



ALBUS : Anomaly detector for Long duration BUrst Searches

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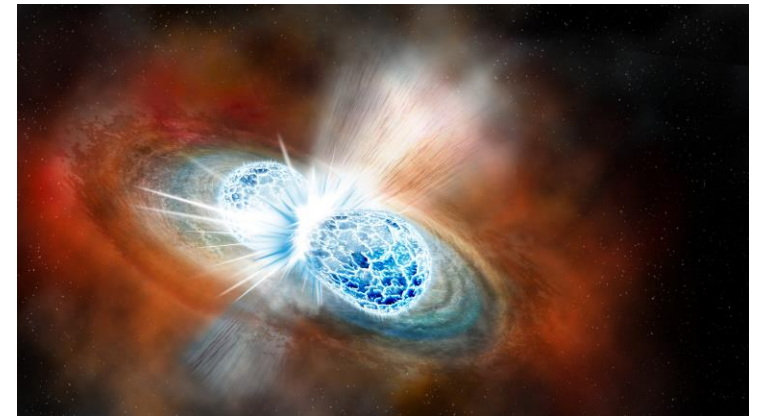
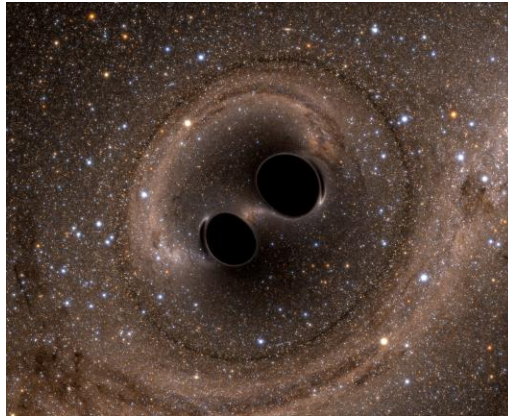
Table of Contents

- 1) What are Bursts ?
- 2) How do we detect them ?
- 3) Convolutional neural networks
- 4) New approach : mimic long-duration burst signals
- 5) Early results
- 6) Improvements and future plans

1) What are Bursts ?

- Gravitational waves detected so far : compact binary coalescences (CBC)

- Black Hole-Black Hole
- Neutron star-Neutron star
- Black Hole-Neutron star



- Expected class of events : Bursts
- Anything that is transient and not a CBC
- two families of bursts : short- (< 2 sec) and long duration (> 2 sec)

1) What are Bursts ?

- What are the phenomena generating long-duration bursts ?

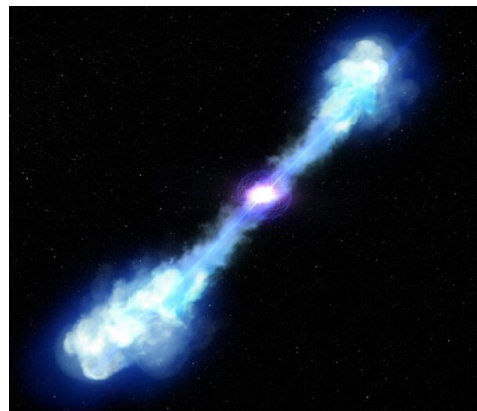
Non-axisymmetric deformations in magnetars



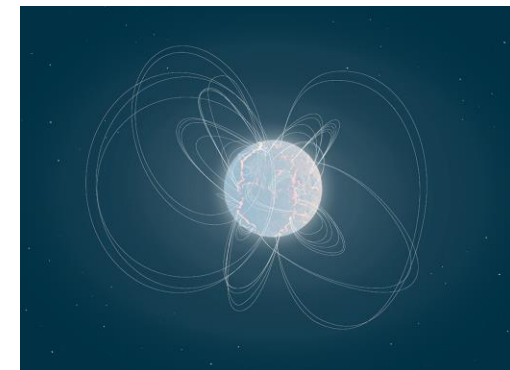
Accretion disk instabilities around black holes



Gamma-ray Bursts



Fallback accretion in newborn neutron stars



2) How do we detect them ?

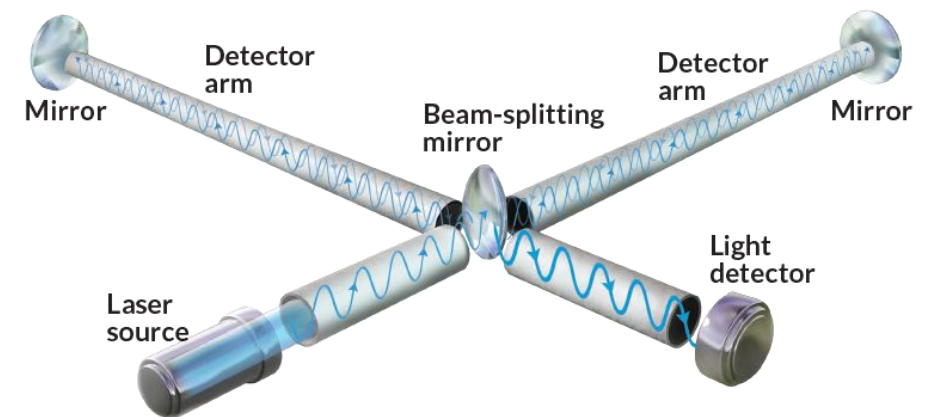
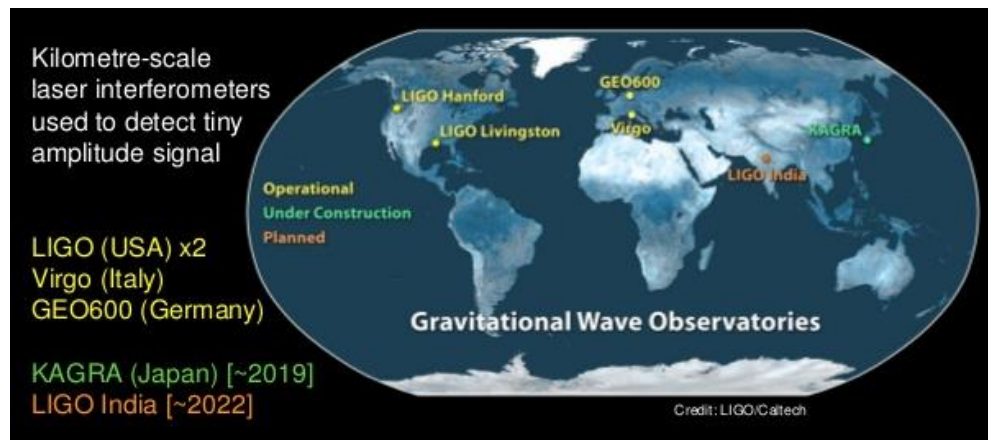
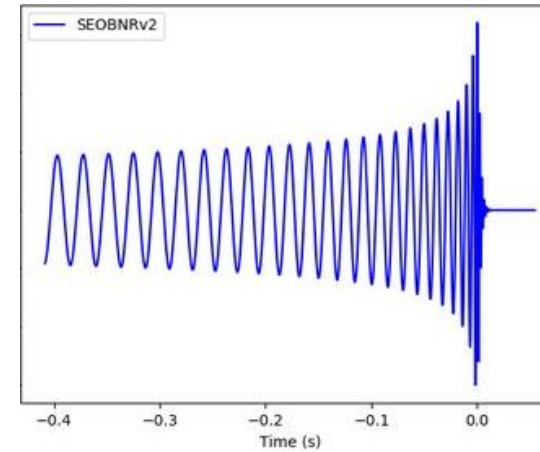
- CBC detection : general relativity \Rightarrow model of collision = waveform
 \Rightarrow then try to match those models to the data (matched filtering)

- Many other phenomena can generate GWs !

\Rightarrow But physics is poorly known...

\Rightarrow Models not accurate enough to apply match filtering.

Solution : use multiple detectors to find correlation in the data



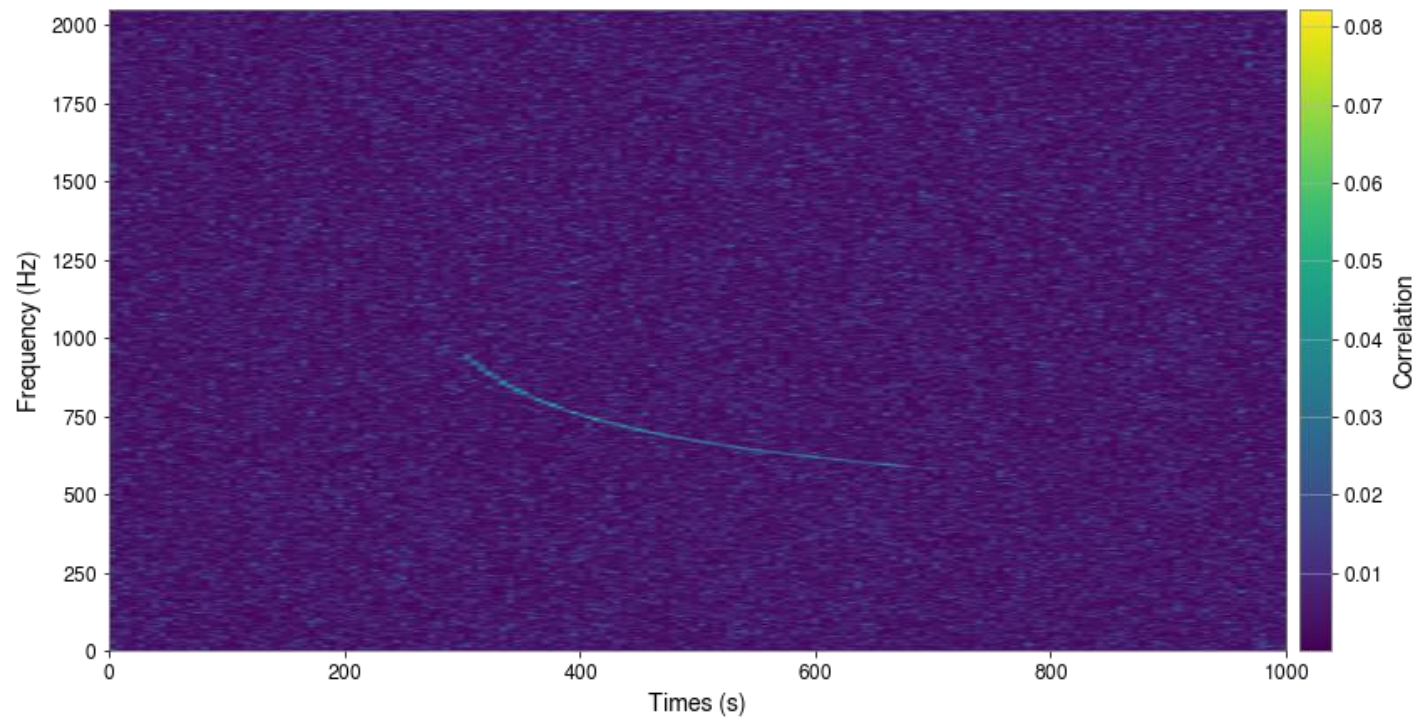
2) How do we detect them ?

- Excess of power method

=> Search in Time-Frequency space => minimal assumption : well represented in that TF space

=> Bursts should be clusters of high-correlation pixels

=> Many sources of noise (seismic, laser noise, suspensions, etc.)



3) Convolutional neural networks

- Class of artificial neural networks employing convolution

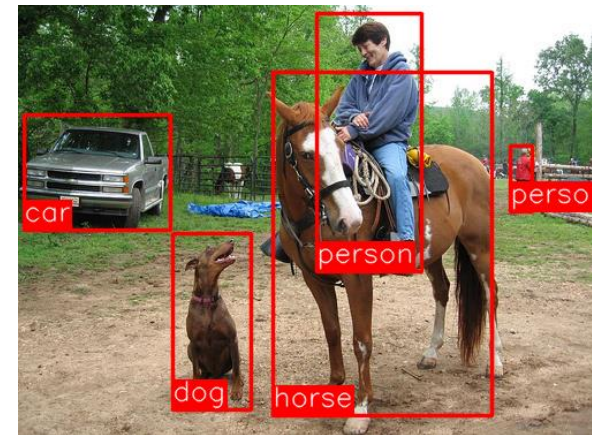
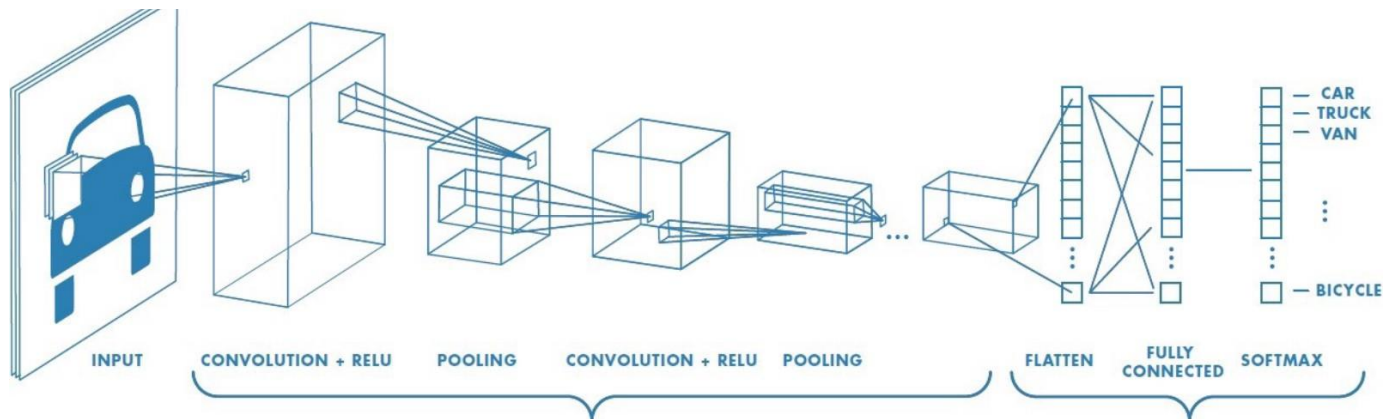
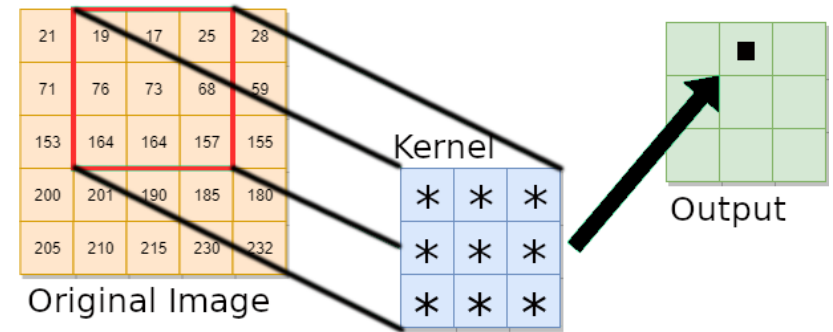
=> easy to use and understand

=> allows to downscale the information

- Image processing applications often require :

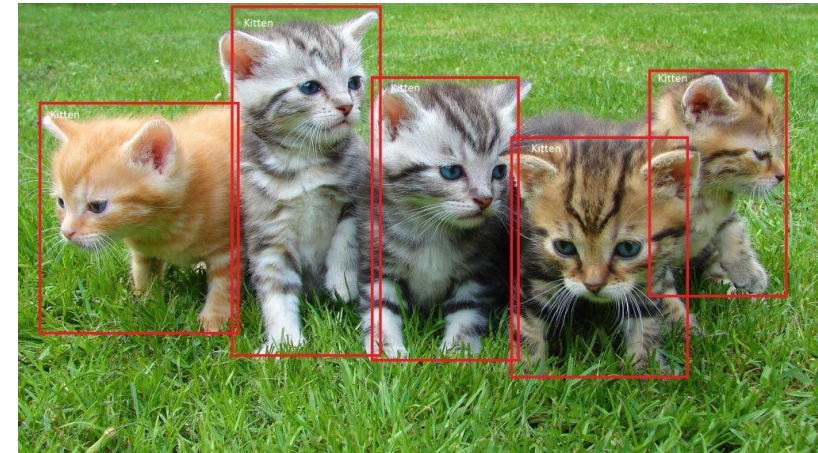
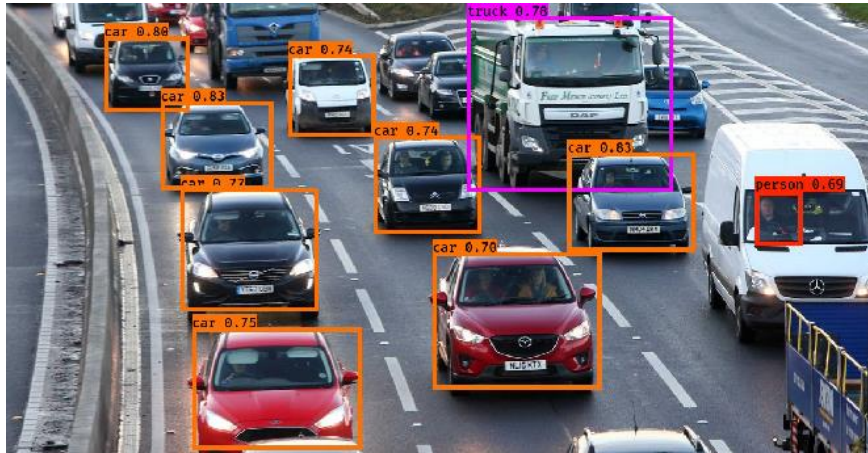
=> classification tasks (medical images, galaxy catalogs, ...)

=> bounding box determination (self-driven cars, face recognition, ...)



3) Convolutional neural networks

- Efficient at recognizing patterns and shapes :

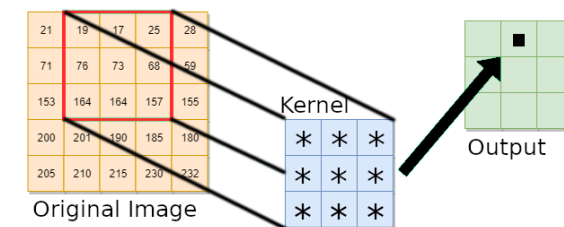


- Note : a neural network is nothing without a well-designed loss function !

=> loss function = what you want to minimize to achieve your goal (classification, prediction, ...)

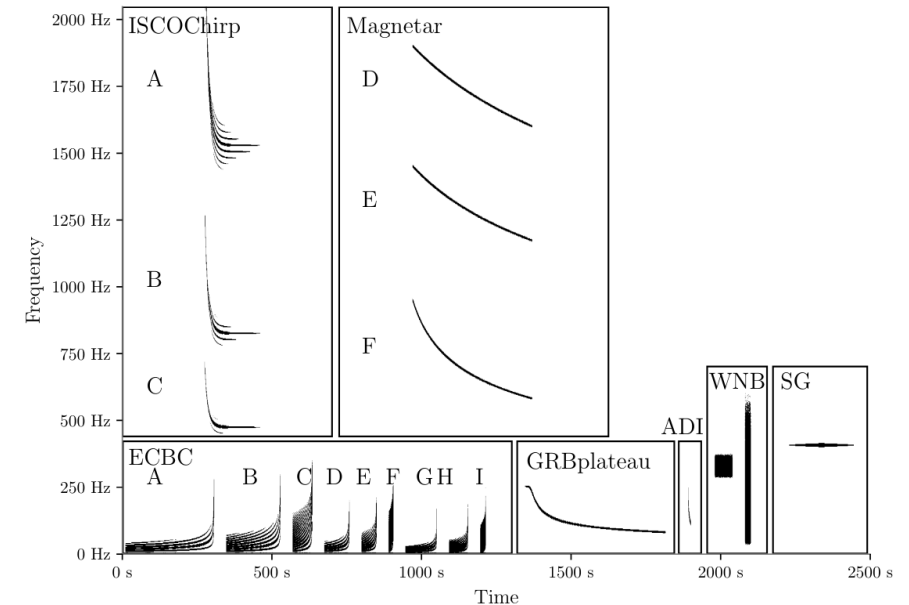
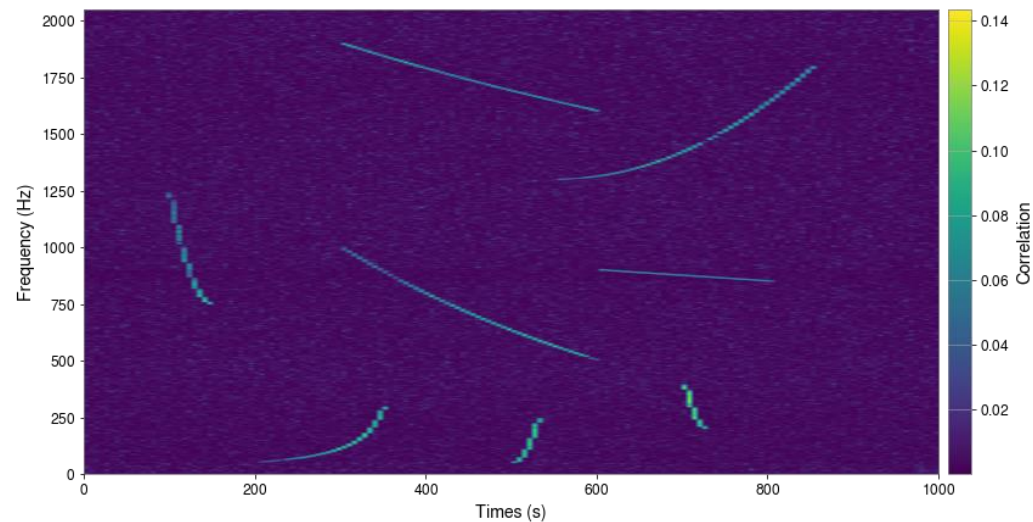
=> loss function gives feedbacks to update the weights (in kernels, ...)

=> bad weight updates = badly conditioned training = bad results



4) New approach : mimic long-duration burst signals

- Problem : can't rely on the long-duration models
 - too many uncertainties in the physical phenomena
 - models cannot be used as patterns to match for
- They all show a "chirp up" or "chirp down" behavior
 - ==> easily mimicked thanks to the *Python Scipy* library !
 - ==> Allow to generate chirps as time series

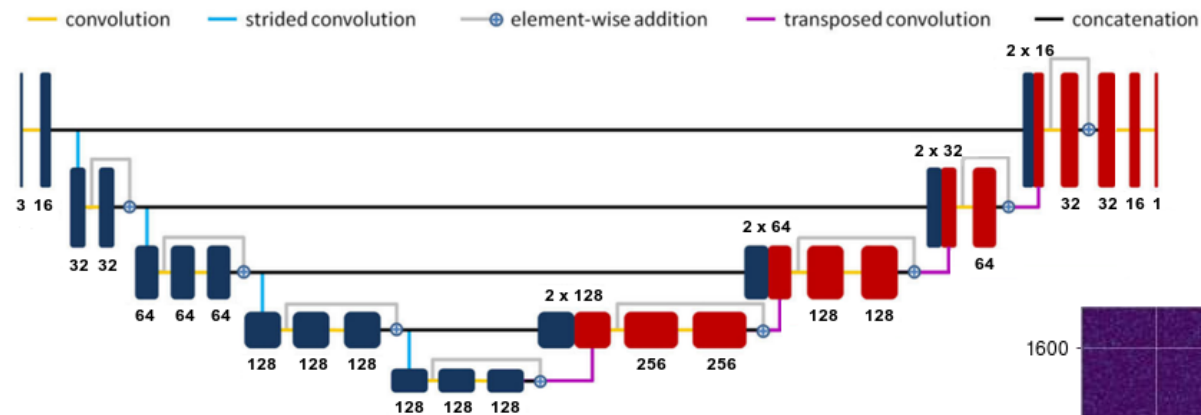


Taken from O3 long-duration paper :

https://dcc.ligo.org/public/0174/P2100078/011/o3_long_duration.pdf

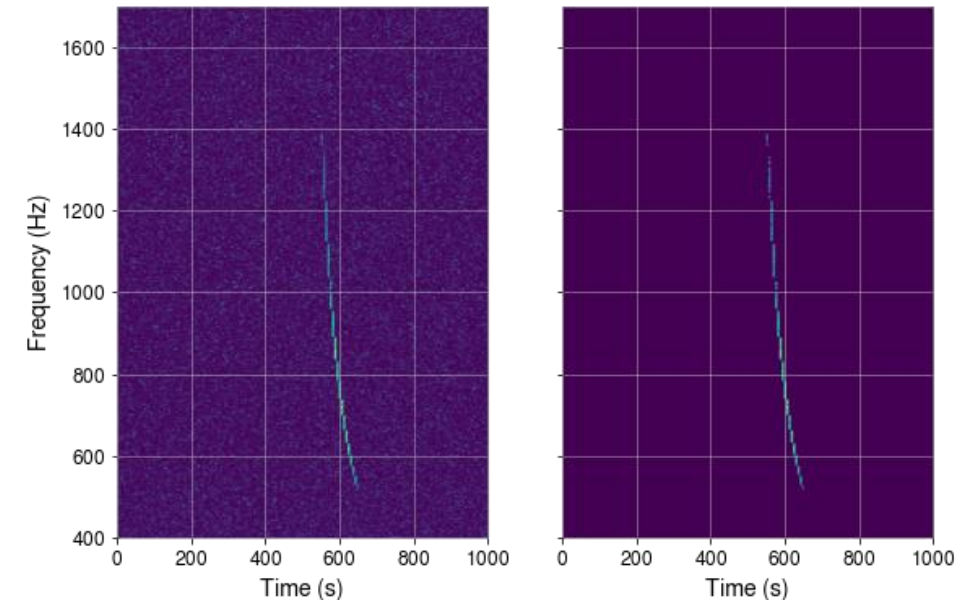
4) New approach : mimic long-duration burst signals

- Inspired by *Xing et al., 2019*. (<https://doi.org/10.1186/s12859-019-3037-5>), coded with PyTorch
- Downscaling and upscaling network



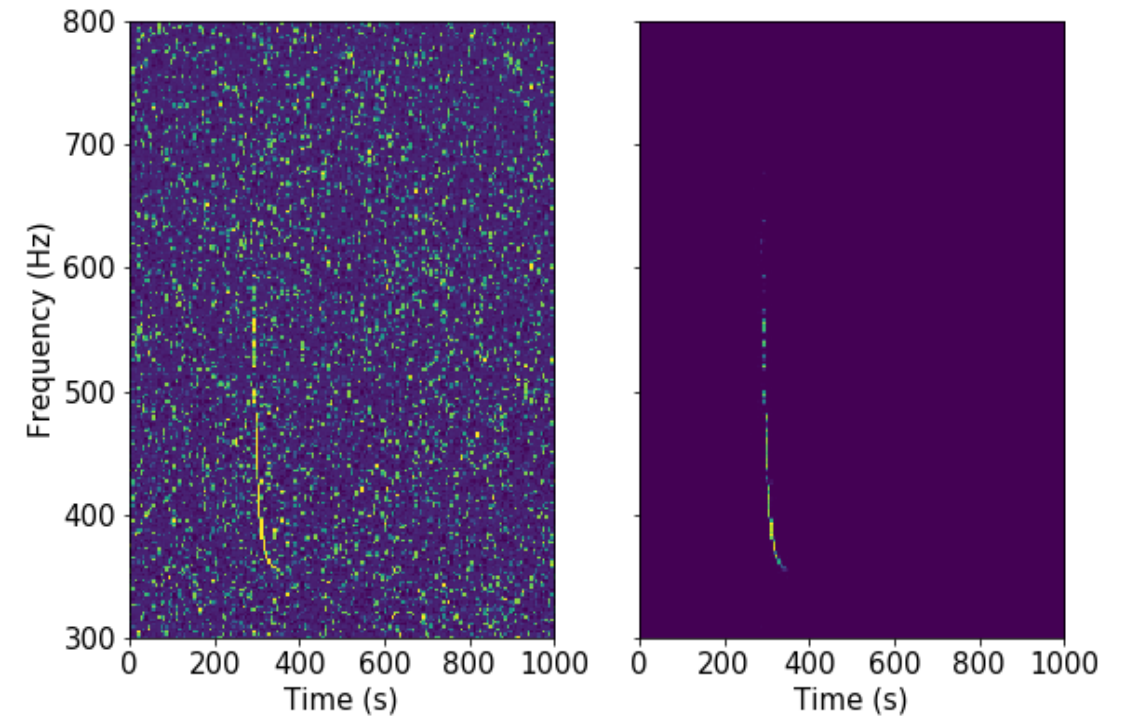
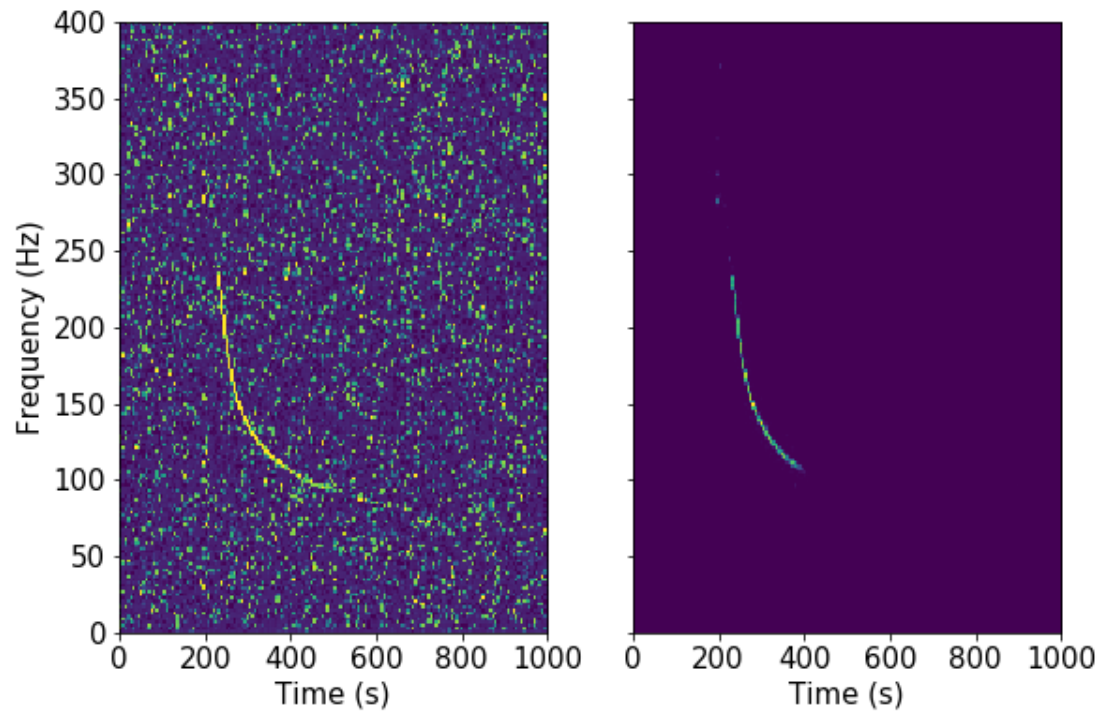
- Method :
 - train the network so that : output (O) \simeq target (T)
 - \implies our target will be injection in empty TF map
 - \implies Empty map for noise-only images

- Loss that is being minimized :
$$MSE = \frac{1}{2} \sum_{i,j} (T_{ij} - O_{ij})^2$$



5) Early Results

- Localization : Time-Frequency maps with injection

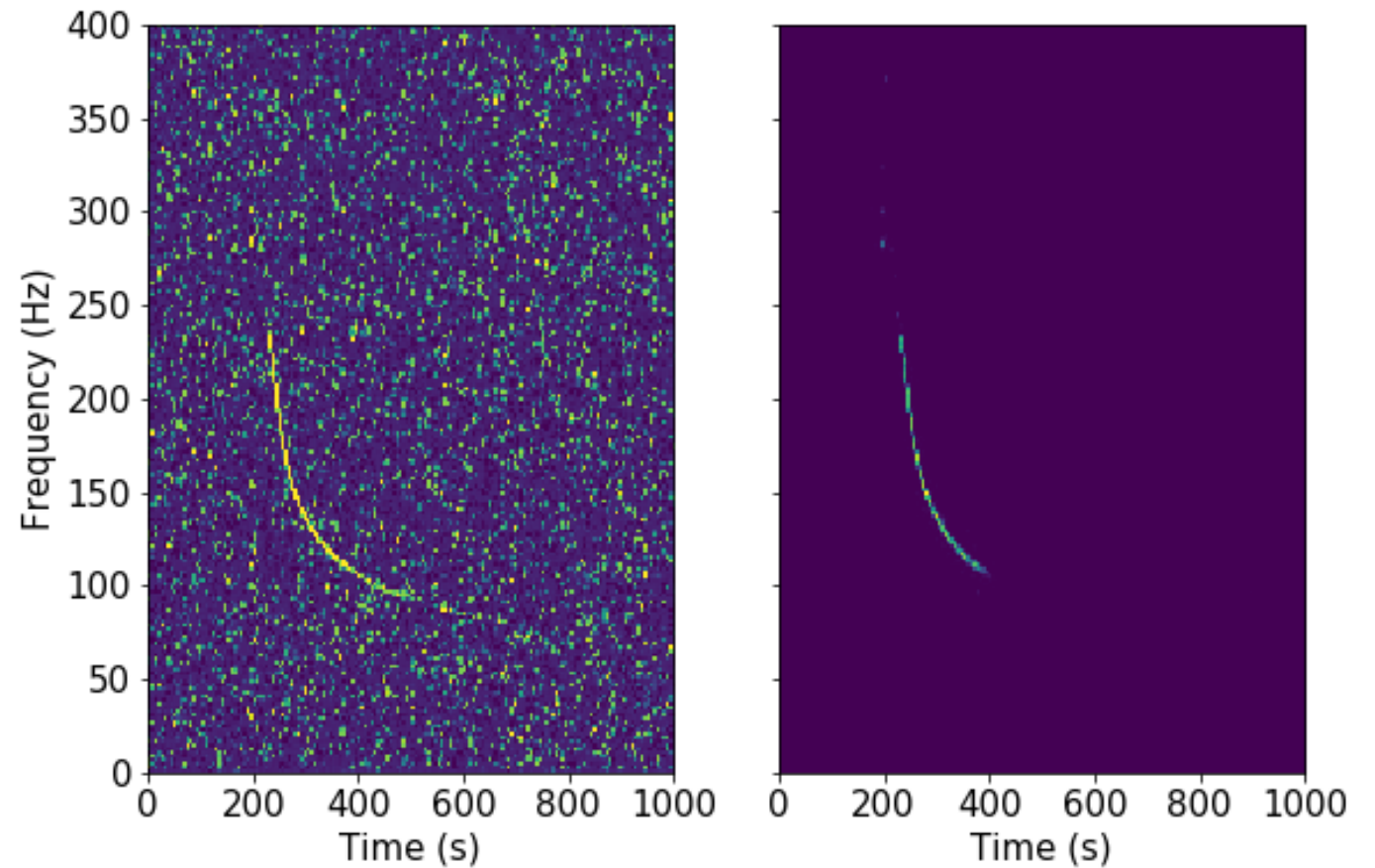


5) Early Results

- Localization : Time-Frequency maps with injection

- next step : learn the connectivity
between pixels

==> What about the time-frequency
maps with only pure noise ?

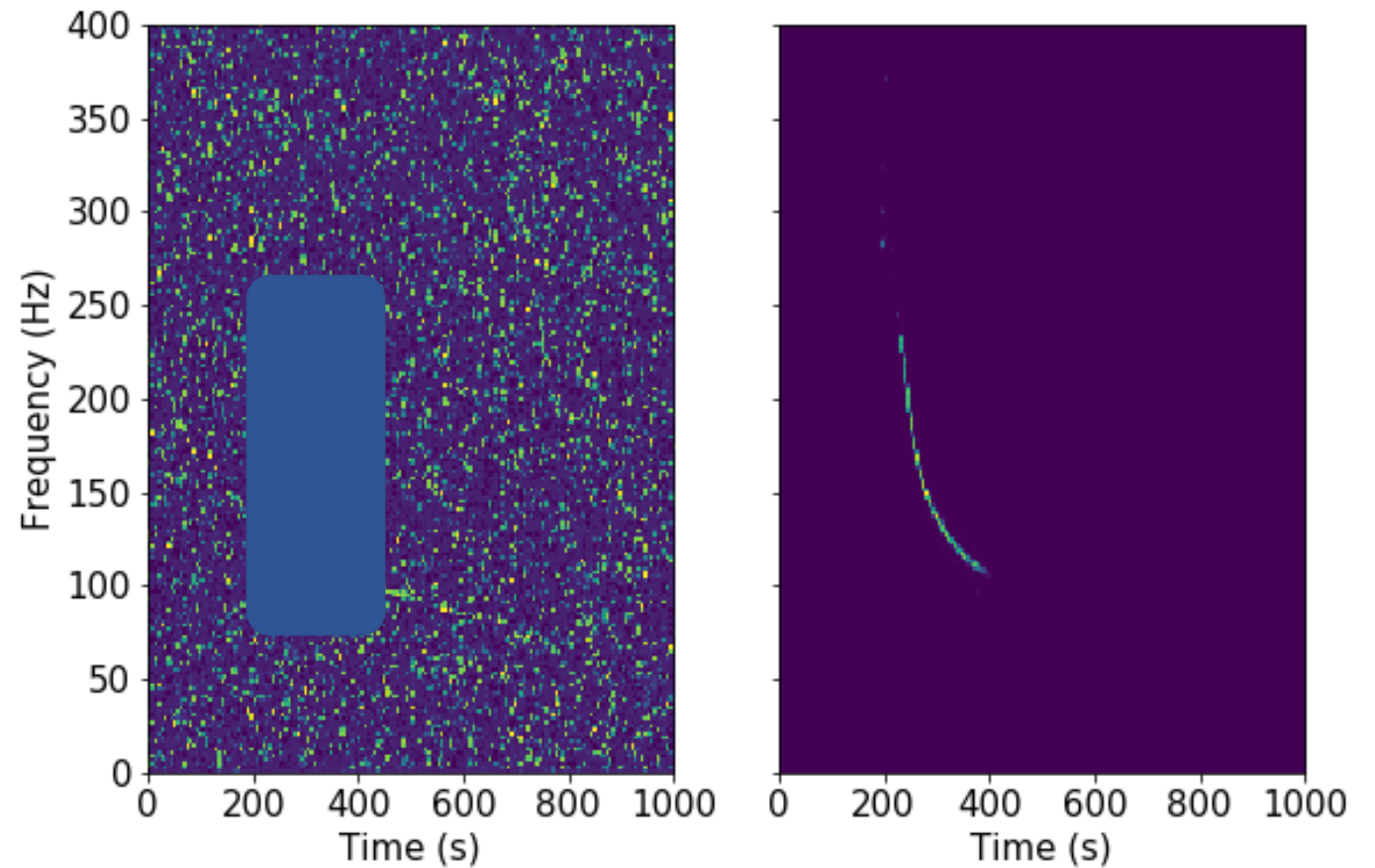


5) Early Results

- Localization : Time-Frequency maps with injection

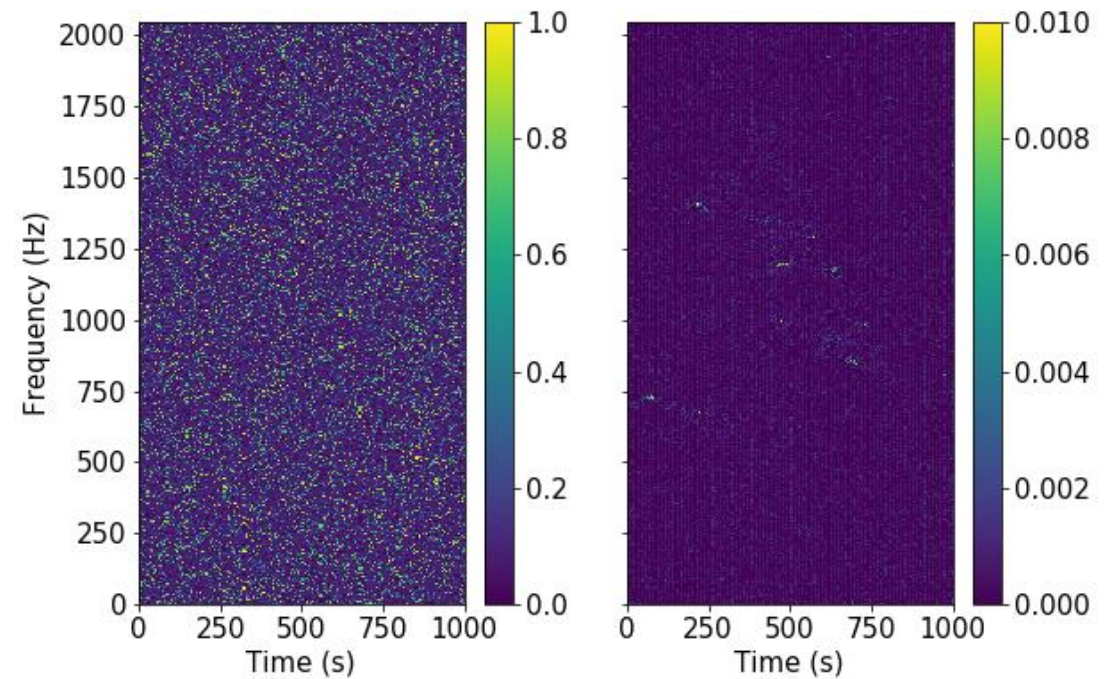
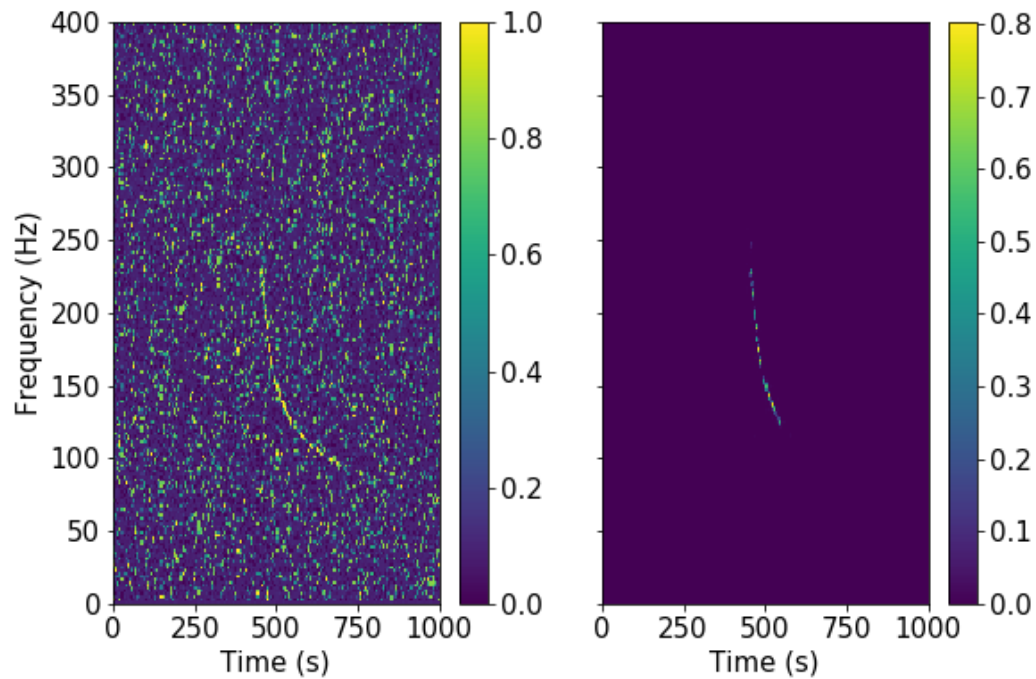
- next step : learn the connectivity
between pixels

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maps with only pure noise ?



5) Early Results

- Localization : Time-Frequency maps with pure noise
 - Values at least 1 order of magnitude lower than injection images
 - Sparse and uncorrelated pixels



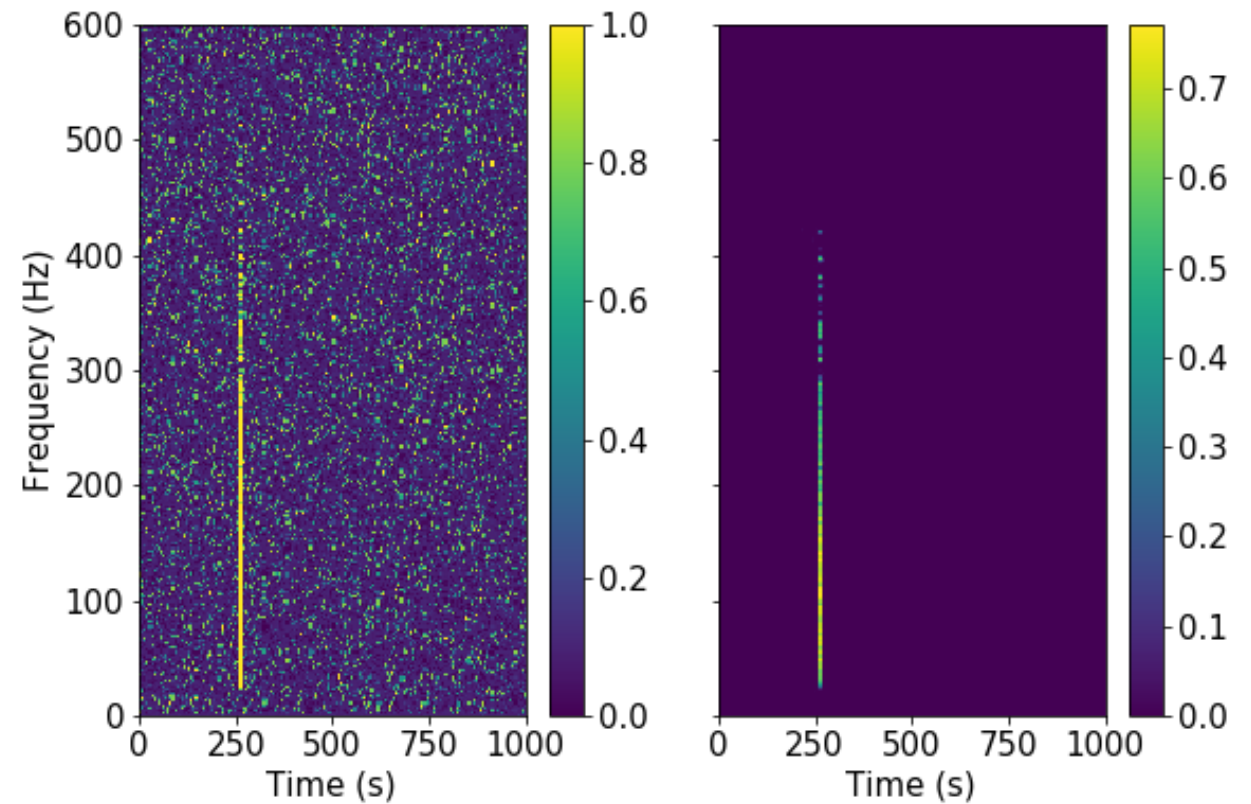
5) Early Results

- Localization : Time-Frequency maps with pure noise

- Instrumental/environmental noise transients (glitches) are detected !

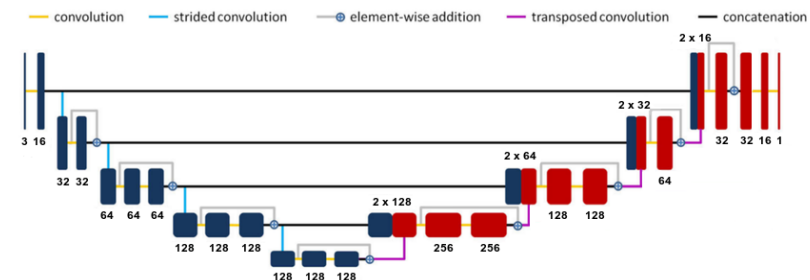
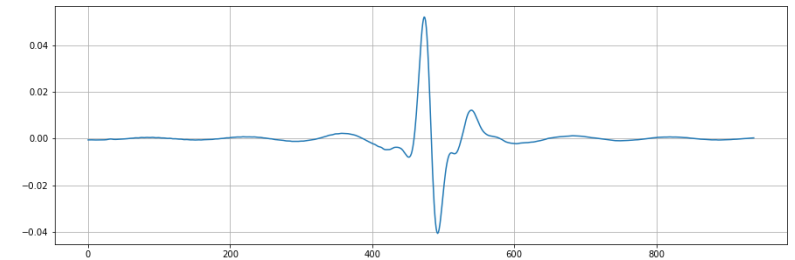
=> limit the sensitivity of our searches

=> need for a tool to remove them



6) Improvements and future plans

- State of the work : internal LVK review start by the end of November
- Implement new training method : Curriculum Learning (train with the easiest samples at first)
=> should increase the performances for low amplitude injections
- Add a classifier to remove glitches
=> see the work of Melissa Lopez and myself (paper out soon)
- Test on new problems (can be adapted to any image shape !)
=> GW background search, GW from supernovae, ...



THE END

Thank you for your attention !

Questions ?

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