $\gamma$  for Al<sub>2</sub>O<sub>3</sub> in olivine can be 906 assumed to be 1. We thus encapsulate all parameters related to spinel as  $\varphi^{Spl}$ , and Eq. A.13 907 can be fitted with coefficients ( $c_i$ ) replacing the regular solution parameters, and related to
- 908 the spinel composition as:

$$\ln\left(\frac{X_{\text{Al}_2\text{O}_3^{\text{OI}}}}{X_{\text{Mg}^{2+}}^{\text{TET}}X_{\text{Al}^{3+}}^{\text{OCT}}}\right) + b = \frac{\varphi^{\text{Spl}}}{c_0 \times T}$$
(A.14)

909 where

RIGHT

$$\varphi^{\text{Spl}} = c_1 X_2 + c_2 X_2^2 + c_3 X_2^{\frac{1}{2}} + c_4 (1 - X_2)(1 + X_4 - X_2) + c_5 (1 - X_2)(X_3) + c_6 (1 - X_2)(X_4) + c_7 (1 - X_2)(X_5) + c_8 X_3 (X_3 + X_4 + X_5) + c_9 X_4 (X_3 + X_4 + X_5) + c_1 0 X_5 (X_3 + X_4 + X_5)$$

 $-c_{11}(X_3)(X_4) - c_{12}(X_3)(X_5) - c_{13}(X_4)(X_5) + c_{14}$ 

910 The coefficients b, and  $c_0-c_{12}$  can be then obtained by fitting the compositions of

911 experimental olivine and spinel pairs.

912

#### 913 APPENDIX B. F-TEST MODEL

914 An *F*-test is a statistical test with the null hypothesis that a distribution follows an *F*-

915 distribution; F-tests have been used in solving geological problems involving regression

916 (e.g., Abouchami et al., 2005). An F-test is commonly used to compare two models that

give similar or identical results but use different numbers of parameters, with the aim of

918 determining which model is statistically better or whether the model with fewer parameters

919 is nested within the model with more parameters. In this case, the *F*-value can be calculated

920 as:

$$F = \frac{(RSS_1 - RSS_2)/(p_1 - p_2)}{RSS_2/(n - p_2)}$$
(B.1)

where  $RSS_i$  represents the residual sum of squares of model *i*;  $p_1$  and  $p_2$  represent the number 921 922 of parameters in models 1 ( $m_1$ ) and 2 ( $m_2$ ), and  $p_1 < p_2$ ; and n is the amount of data points 923 used in the regressions. While comparing the two models, the null hypothesis is that  $m_2$  is 924 not better than  $m_1$  (i.e., that  $m_2$  is an overfitting of  $m_1$ ). When comparing the *F*-value with 925 the critical *F*-value, which can be calculated from the *F*-distribution with  $(p_2 - p_1, n - p_2)$ degrees of freedom, if F exceeds the critical value, the associated p-value is small ( $<1 - \alpha$ ; 926 e.g.,  $\alpha = 0.05$  for 95% confidence) and then null hypothesis is rejected, meaning that  $m_2$  is 927 928 indeed statistically better. If F is below the critical value, the associated p-value is larger than  $\alpha$  and  $m_1$  is statistically better than  $m_2$ . 929

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#### **1309** Figure captions

1310 1311 Fig. 1. Spinel compositions in calibration experiments and natural rocks. (a) Ternary projection of spinel prism, Al-Cr-Fe<sup>3+</sup>. (b) Spinel Cr# as a function spinel Mg# (c) Spinel 1312 TiO<sub>2</sub> (wt. %) as a function of spinel  $Fe^{3+}/(Fe^{3+} + Cr + Al)$ . (d) Spinel  $Fe^{3+}/(Fe^{3+} + Cr + Al)$ 1313 as a function of spinel  $Fe^{2+}/(Fe^{2+}+Mg)$ . Data from Barr et al. (2009), Mitchell and Grove 1314 1315 (2015), Charlier et al. (2018), Parman and Grove (2004), and Thy (1995) are remeasured 1316 published experiments. Data from Matzen et al. (2011) and Hanson & Jones (1998) are used 1317 as part of the test dataset. See text for details on the calibration and test dataset. Natural 1318 samples are compiled from published studies: MORB (Coogan et al., 2014; Matthews et al., 1319 2021;), Iceland (Spice et al., 2016; Matthews et al., 2016), Skye (Spice et al., 2016), 1320 Emeishan (Xu & Liu, 2016; Zhang et al., 2021; Li et al., 2021; Wu et al., 2022), Etendeka 1321 (Jennings et al., 2019), Tortugal (Trela et al., 2017). † represents experiments used in the 1322 training dataset, \* represents experiments used in the test dataset. Error bar represents  $l\sigma$ 1323 standard deviation and is not observable if it is smaller than the symbol. 1324 Fig. 2. (a)  $Al_2O_3$  and (b)  $Cr_2O_3$  concentrations in olivine ( $\mu g/g$ ) as a function of Fo content (mol.%). Data sources are as in Fig. 1. Error bar represents 10 standard deviation and is not 1325 1326 observable if it is smaller than the symbol. 1327 1328 Fig. 3. Representative backscattered electron images of experimental products. Quench texture is observed in high temperature experiments (>1500 °C), euhedral olivine and spinel 1329 1330 indicate the attainment of equilibrium in the experiments. Abbreviations: Gl, glass; Cr-Spl, 1331 Cr-spinel; Ol, olivine. 1332 1333 Fig. 4. Melt compositional variations as functions of melt MgO content (wt.%). 1334 Experimental data sources are as in Fig. 1. Error bar represents  $1\sigma$  standard deviation and is 1335 not observable if it is smaller than the symbol. 1336 1337 1338 Fig. 5. Comparison of crystallization temperatures calculated using (a) the model of Coogan et al. (2014) and (b-d) Eqs. 5, 7, and 8, respectively. The green line represents the results of 1339 1340 a Monte Carlo simulation showing the robustness of the fit when  $1\sigma$  standard errors on 1341 spinel and olivine compositions are considered. The  $r^2$  values represent the robustness of the 1342 linear regressions between the measured temperatures and predicted temperatures. The 1343 brown and purple fields represent temperatures within  $\pm 50$  °C and  $\pm 100$  °C. Data sources as 1344 in Fig. 1. 1345 Fig. 6. Effects of spinel Cr# on (a) the OSAT of Coogan et al. (2014) and (b-d) the 1346 1347 extended ØSAT using Eqs. 5, 7, and 8, respectively. Data sources as in Fig. 1. 1348 1349 Fig. 7. Effects of spinel Mg# on (a) the OSAT of Coogan *et al.* (2014) and (b–d) the 1350 extended OSAT using Eqs. 5, 7, and 8, respectively. Data sources as in Fig. 1. 1351 Fig. 8. Effects of spinel  $Fe^{3+}/(Fe^{3+} + Cr + Al)$  on (a) the OSAT of Coogan *et al.* (2014) and 1352 1353 (b-d) the extended OSAT using Eqs. 5, 7, and 8, respectively. Data sources as in Fig. 1. 1354 1355 Fig. 9. Effects of pressure (MPa) on (a) the OSAT of Coogan et al. (2014) and (b-d) the 1356 extended OSAT using Eqs. 5, 7, and 8, respectively. Data sources as in Fig. 1. 1357 1358 Fig. 10. Modeled effects of (a) Fe-Mg re-equilibration in spinel on temperatures estimated 1359 using Eq. 5 and (b) Cr re-equilibration (decoupled substitution) in olivine on temperatures 1360 estimated using Eq 8. (c) Cr re-equilibration (coupled substitution with Al) in olivine on 1361 temperatures estimated using Eq 8. 1362

Fig. 11. Estimated crystallization temperatures for natural samples. Grav symbols are results using the model of Coogan et al. (2014). Blue symbols are results using the extended OSAT of this study that passed the Z-test protocol described in the text. Red lines are the calculated equilibrated olivine compositions and temperatures from our forward thermodynamic modeling using MAGEMin. See text for more details. The histograms along the left axis show compare the distribution of results using our model to those using the model of Coogan et al. (2014). The histogram along the bottom axis shows the distribution of olivine Fo contents, which is the same in both datasets. Fig. 12. Mantle lithologies and melting temperatures determined by our inversion analysis of natural samples using our extended OSAT, the contour lines represent density plot of the results. Fig. 13. Comparison of estimated mantle potential temperatures in this study (colored) with those of Matthews et al. (2021; gray). The temperature excess on the right axis is calculated here relative to the median  $T_p$  of MORB (Siqueiros). The heights of the boxes show the interquartile range (25<sup>th</sup>-75<sup>th</sup>), and the upper and lower whisker bars show maximum and minimum values beyond which, data points represent outliers. 





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1501	Table.	1.	Starting	compo	sitions	used	in	this	stud	y
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	10010. 1. Duu	ung oc	Jub 600	nuons	ubeu II	i tino i	study.								-
	Sample	Ind	Si	Ti	$Al_2$	Fe	Mn	Mg	Ca	Na <sub>2</sub>	$K_2$	$P_2$	Cr <sub>2</sub>	<b>C</b>	-
	Name	ex <sup>a</sup>	$O_2$	$O_2$	$O_3$	$O_t$	Ο	Ō	Ο	Ο	0	$O_5$	$O_3$	Sum	
	natural				<u> </u>	ť						U			-
	rocks														
	reens		40	13	69	12	0.1	28	81	1.0	0	04	03	100	
	Ny17-135	Α	+0. 52	2	1	12.	0.1	20.	0.1	2	70	0. <del>4</del>	5	100.	
	NI17		22	) 10		21 11	0	14	0	5	/0	4	5	100	
	Ny17-	В	37.	1.2	6.4	11.	0.1	<u>31</u> .	/.6	0.9	0.	0.4	1.9	100.	S N
	135_Cr2*		95	4	/	43	/	08	6	6	66	1	/	00	) /
	Ny17-	С	39.	3.1	10.	11.	0.2	10.	16.	3.0	2.	1.4	0.0	100.	
	161_Cr	U	67	9	61	50	2	99	30	1	97	8	6	00	
	Ny17-	р	39.	3.1	10.	11.	0.2	10.	16.	2.9	2.	1.4	1.4	100.	
	161_Cr2	D	11	4	46	34	1	84	07	7	93	6	(8)	00	
	NYAM_Cr	Г	41.	1.1	4.3	12.	0.0	31.	6.2	0.5	0.	0.1	1.9⁄	100.	
	2*	E	88	2	4	03	7	08	3	0	65	< 4	8	00	
		_	50.	0.7	16.	9.3	0.1	10.	10.	2.1	Ø.	0.0	0.0	100.	
	<b>O</b> s82	F	21	0	33	9	5	23	29	8	33	9	9	00	
			<u>4</u> 9	06	16	92	01	10	10	21	-0	0.0	14	100	
	Os82_Cr	G	ч <i>у</i> . 52	0.0	11	6	5	00	15	<u>2</u> .1	23	0.0	6	100.	
	DI		20	00	11	72	0.1	22	12		<i>y</i> 55	2	1.0	100	
	PI- 052 C-2*	Η	39. 42	0.8	4.5	/.5	0.1	52. 02	12.	0.1	0.	0.0	1.9	100.	
	052_Cr2*		42	0	5	9	2	93	68		11	0	8	100	
	TO1	Ι	44.	0.7	3.9	11.	0.1	32.	4.8	0.1	0.	0.1	0.4	100.	
	_		11	9	4	88	1	86	12	2	05	9	2	00	
	TO1 Cr	T	44.	0.7	3.9	11.	0.1	32.	4.8	0.1	0.	0.1	0.7	100.	
	101_01	5	62	9	3	84	7	74	0	2	05	9	8	00	
	TO1 C-2	V	44.	0.7	3.8	11	0.1	32.	4.7	0.1	0.	0.1	1.9	100.	
	101_012	К	07	8	8	- 69	7	34	4	2	05	9	8	00	
	<b>TO</b>	Ŧ	45.	0.8	4.4	11.	0.1	31.	5.1	0.1	0.	0.2	0.4	100.	
	102	L	18	9	8	84	7	35	8	4	18	0	1	00	
			45	0.8	44	11	0.1	31	51	0.1	0	0.2	07	100	
	TO2_Cr	Μ	01	8		80	7	23	6	4	18	0	9	00	
			11	0.8		11	01	30	51	01	0	$\hat{0}_{2}$	10	100	
	TO2_Cr2	Ν	~ 16	0.0	1.7	66	7	90. 85	0	1	18	0.2	0	100.	
	and hatio		40	<u> </u>	1	00	/	85	0	4	10	0	0	00	
	syninelic	ĺ	>	/											
	material			0 <b>-</b>	•	10	0.0	21		0.1		0.1	1.0	100	
	TOI_syn_	0	45.	0.7	3.8	10.	0.2	31.	4.6	0.1	1.	0.1	1.0	100.	
	Crl		89	4	4	22	0	91	6	5	08	6	0	00	
	TO1_syn_	Р	45.	0.7	3.8	10.	0.1	31.	4.6	0.1	1.	0.1	1.9	100.	
	Cr2	-	44	3	0	12	9	60	2	5	06	6	8	00	
	TO2_syn_	0	43.	2.2	4.8	11.	0.2	34.	0.5	0.2	0.	0.2	1.0	100.	
	Cr1	Q	73	6	6	60	4	87	6	2	20	1	0	00	
Á	TO2_syn_	р	43.	2.2	4.8	11.	0.2	34.	0.5	0.2	0.	0.2	1.9	100.	
( )	Cr2	К	30	3	1	49	4	52	6	2	20	1	8	00	
	Nv17 svn	<i>c</i> .	38.	1.2	6.5	10.	0.2	32.	7.7	0.9	0.	0.4	1.0	100.	
	Cr1	S	42	6	3	40	4	07	4	9	68	3	0	00	
<u> </u>	Ny17 syn		38	12	64	10	$0^{2}$	31	76	0,0	0	04	19	100	
	Cr?	Т	04	1.2 5	7	30	0.2 Л	75	6	Q.)	67	2	1.) Q	00.	
	CIZ		0 <del>4</del> 45	0.0	06	12	+ 0 1	ני רי	76	10	07	01	05	100	
	2Cr'+1Cr	U	4J.	0.0	9.0	12.	0.1	22. 71	7.0	1.0	0.	1	0.5	100.	
			UI	4	9	31	U	/1	/	U	04	1	2	00	

1502 <sup>a</sup> Indexes are used to indicate which starting composition was used during the experiments in Table. 2.

1503 1504

\* Indicates starting composition was doped with forsterite or MgO to increase liquidus.

Table. 2. Experimental conditions of extended,	re-measured and literature data.
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	Run No.	Reference	T (°C)	Pressure (MPa)	logfO 2	ΔQF M	H2O (wt%)	Duration (h)
	This study				2		<u> </u>	
	CAG03-K <sup>a, b</sup>	This study	1606.0 0	0.10	-4.53	0.00	0.00	7.00
	CAG03-N <sup>a</sup>	This study	1606.0 0	0.10	-4.53	0.00	0.00	7.00
	CAG03-H <sup>a</sup>	This study	1606.0 0	0.10	-4.53	0.00	0.00	7.00
	CAG04-N <sup>a</sup>	This study	1568.0 0	0.10	-4.83	0.00	0.00	7.00
	CAG04-E <sup>a</sup>	This study	1568.0	0.10	-4.83	0.00	0.00	7.00
	CAG05-R <sup>a</sup>	This study	1571.0	0.10	-4.81	0.00	0.00	24.00
	CAG05-S <sup>a</sup>	This study	1571.0 0	0.10	-4.81	0.00	0,00	24.00
	CAG06-P <sup>a</sup>	This study	1571.0 0	0.10	-4.81	0.00	0.00	24.00
	CAG06-R <sup>a</sup>	This study	1571.0 0	0.10	-4.81	0.00	0.00	24.00
	CAG06-T <sup>a</sup>	This study	1571.0	0.10	-4.81	0.00	0.00	24.00
	CAG07-T <sup>a</sup>	This study	1600.0	0.10	-4.58	0,00	0.00	24.00
	NAB01-C	This study	1190.0 0	0.10	-8.53	0.00	0.00	72.00
	NAB01-G	This study	1190.0	0.10	-8.53	0.00	0.00	72.00
	NAB01-F	This study	1190.0	0.10	-8.53	0.00	0.00	72.00
	NAB02-D	This study	1174.0	0.10	-8.72	0.00	0.00	94.00
	NAB02-G	This study	1174.0	0.10	-8.72	0.00	0.00	94.00
	NAB02-F	This study	1174.0	0.10	-8.72	0.00	0.00	94.00
	NAB21-A	This study	1450.0	0.10	-5.84	0.00	0.00	24.00
	NAB24-I	This study	1401.0	0.10	-6.29	0.00	0.00	72.00
	NAB24-L	This study	1401.0	0.10	-6.29	0.00	0.00	72.00
	NAB31-L	This study	1401.0	0.10	-4.29	2.00	0.00	48.00
	NAB31-A	This study	1401.0	0.10	-4.29	2.00	0.00	48.00
	NAB41-J	This study	1469.0	0.10	-5.67	0.00	0.00	48.00
	NAB41-M	This study	1469.0	0.10	-5.67	0.00	0.00	48.00
	NAB45-K <sup>a, c</sup>	This study	1529.0	0.10	-5.15	0.00	0.00	24.00
	NAB45-N <sup>a, c</sup>	This study	1529.0	0.10	-5.15	0.00	0.00	24.00
(	NAB45-B <sup>a, c</sup>	This study	1529.0	0.10	-5.15	0.00	0.00	24.00
	NAB45-H <sup>a, c</sup>	This study	1529.0	0.10	-5.15	0.00	0.00	24.00
	NAB49-Q	This study	1500.0	0.10	-5.40	0.00	0.00	24.00
	NAB49-R	This study	1500.0	0.10	-5.40	0.00	0.00	24.00
$\mathbf{\nabla}$	NAB51-R	This study	1500.0	0.10	-5.40	0.00	0.00	24.00
	NAB51-S	This study	1500.0	0.10	-5.40	0.00	0.00	24.00
	NAB52-R <sup>a</sup>	This study	1529.0	0.10	-5.15	0.00	0.00	24.00
	NAB53-O	This study	1445.0	0.10	-5.89	0.00	0.00	72.00
	NAB53-O	This study	0 1445.0	0.10	-5.89	0.00	0.00	72.00
	X		0		2.37			

	NAB53-R	This study	1445.0	0.10	-5.89	0.00	0.00	72.00	
	NAB53-S	This study	1445.0	0.10	-5.89	0.00	0.00	72.00	
	NAB57-R <sup>a</sup>	This study	1530.0 0	0.10	-5.15	0.00	0.00	24.00	
	NAB57-T <sup>a</sup>	This study	1530.0 0	0.10	-5.15	0.00	0.00	24.00	
	NAB58-R	This study	1397.0 0	0.10	-6.33	0.00	0.00	72.00	$\succ$
	*NAB88-R	This study	1350.0 0	0.10	-8.39	-1.60	0.00	118.00	$\searrow$
	NAB89-U *NAB89-R	This study	1253.0 0	0.10	-9.41	-1.60	0.00	120.00	
		This study	1253.0 0	0.10	-9.41	-1.60	0.00	120.00	
	NAB90-U	This study	1305.0 0	0.10	-8.85	-1.60	0.00	36.00	
	*NAB91-Q	This study	1305.0 0	0.10	-8.85	-1.60	0.00	36.00	
	*NAB91-R	This study	1305.0 0	0.10	-8.85	-1.60	0.00	36.00	
	Re-analyzed experim	ents		0.00					
	com1-12	Barr et al. (2009)	1320.0 0	0.10	-7.09	0.00	0.00		
	com1-2	Barr et al. (2009)	1350.0 0	0.10	-6.79	0,00	0.00		
	com1-4	Barr et al. (2009)	1300.0 0	0.10	-7.30	0.00	0.00		
	213.00	Thy (1995)	1240.0 0	0.10	7.95	0.00	0.00		
	214.00	Thy (1995)	1251.0 0	0.10	-7.83	0.00	0.00		
	215.00	Thy (1995)	1231.0 0	0.10	-8.05	0.00	0.00		
	216.00	Thy (1995)	1221.0 0	0.10	-8.17	0.00	0.00		
	218.00	Thy (1995)	1200.0 0	0.10	-8.41	0.00	0.00		
	224.00	Thy (1995)	$\begin{array}{c} 1260.0 \\ 0 \end{array}$	0.10	-7.73	0.00	0.00		
	228.00	Thy (1995)	1279.0 0	0.10	-7.52	0.00	0.00		
	230.00	Thy (1995)	1300.0 0	0.10	-7.30	0.00	0.00		
	233.00	Thy (1995)	1311.0 0	0.10	-7.18	0.00	0.00		
	234.00	Thy (1995)	1320.0 0	0.10	-7.09	0.00	0.00		
	235.00	Thy (1995)	1328.0 0	0.10	-7.01	0.00	0.00		
	c553	Mitchell & Grove (2015)	1215.0 0	1200.00	-8.78	-0.54	6.00		
	c559	Mitchell & Grove (2015)	1250.0 0	1200.00	-8.42	-0.59	3.70		
	on66-ha04	Charlier et al. (2018)	1230.0 0	500.00	- 10.58	-2.52	0.00		
( /	twm-f0.6-b1255	Charlier et al. (2018)	1360.0 0	1350.00	-9.03	-2.34	0.00		
	twm71-b1295	Charlier et al. (2018)	1300.0 0	800.00	-9.72	-2.42	0.00		
	twm71-ha04	Charlier et al. (2018)	1230.0 0	500.00	- 10.58	-2.52	0.00		
	w7	Parman & Grove (2004)	1320.0 0	1200.00	-8.09	-1.00	7.40		
	Literature data								
	w39	Wan et al. (2008)	1250.0 0	0.10	-9.55	-1.71	0.00		
	w83_1	Wan et al. (2008)	1250.0 0	0.10	-9.55	-1.71	0.00		
	w83_2	Wan et al. (2008)	1250.0 0	0.10	-9.55	-1.71	0.00		
	w83_3	Wan et al. (2008)	1250.0 0	0.10	-9.55	-1.71	0.00		

w83_4	Wan et al. (2008)	1250.0 0	0.10	-9.55	-1.71	0.00
w83_5	Wan et al. (2008)	1250.0 0	0.10	-9.55	-1.71	0.00
w83_6	Wan et al. (2008)	1250.0	0.10	-9.55	-1.71	0.00
w83_7	Wan et al. (2008)	1250.0	0.10	-9.55	-1.71	0.00
w35	Wan et al. (2008)	1300.0 0	0.10	-8.95	-1.65	0.00
w71_1	Wan et al. (2008)	1300.0 0	0.10	-8.95	-1.65	0.00
w71_2	Wan et al. (2008)	1300.0 0	0.10	-8.95	-1.65	0.00
w72_1	Wan et al. (2008)	1300.0 0	0.10	-8.95	-1.65	0.00
w72_2	Wan et al. (2008)	1300.0 0	0.10	-8.95	-1.65	0.00
w73	Wan et al. (2008)	1300.0 0	0.10	-8.95	-1.65	0.00
w36_1	Wan et al. (2008)	1350.0 0	0.10	-8.35	-1.56	0.00
w36_2	Wan et al. (2008)	1350.0 0	0.10	-8.35	-1.56	0.00
w55	Wan et al. (2008)	1350.0 0	0.10	-8.35	-1.56	0.00
w55_1	Wan et al. (2008)	1350.0 0	0.10	-8.35	-1.56	0.00
w55_2	Wan et al. (2008)	1350.0 0	0.10	-8,35	-1.56	0.00
w62_1	Wan et al. (2008)	1350.0 0	0.10	8.35	-1.56	0.00
w62_2	Wan et al. (2008)	1350.0 0	0.10	-8.35	-1.56	0.00
w66	Wan et al. (2008)	1350.0 0	0.10	-8.35	-1.56	0.00
w66_1	Wan et al. (2008)	1350.0 0	0.10	-8.35	-1.56	0.00
w61	Wan et al. (2008)	1400.0 0	0.10	-7.85	-1.55	0.00
w61_1	Wan et al. (2008)	1400.0 0	0.10	-7.85	-1.55	0.00
w86	Wan et al. (2008)	1400.0 0	0.10	-7.85	-1.55	0.00
w86_1	Wan et al. (2008)	1400.0 0	0.10	-7.85	-1.55	0.00
w86_2	Wan et al. (2008)	1400.0 0	0.10	-7.85	-1.55	0.00
w60_1	Wan et al. (2008)	1450.0 0	0.10	-7.35	-1.51	0.00
w60_2	Wan et al. (2008)	1450.0 0	0.10	-7.35	-1.51	0.00
w78	Wan et al. (2008)	1450.0 0	0.10	-7.35	-1.51	0.00
w328*	Coogan et al. (2014)	1250.0 0	0.10	-0.68	7.16	0.00
w328*_1	Coogan et al. (2014)	1250.0 0	0.10	-0.68	7.16	0.00
w290_1	Coogan et al. (2014)	1350.0 0	0.10	-0.68	6.11	0.00
w330	Coogan et al. (2014)	1350.0 0	0.10	-0.68	0.49	0.00
w330_1	Coogan et al. (2014)	1350.0 0	0.10	-6.30	0.49	0.00
w330_2	Coogan et al. (2014)	1350.0 0	0.10	-6.30	0.49	0.00
w332	Coogan et al. (2014)	1350.0 0	0.10	-7.30	-0.51	0.00
w332_1	Coogan et al. (2014)	1350.0 0	0.10	-7.30	-0.51	0.00
w338	Coogan et al. (2014)	1350.0 0	0.10	-5.50	1.29	0.00
matzen2011_15	Matzen et al. (2011)	1349.0 0	0.10	-6.87	-0.07	0.00
matzen2011_6	Matzen et al. (2011)	1302.0	0.10	-7.36	-0.08	0.00

matzen2011_8	Matzen et al. (2011)	1302.0 0	0.10	-7.36	-0.08	0.00	
matzen2011_46	Matzen et al. (2011)	1300.0 0	0.10	-7.41	-0.11	0.00	
H&J_1998_FAD1b	Hanson & Jones (1998)	1320.0 0	0.10	-3.00	4.09	0.00	
H&J_1998_FAD2	Hanson & Jones (1998)	1320.0 0	0.10	-3.00	4.09	0.00	
H&J_1998_FAD3	Hanson & Jones (1998)	1320.0 0	0.10	-3.00	4.09	0.00	
H&J_1998_FAS1	Hanson & Jones (1998)	1320.0 0	0.10	-3.00	4.09	0.00	
H&J_1998_FAS1_	Hanson & Jones (1998)	1320.0	0.10	-3.00	4.09	0.00	

<sup>a</sup> indicates 1–2 hour thermal oscillation (±10°C) was applied;
 <sup>b</sup> Experiment number ends with the starting composition noted in Table. 1;
 <sup>c</sup> indicates chromite seeds were applied.
 \* indicates experiments contain clinopyroxene

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Table. 3. Scenarios for selecting the appropriate model for OSAT.

	Scenarios	Model selection		
	Compare 3 models (Eq. 5, Eq. 7 & Eq. 8)			
	$S1: T\_Eq5 > T\_Eq7, T\_Eq5 > T\_Eq8$	Eq.5		
	If not S1, apply Z-test:			
	$Z_{Eq5\&Eq7} > Z_{Eq5\&Eq8} > 1.35$	Eq.7		
	$Z_{Eq5\&Eq8} > Z_{Eq5\&Eq7} > 1.35$	Eq.8		
	$Z_{Eq5\&Eq8} < 1.35, Z_{Eq5\&Eq7} < 1.35$	Eq.5		
	$Z_{Eq5\&Eq7} > 1.35 > Z_{Eq5\&Eq8}$	Eq.7		
	$Z_{Eq5\&Eq8} > 1.35 > Z_{Eq5\&Eq7}$	Eq.8		
	Compare 2 models (Eq. 5 & Eq. 7)			
	S2: T_Eq5 > T_Eq7	Eq.5		
	If not S2, apply Z-test:			
	Z_(Eq5&Eq7) > 1.35	Eq.7	$\langle \rangle$	$\sim$
	Z_(Eq5&Eq7) < 1.35	Eq.5		<b></b>
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Table. 4. Comparison of the temperature estimates of models.

		_	Eq.	. 5	Eq	. 7	Eq.	8	Model of C al. (20	boogan et 014)
	Location	observa tions	Min- Max (°C)	$T_{cry}$ (°C) <sup>a</sup>	Mın- Max (°C)	<i>T<sub>cry</sub></i> (°C)	M1n- Max (°C)	<i>T<sub>cry</sub></i> (°C)	Min-Max (°C)	<i>T<sub>cry</sub></i> (°C)
	MORB (Siqueiros)	22	1167- 1296	$1250^{+4}_{-7}$	1096- 1260	$1197^{+6:}_{-8}$	1115- 1278	$1218^{+5}_{-8}$	1121- 1292	1230+59
	Iceland	85	1145- 1362	$1274_{-1}^{+5}$	1112- 1350	$1247^{+8}_{-8}$	1169- 1345	$1264_{-6}^{+6}$	1183- 1383	1283 <sup>+84</sup> -66
	Skye	35	1313- 1436 1131-	1388_5	1290- 1425 1061-	1348_4!	1307- 1436 1166-	1376 <sup>+5</sup>	1337- 1474 1188-	1409 <sup>+60</sup> -65
	Emeishan	48	1391 926-	$1202^{+1}_{-6}$	1361 1157-	$1145^{+1}_{-6}$	1397 1178-	$1238_{-4}^{+9}$	1439 1197-	$1271_{55}^{+10}$
	Etendeka	62 157	1420 1109-	$11/8_{-1}^{+8}$	1447 1191-	1268 <u>-9</u>	1455 1217-	$1295_{-8}^{+7}$	1508 1250-	1492+89
1515	<sup>a</sup> $T_{cry}$ is report	ted as media	1506 n value wi	th 95 <sup>th</sup> and 5	1471 <sup>th</sup> percent	ile quoted.	1522	1720-1	1601	
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1517 1518 Table. 5. Summary of the posterior distributions of the  $T_p$ ,  $\phi_{Px}$ ,  $\phi_{Hz}$  in the inversion model as a comparison with the results from Matthews et al. (2021).

			This study		Matt	hews et al. (2	.021)	-
Location	$T_{cry} (^{\circ}\mathrm{C})^{\mathrm{a}}$	$T_p$ (°C)	${\it \Phi}_{Px}$	$arPsi_{ extsf{Hz}}$	$T_p$ (°C)	${\it \Phi}_{Px}$	$arPsi_{ extsf{Hz}}$	
MORB	1266±22	$1355^{+24}_{-19}$	$0.02\substack{+0.02 \\ -0.02}$	$0.34\substack{+0.25 \\ -0.30}$	$1364^{+18}_{-23}$	$0.02\substack{+0.02 \\ -0.02}$	$0.40\substack{+0.16 \\ -0.37}$	-
Iceland	1363±24	$1518^{+21}_{-17}$	$0.08\substack{+0.04\\-0.04}$	$0.23\substack{+0.18 \\ -0.17}$	$1525^{+21}_{-18}$	$0.08\substack{+0.05 \\ -0.04}$	$0.27\substack{+0.16 \\ -0.20}$	
Skye	1436±24	$1550^{+80}_{-82}$	$0.12\substack{+0.16 \\ -0.10}$	$0.51\substack{+0.32 \\ -0.39}$	$1566^{+73}_{-70}$	$0.10\substack{+0.15 \\ -0.10}$	$0.47\substack{+0.28 \\ -0.39}$	
Emeishan	1391±24	$1481^{+87}_{-78}$	$0.10\substack{+0.14 \\ -0.09}$	$0.53\substack{+0.29 \\ -0.36}$	$1555^{+100}_{-97}$	$0.13\substack{+0.14 \\ -0.10}$	$0.55\substack{+0.31 \\ -0.38}$	
Etendeka	1447±34	$1577^{+106}_{-98}$	$0.11\substack{+0.19 \\ -0.10}$	$0.53\substack{+0.31 \\ -0.38}$	$1599^{+104}_{-79}$	$0.10\substack{+0.15 \\ -0.09}$	$0.51\substack{+0.32 \\ -0.38}$	
Tortugal	1506±24	$1648^{+193}_{-91}$	$0.15\substack{+0.29 \\ -0.13}$	$0.50\substack{+0.31 \\ -0.38}$	$1813^{+157}_{-149}$	$0.29\substack{+0.18 \\ -0.24}$	$0.45^{+0.32}_{-0.27}$	
Crystaniz				i model.			55	
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	$X_2 = X_{Mg^{2+}}^{TET} + 2X_{Mg^{2+}}^{OCT}$	$X_{Mg^{2+}}^{TET} = \frac{X_2 + s_1}{2}$
	$X_3 = X_{Cr^{3+}}^{TET} + X_{Cr^{3+}}^{OCT}$	$X_{Fe^{2+}}^{TET} = X_4 + s_2 + s_3 + s_4 - \frac{X_2 + s_1}{2}$
	$X_4 = 2X_{Ti^{4+}}^{OCT}$	$X_{Fe^{3+}}^{TET} = X_5 - s_4$
	$X_5 = X_{Fe^{3+}}^{TET} + 2X_{Fe^{3+}}^{OCT}$	$X_{AI}^{TET} = 1 - X_3 - X_4 - X_5 - s_2$
	$s_1 = X_{Mg^{2+}}^{TET} - 2X_{Mg^{2+}}^{OCT}$	$X_{Cr^{3+}}^{TET} = X_3 - s_3$
	$s_2 = \frac{2X_{Al^{3+}}^{OCT} - X_{Al^{3+}}^{TET}}{2}$	$X_{Mg^{2+}}^{OCT} = \frac{X_2 - s_1}{4}$
	$s_3 = \frac{2X_{Cr^{3+}}^{OCT} - X_{Cr^{3+}}^{TET}}{2}$	$X_{Fe^{2+}}^{OCT} = \frac{1 - s_2 - s_3 - s_4}{2} - \frac{X_2 - s_1}{4}$
	$s_4 = \frac{2X_{Fe^{3+}}^{OCT} - X_{Fe^{3+}}^{TET}}{2}$	$X_{Fe^{3+}}^{OCT} = \frac{X_5 + s_4}{2}$
	$X_{Fe^{2+}}^{TET} + X_{Fe^{2+}}^{TET} + X_{Al^{3+}}^{TET} + X_{Cr^{3+}}^{TET} + X_{Fe^{3+}}^{TET} = 1$	$X_{AI^{3+}}^{OCT} = \frac{1 - X_3 - X_4 - X_5 + s_2}{2}$
	$X_{Fe^{2^+}}^{OCT} + X_{Mg^2}^{OCT} + X_{AI^{3^+}}^{OCT} + X_{Cr^{3^+}}^{OCT} + X_{Fe^{3^+}}^{OCT} + X_{Ti^{4^+}}^{OCT} = 1$	$\mathbf{X}_{\mathrm{Grav}}^{\mathrm{OCT}} = \frac{\mathbf{X}_3 + \mathbf{s}_3}{2}$
		$X_{T1^{4+}}^{OCT} = \frac{X_4}{2}$
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Table. A1. Independent compositional, ordering variables and definitions of site mole fractions

		Coefficients <sup>a</sup>	Statistical pa	rameters
		$\overline{Eq}$	. 5	
	c0	-0.168	$r^2$	0.97
	c1	1.487	SEE	20.20
	c2	-0.593	RMSE	29.00
	c3	-0.630	<i>p</i> -value	1.11E-16
	c4	0.390	$r^2_{\rm MCCV}$	0.97
	c5	0.009	SEE <sub>MCCV</sub>	20.20
	сб	-2.492	RMSE <sub>MCCV</sub>	23.90
	c7	0.065		
	c8	-0.031		
	c9	-4.141		
	c10	-0.428		
	c11	-4.637		
	c12	-0.054		
	c13	-10.803		
	c14	2.714		
	b	0.654		
		Eq	. 7	
	Cr#	1.144(0.136)	$r^2$	0.83
	lnK <sub>DAl</sub>	-0.865(0.046)	SEE	44.80
	const	0.740(0.317)	RMSE	34.40
			<i>p</i> -value	2.88E-36
			$r^2_{\rm MCCV}$	42.10
			SEE <sub>MCCV</sub>	0.92
			RMSE <sub>MCCV</sub>	43.30
		Eq	. 8	
	Cr#	0.543(0.121)	$r^2$	0.91
	lnK <sub>DAl</sub>	-0.657(0.040)	SEE	31.70
	InK <sub>DCr</sub>	-0.389(0.041)	RMSE	39.10
	const	0.049(0.241)	<i>p</i> -value	3.79E-47
			$r_{\rm MCCV}^2$	0.94
. ( /	<b>S</b> Y		SEE <sub>MCCV</sub>	35.80
	<u>a 1</u>	• • • • •	RMSE <sub>MCCV</sub>	34.20
1526	for the li	r in bracket indic	cates the $1\sigma$ er	ror of coeffi
1528	for the h	ineur regression.		

**Table. A2.** Coefficients and statistical parameters of regressions for OSAT.