

# The Distribution of Heat-producing Elements during Mercury Evolution

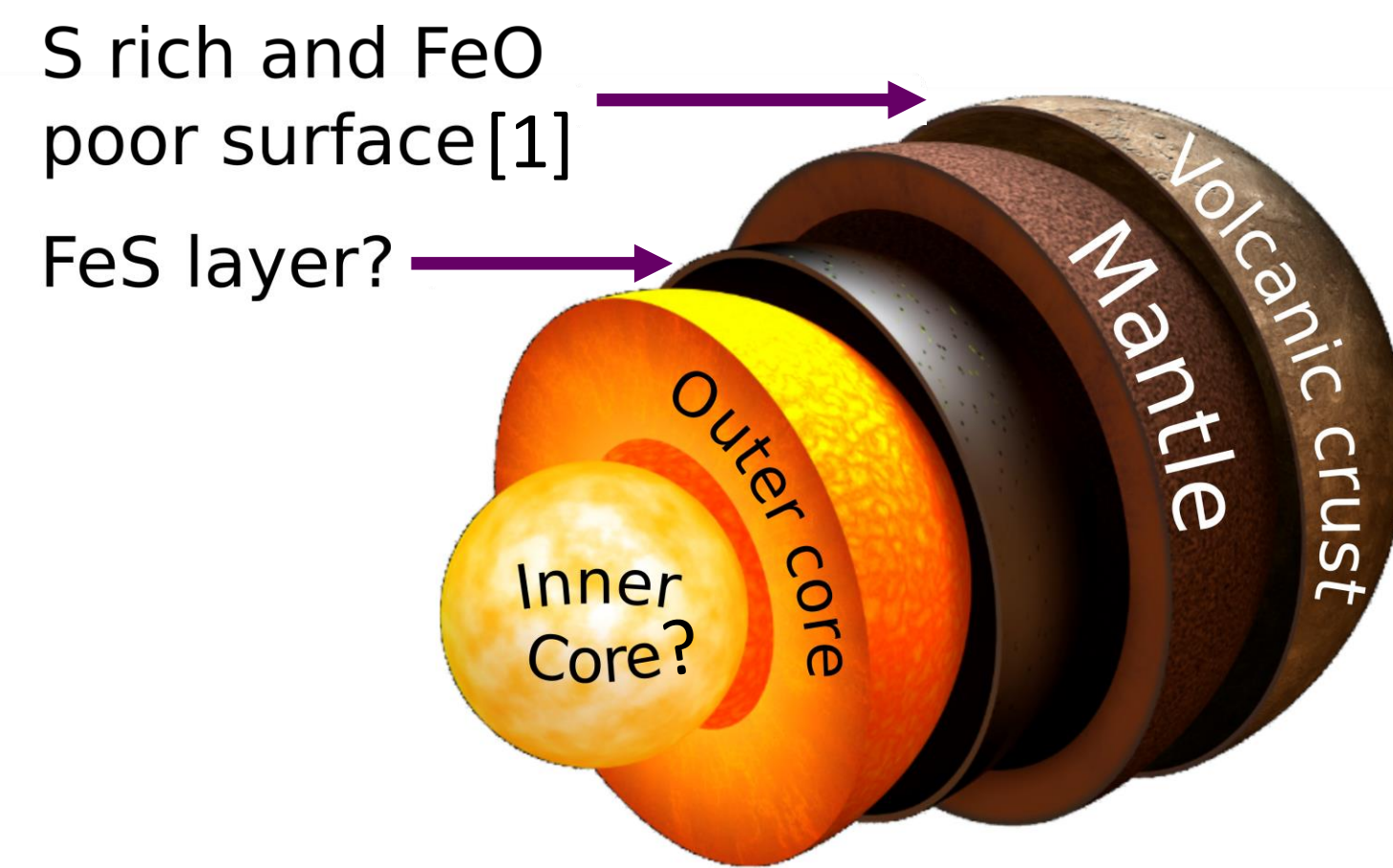
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## Motivation

To understand the distribution of U, Th and K in Mercury and link it with the planet's thermal evolution and its volatile budget.



The interior of Mercury. After Charlier & Namur, 2019 [2]

The major scientific questions are:

- Do sulfides in the mantle and crust of Mercury represent a reservoir for heat-producing elements?
- What is the core-mantle-crust distribution of heat-producing elements?

## Experimental approach

### Starting material: an analogue of Mercury's silicate mantle

Synthetic powder corresponding to the silicate part of an enstatite chondrite [3]

+ U, Th, K ; + Si metal; ± FeS, CaS, S

Different Si/SiO<sub>2</sub> ratio to control oxygen fugacity

Some experiments are iron-free

### Experimental equipment and conditions

Phase equilibria experiments were performed at pressure, temperature, and oxygen fugacity conditions relevant to Mercury's magma ocean using

- Internally Heated Pressure Vessel: 0.1 GPa, 1520 – 1600°C
- Piston-cylinder apparatus: 1 – 3 GPa, 1400 – 1700°C
- Multi-anvil apparatus: 6 GPa, 1700°C

Oxides	wt%
SiO <sub>2</sub>	61.26
TiO <sub>2</sub>	0.10
Al <sub>2</sub> O <sub>3</sub>	2.76
Cr <sub>2</sub> O <sub>3</sub>	0.82
Fe <sub>2</sub> O <sub>3</sub>	0.00
FeO	0.00
MnO	0.43
MgO	31.37
CaO	1.65
Na <sub>2</sub> O	1.55
K <sub>2</sub> O	0.19
P <sub>2</sub> O <sub>5</sub>	0.90
Total	101.03



The piston-cylinder in the University of Liège

### Analytical methods

Different methods were used to analyze the texture and chemistry of the samples recovered from experiments quenched at high temperature:

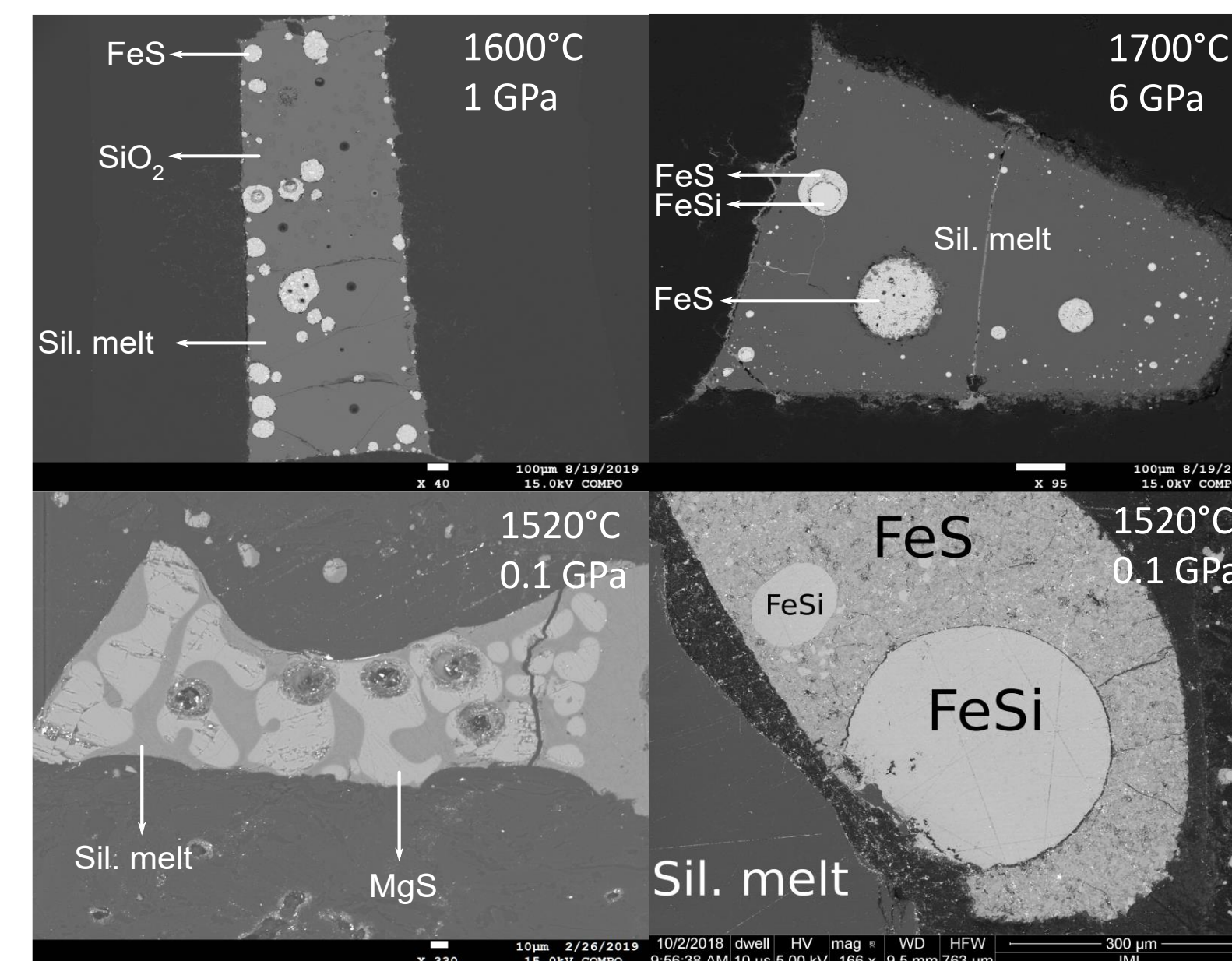
- Optical microscope, Scanning Electron Microscope (SEM) + Energy Dispersive Spectroscopy (EDS) for characterizing textures and phases
- Electron Probe Micro-Analyzer (EPMA) for quantifying the major and minor elements composition
- Laser Ablation – Inductively Coupled Plasma – Mass Spectrometry (LA-ICPMS) for quantifying minor and trace elements

### References:

[1] Weider S.Z., Nittler L.R., Starr R.D., Crapster-Pregont E.J., Peplowski P.N., Denevi B.W., Head J.W., Byrne P.K., Hauck S.A., Ebel D.S., Solomon S.C. (2015) Evidence for geochemical terranes on Mercury: Global mapping of major elements with MESSENGER's X-Ray Spectrometer. *Earth and Planetary Science Letters*. Vol. 416, 109-120. [2] Charlier B. & Namur O. (2019) The origin and differentiation of planet Mercury. *Elements*. Vol. 15, 9-14. [3] Berthet S., Malavergne V., Righter K. (2009) Melting of the Indarch meteorite (EH4 chondrite) at 1GPa and variable oxygen fugacity: implications for early differentiation processes. *Geochimica et Cosmochimica Acta*. Vol. 73 (20), 6402-6420. [4] Morard G. & Katsura T. (2010) Pressure-temperature cartography of Fe-S-Si immiscible system. *Geochimica et Cosmochimica Acta*. Vol. 74, 3659-3667. [5] Namur O., Charlier B., Holtz F., Cartier C., McCammon C. (2016) Sulfur solubility in reduced mafic silicate melts: Implications for the speciation and distribution of sulfur on Mercury. *Earth and planetary science letters*. Vol. 448, 102-114. [6] Boujibar A., Habermann M., Righter K., Ross D.K., Pando K., Righter M., Chidester B.A., Danielson L.R. (2019) U, Th and K partitioning between metal, silicate, and sulfide and implications for Mercury's structure, volatile content, and radioactive heat production. *American Mineralogist*. Vol. 104, 1221-1237. [7] Wohlers A. & Wood B. (2017) Uranium, thorium and REE partitioning into sulfide liquids: implications for reduced S-rich bodies. *Geochimica et Cosmochimica Acta*. Vol. 205, 226-244.

## Phases equilibria and textures

- The quenched samples present the following phases: silicate melts + FeS + Fe-Si (± SiO<sub>2</sub>, ± Si and ± enstatite)
- FeS and Fe-Si phases are large enough for high quality and numerous LA-ICP-MS data to be collected
- Sulfides rich in Mg, Ca, Cr, Mn and Fe are present in some experiments
- Other sulfides are more difficult to measure
- Immiscibility between FeS and Fe-Si. FeS surrounds Fe-Si as a result of wetting properties [4]
- In the Fe-free, CaS-doped experiments, metallic phases composed of Si, Cr, Mn, P are observed
- In some experiments at low  $fO_2$  conditions ( $\sim IW - 3$ ), Fe-P is sometimes present



Examples of backscattered SEM images of recovered samples. Experimental conditions indicated on the top right of each image. Samples are contained in graphite capsules

## Oxygen fugacity ( $fO_2$ ) determination

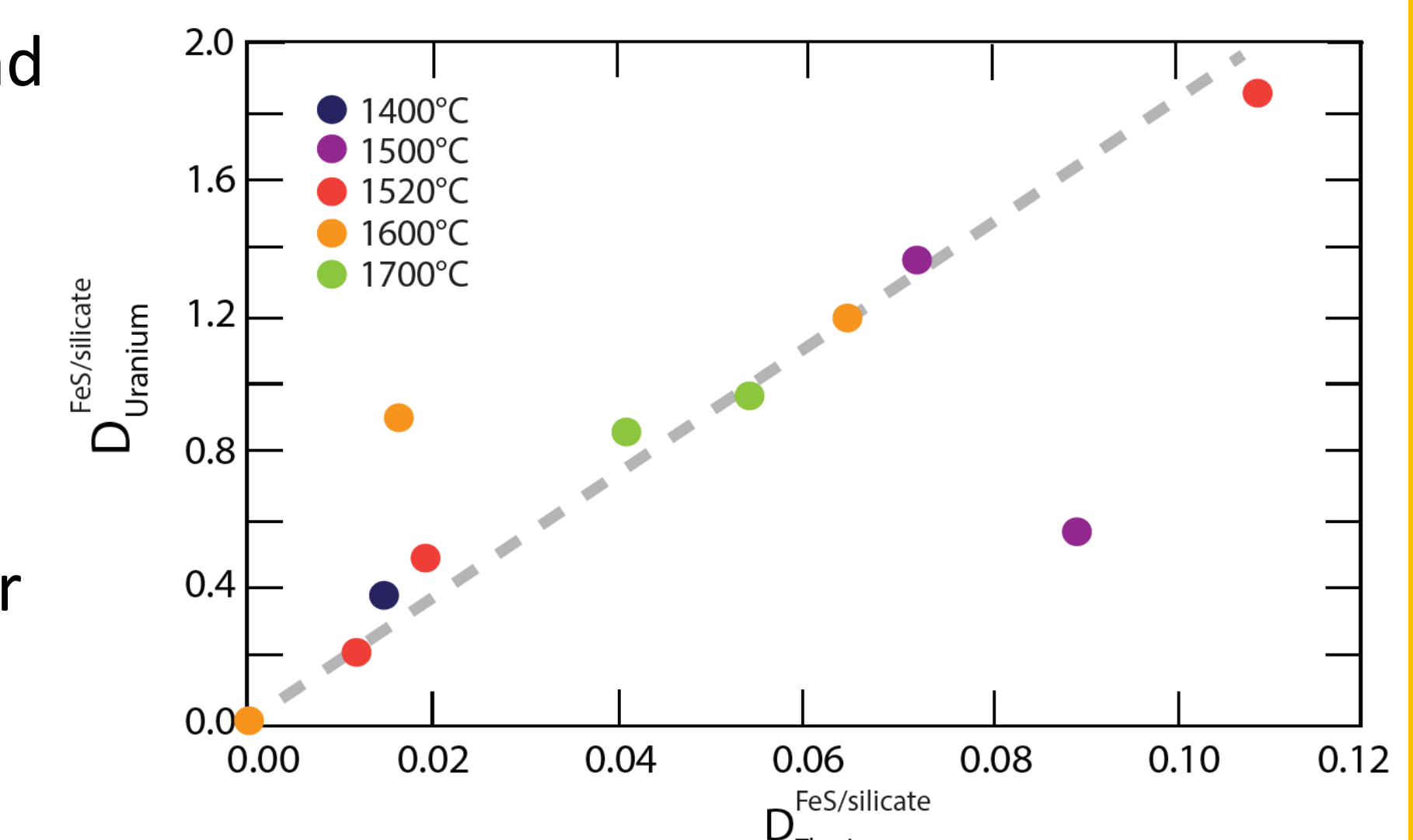
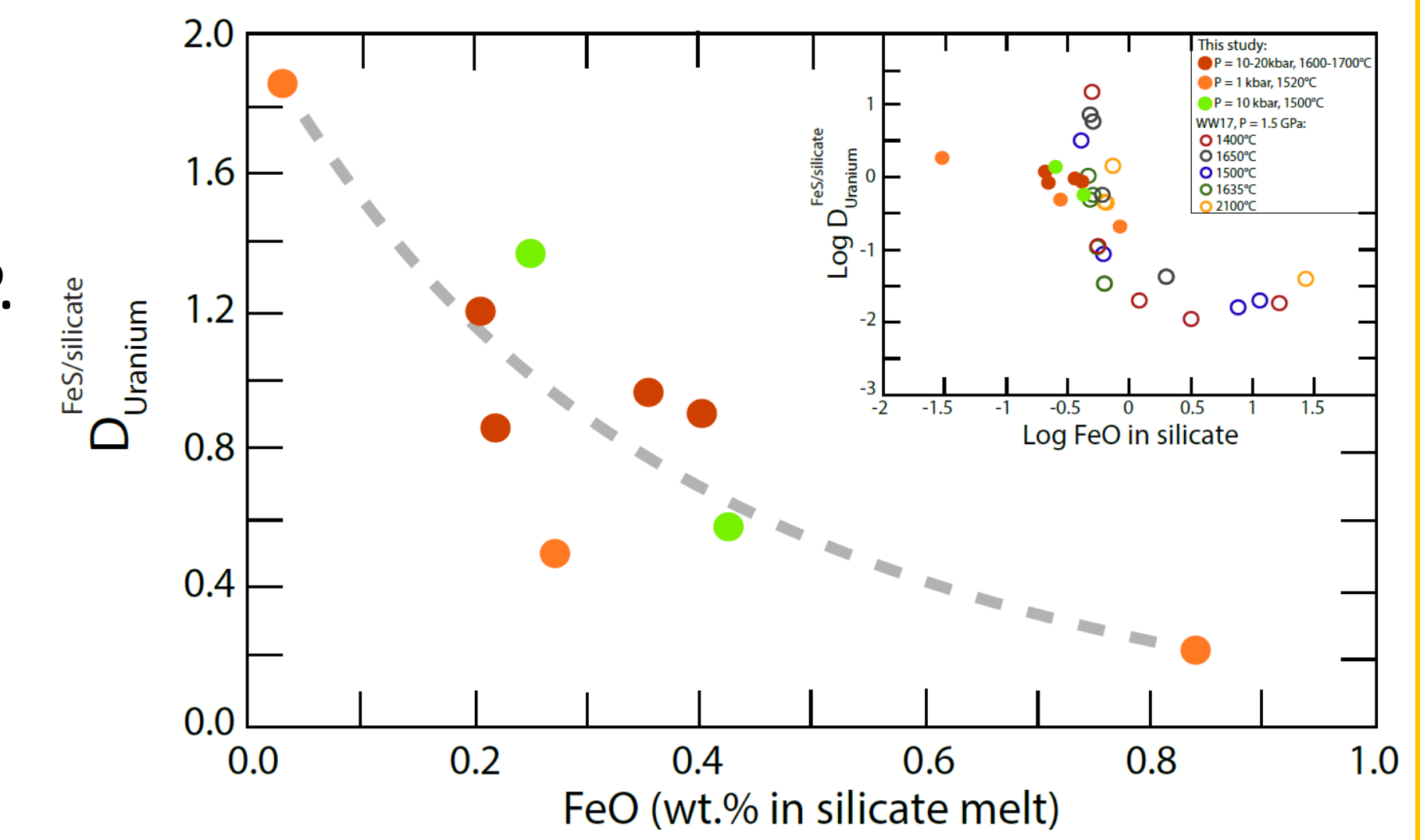
Oxygen fugacity has been estimated using the 2 following methods:

- (1)  $Fe_{Metal} + 1/2 O_2 = FeO_{Silicate}$
- (2) Sulfur Content at Sulfide Saturation (SCSS) [5]

Partition coefficients are expressed as a function of FeO in the silicate melt.  $fO_2$  estimates obtained with (1) are similar to estimates provided by (2)

## Partition coefficients

- $D_{U}^{FeS/sil}$  increases with decreasing  $fO_2$ . Future data over a range of pressure will help constrain the effect of P. No clear effect of T on partitioning is observed
- Contrastingly, literature data suggest exponential partitioning of U at very low  $fO_2$  [6,7]
- The ratio between  $D_{U}^{FeS/sil}$  and  $D_{Th}^{FeS/sil}$  seems constant and independent of  $fO_2$ , T or P
- K partitions less in the sulfides than U for the same  $fO_2$ . New data at extremely low  $fO_2$  are required to better constrain K partitioning
- Current dataset of partitioning data of CaMgS doesn't allow constraining partitioning data. However,  $D_{U}^{sul/sil}$  and  $D_{Th}^{sul/sil}$  are higher than their FeS counterparts



Heat-producing elements partitioning results. Top panel:  $dU^{FeS/sil}$  as a function of FeO content in the silicate melt. Inset: comparison with data from Wohlers & Wood, 2017. Bottom panel:  $dU^{FeS/sil}$  as a function of  $dTh^{FeS/sil}$

- Heat-producing elements do not partition into Fe-Si. There is no clear trend between U and Th partitioning and  $fO_2$
- Fe-P incorporates U and Th ( $D_{U}^{FeP/sil} \sim 4$ ). More experiments are needed to understand the partitioning behavior in the Fe-P phases

## Take-home message

- U, Th and K display increasing chalcophile behavior at low  $fO_2$
- FeS does not fractionate U and Th
- An FeS layer should incorporate large amounts of U, Th and K
- An Fe-Si core should be mainly depleted in heat-producing elements
- Accessory phases such as Fe-P could be a secondary reservoir for U, Th, K

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