# The frontier of simulation-based inference

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# **Scientific simulators**





 $heta, z, x \sim p( heta, z, x)$ 



This results in the likelihood  $p(x| heta) = \int p(x,z| heta) dz$  to be intractable.

### **Problem statement**

Start with

- a simulator that lets you generate N samples  $x_i \sim p(x_i | heta_i)$ ,
- observed data  $x_{
  m obs} \sim p(x_{
  m obs}| heta_{
  m true})$ ,
- a prior  $p(\theta)$ .

Then,



# **Inference algorithms**



# **Inference algorithms**



# Approximate Bayesian Computation (ABC)



#### Issues

- How to choose x'?  $\epsilon$ ?  $|| \cdot ||$ ?
- No tractable posterior.
- Need to run new simulations for new data or new prior.

# **Neural Ratio Estimation (NRE)**

The Bayes rule can be rewritten as

$$p( heta|x) = rac{p(x| heta)p( heta)}{p(x)} = r(x| heta)p( heta) pprox \hat{r}(x| heta)p( heta),$$

where  $r(x| heta) = rac{p(x| heta)}{p(x)}$  is the likelihood-to-evidence ratio.

The ratio can be learned with machine learning, even neither the likelihood nor the evidence can be evaluated!



The solution d found after training approximates the optimal classifier

$$d(x, heta)pprox d^*(x, heta)=rac{p(x, heta)}{p(x, heta)+p(x)p( heta)}.$$

Therefore,

$$r(x| heta) = rac{p(x| heta)}{p(x)} = rac{p(x, heta)}{p(x)p( heta)} pprox rac{d(x, heta)}{1-d(x, heta)} = \hat{r}(x| heta).$$





### **Showtime!**

Some applications of physics and astrophysics.



#### **Case 1: Hunting new physics at particle colliders**



With enough training data, NRE gets the likelihood-ratio statistic right.

Using more information from the simulator improves sample efficiency substantially.

**Case 2: Dark matter substructure from gravitational lensing** 







#### **Case 3: Constraining dark matter with stellar streams**



Palomar 5 (Pal5) stream Pal5 was discovered in 2001 as the first thin stream formed from a globular cluster. Its current orbit takes it far over the galactic center.



#### GD1 stream -

Discovered in 2006, GD1 is the longest known thin stream, stretching across more than half the northern sky. It contains a gap that could Image chaitse spark of a dark matter collision 500 million years ago.

Gap

Milky Way



Preliminary results for GD-1 suggest a preference for CDM over WDM.

50

3.5

3.0

10

20

30

 $m_{\rm WDM}$ 

40

50

0.005

0.000

Ó

10

20

30

 $m_{\rm WDM}$ 

40

# **Diagnosing inference**

How to assess that approximate posteriors are not too wrong?

#### Coverage

- For every  $x, \theta \sim p(x, \theta)$  in a validation set, compute the  $1 \alpha$  credible interval based on  $\hat{p}(\theta|x) = \hat{r}(x|\theta)p(\theta)$ .
- The fraction of samples for which  $\theta$  is contained within the interval corresponds to the empirical coverage probability.

If the empirical coverage is larger that the nominal coverage probability  $1-\alpha$ , then the ratio estimator  $\hat{r}$  passes the diagnostic.





All benchmarked algorithms can produce non-conservative posterior approximations.

### **The frontier of**

### simulation-based inference

# Averting a crisis in simulationbased inference?

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# **Summary**

- Much of modern science is based on simulators making precise predictions, but in which inference is challenging.
- Machine learning enables powerful inference methods, which work in problems from the smallest to the largest scales.
- Advances in simulation-based inference will translate into scientific progress.
- However, further work is needed to make these methods more robust and reliable.

The end.