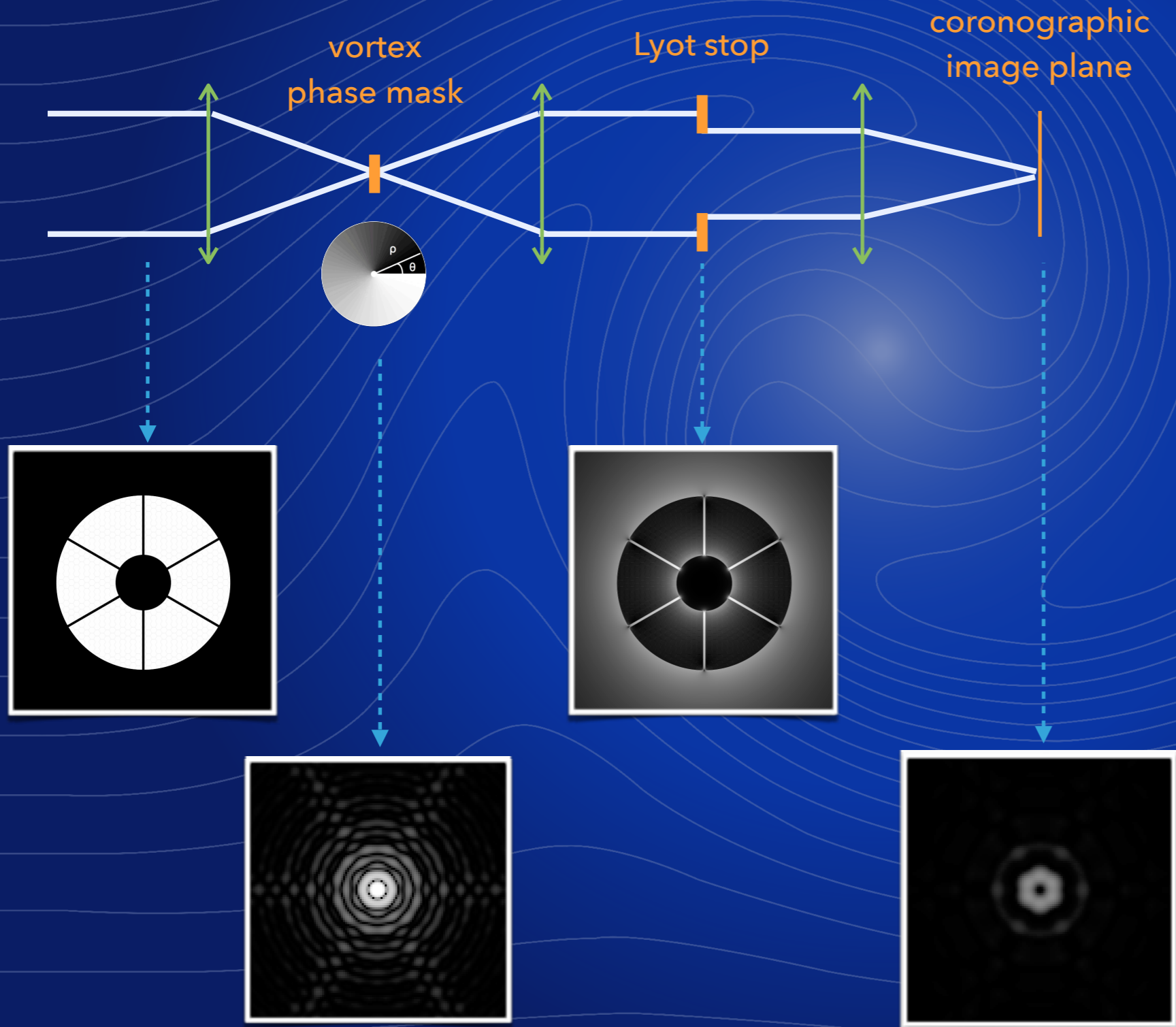


VORTEX

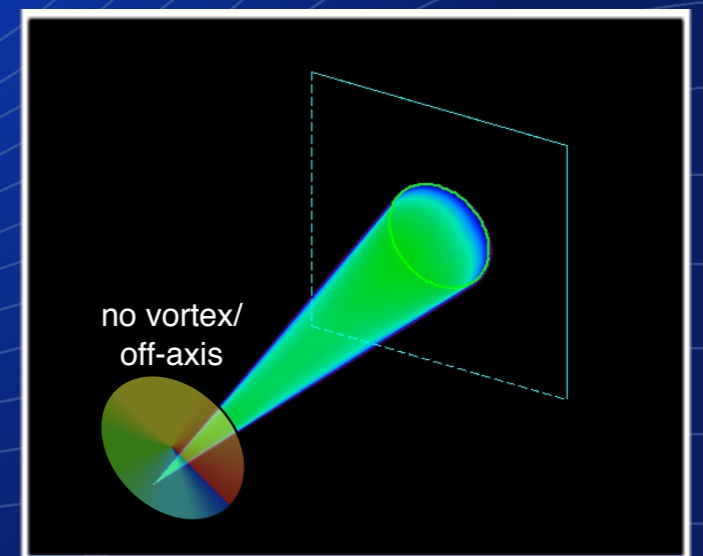
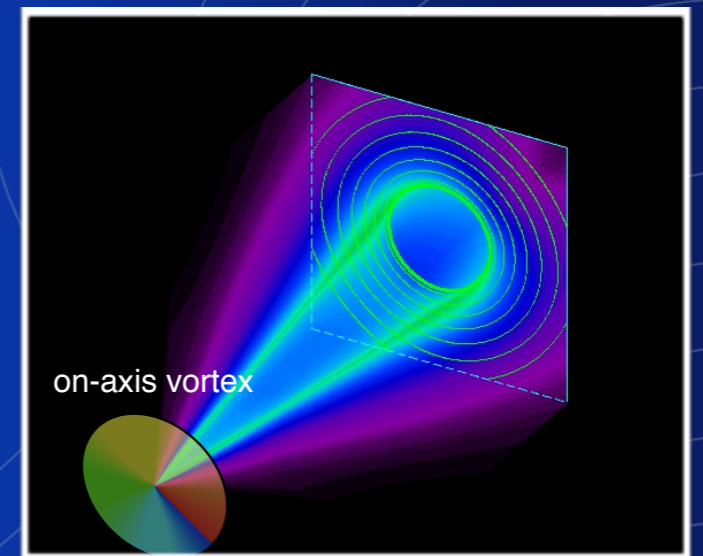
OLIVIER ABSIL

AROUND THE WORLD WITH THE INFRARED VORTEX CORONAGRAPH

THE VORTEX CORONAGRAPH IN A NUTSHELL

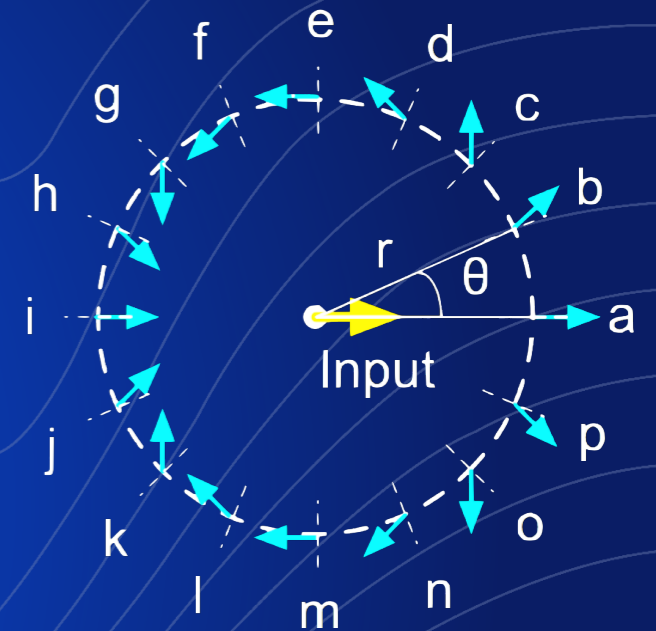
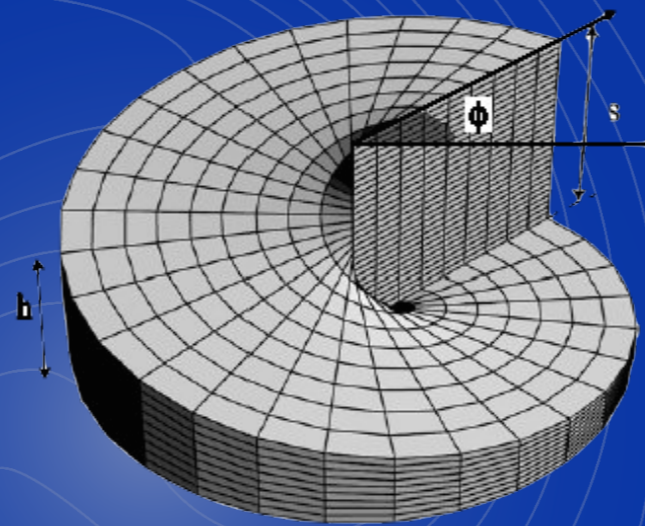


perfect on-axis cancellation
for a circular aperture



IMPLEMENTATIONS OF THE VORTEX PHASE MASK

- ▶ scalar vortex
 - * helical piece of glass
- ▶ vector vortex = spatially variant HWP
 - * liquid crystal polymers
 - * subwavelength gratings
 - * photonic crystals



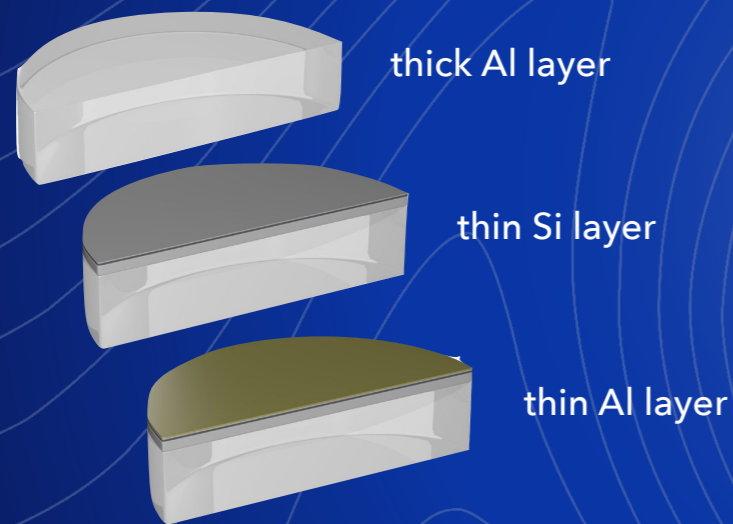
Annular Groove Phase Mask



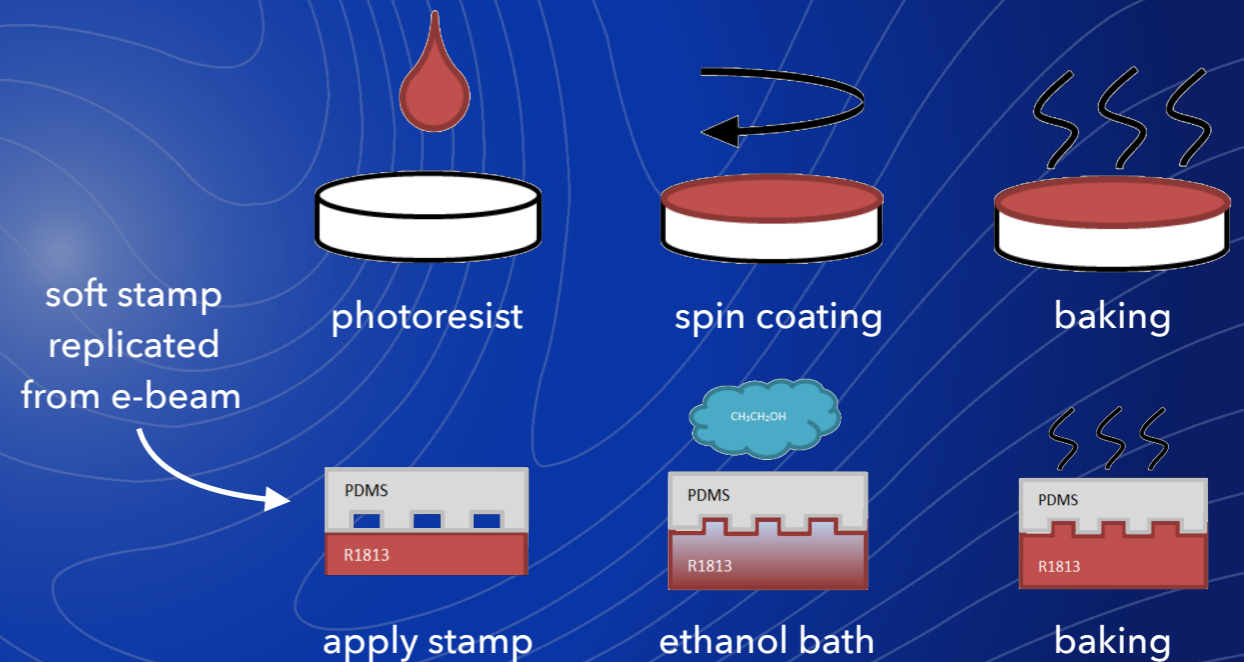
MANUFACTURING DIAMOND AGPM @ UPPSALA

Vargas Catalan et al. (2016)

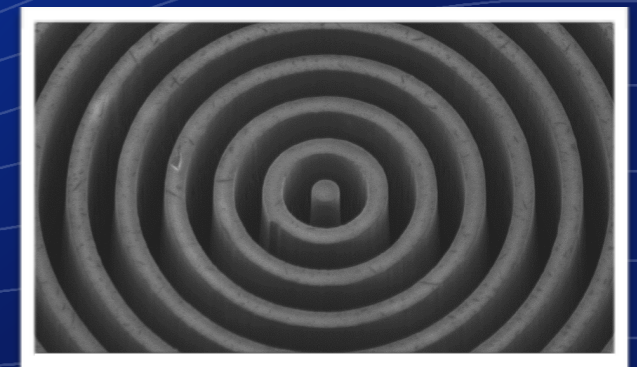
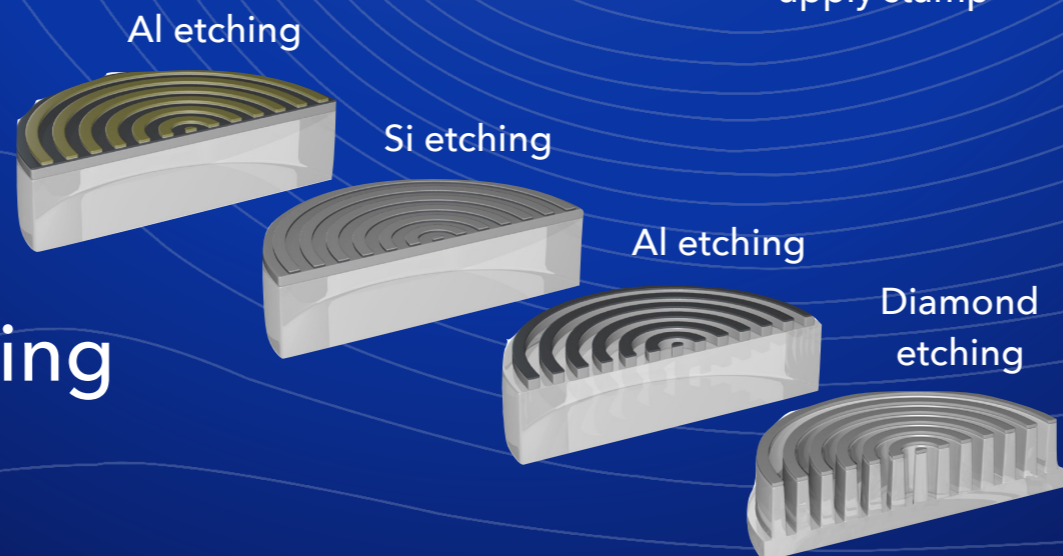
1. diamond coated with Al and Si layers (sputtering)



2. e-beam pattern transferred with solvent-assisted moulding

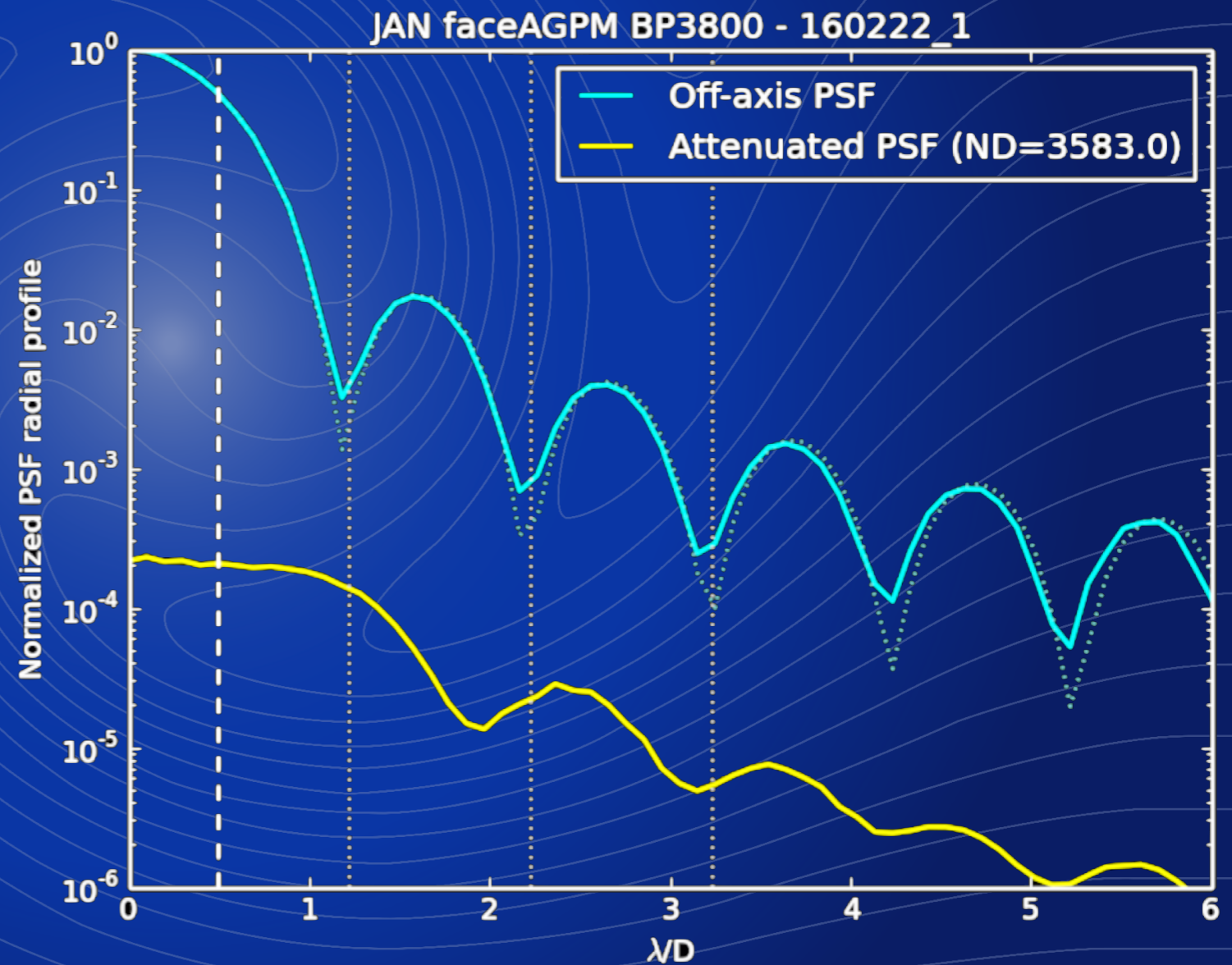


3. reactive ion etching



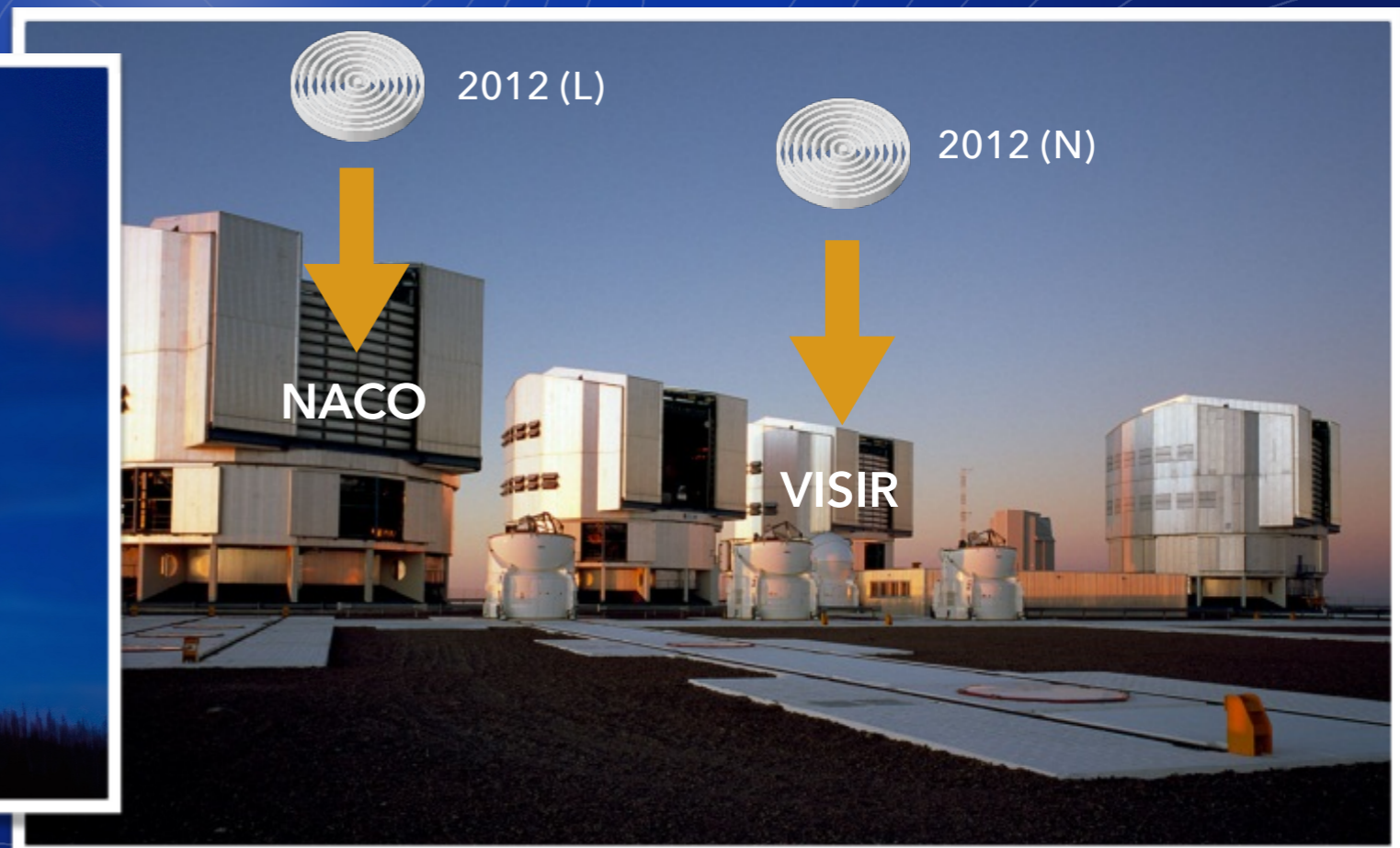
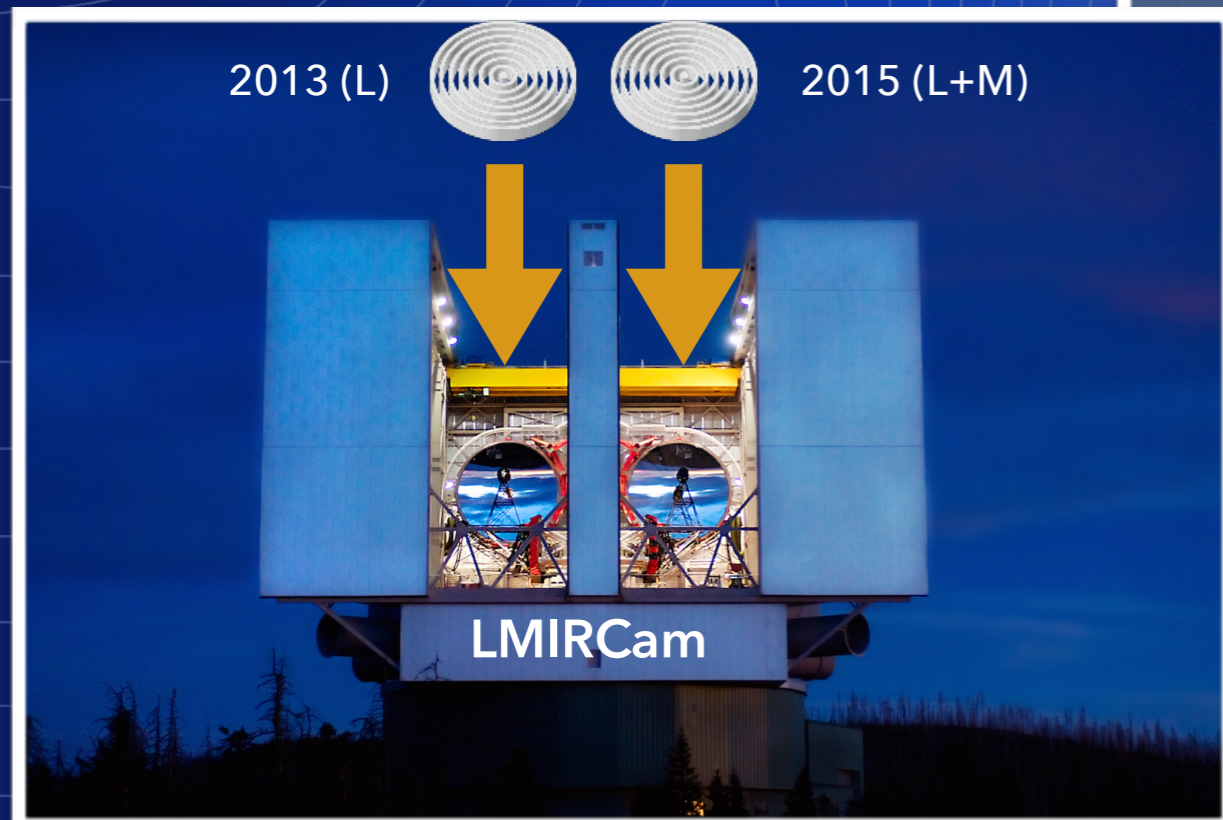
BEST PERFORMANCE IN THE LAB - 2018 UPDATE

- ▶ dedicated test bench (VODCA) now available at ULiège
- ▶ 10+ science-grade L-band AGPMs etched & tested
- ▶ broadband rejection > 1000 : 1



INSTALLATION AND COMMISSIONING

- ▶ piggyback on existing coronagraphic IR cameras
- ▶ short commissioning runs (1-2 nights)

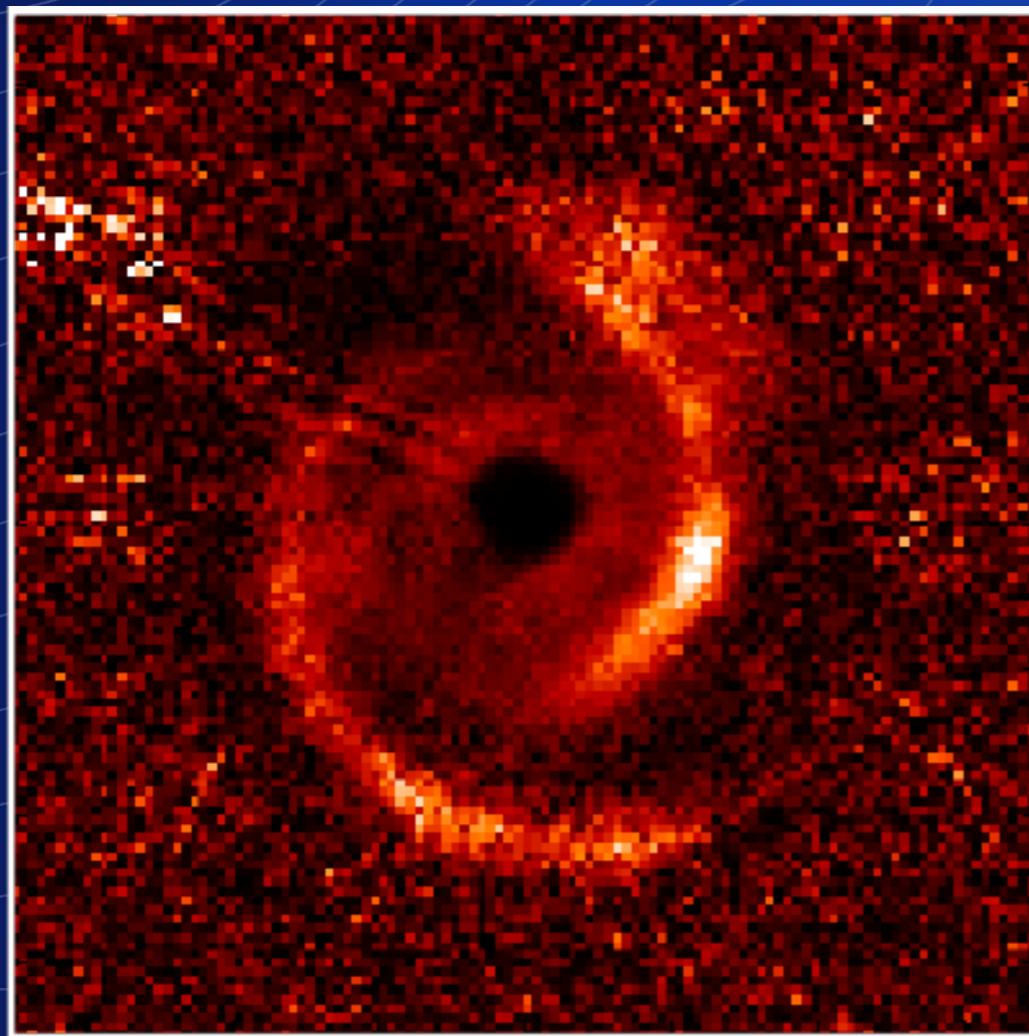




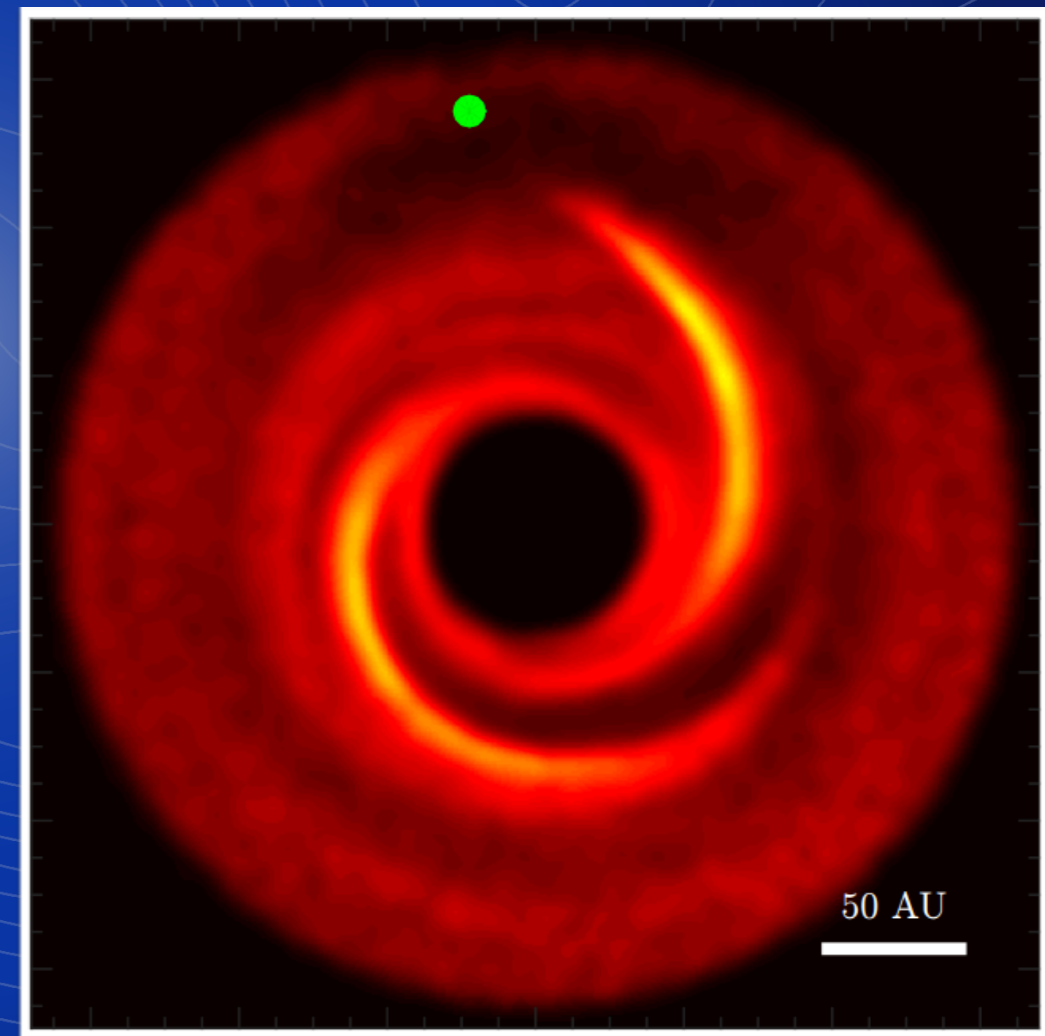
PERFORMANCE AND MAIN RESULTS SO FAR

NIRC2 & NACO TRANSITION DISK SURVEY

SPHERE/IRDIS Y band polarimetry (Benisty et al. 2015)

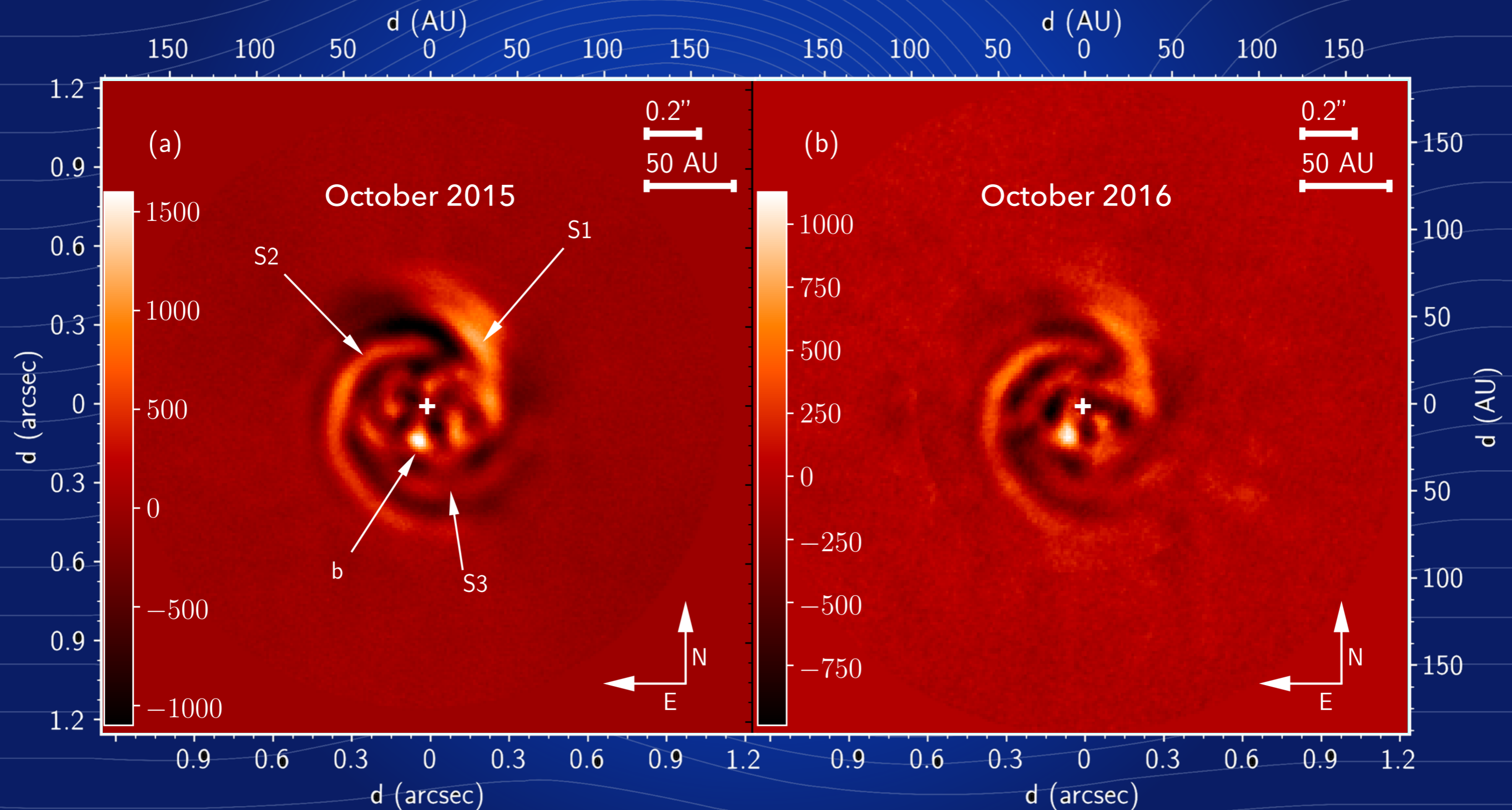


Protoplanet prediction (Dong et al. 2015)



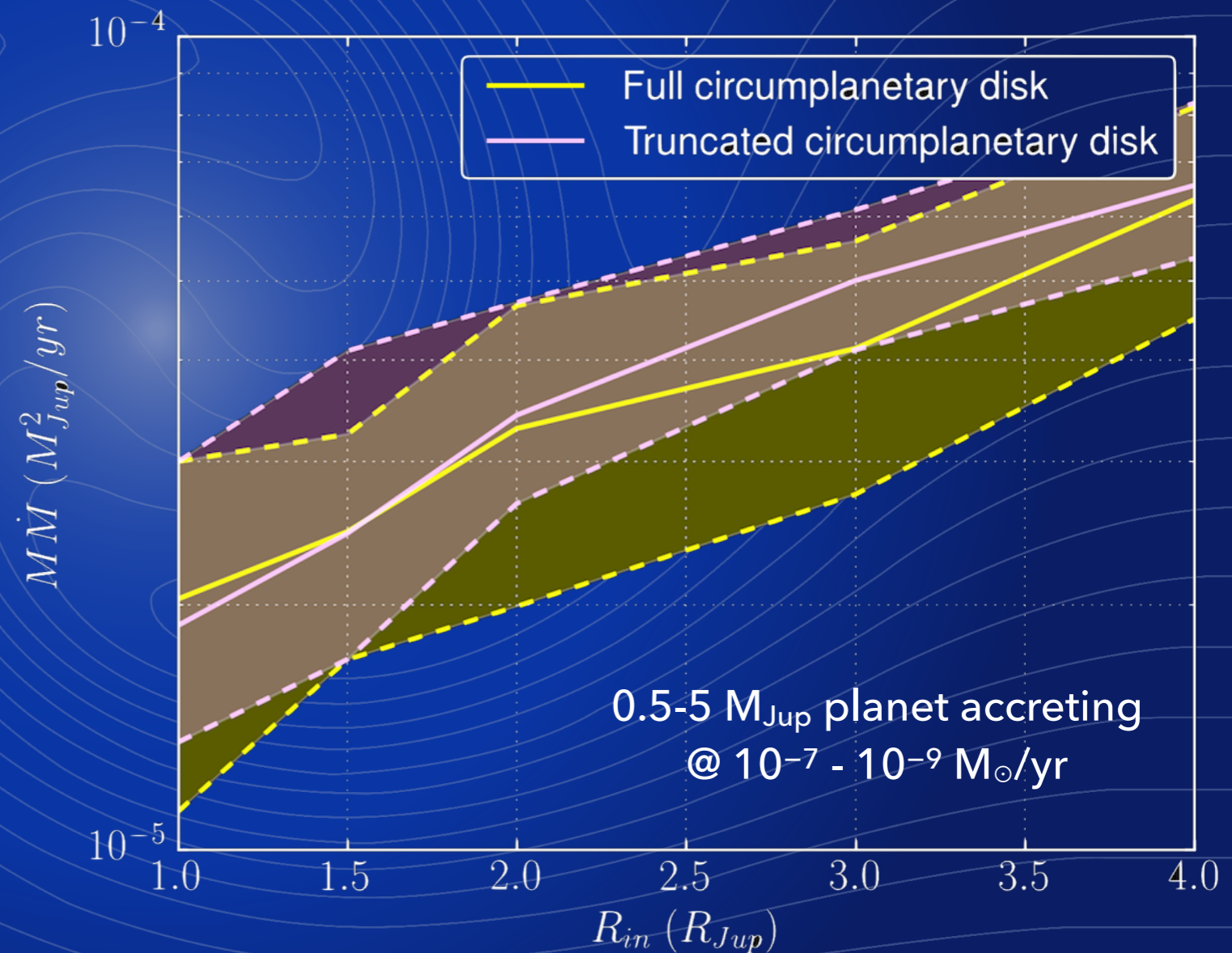
goal: search for protoplanets at the origin of disk structures

THE KECK/NIRC2 + VORTEX VIEW OF MWC758

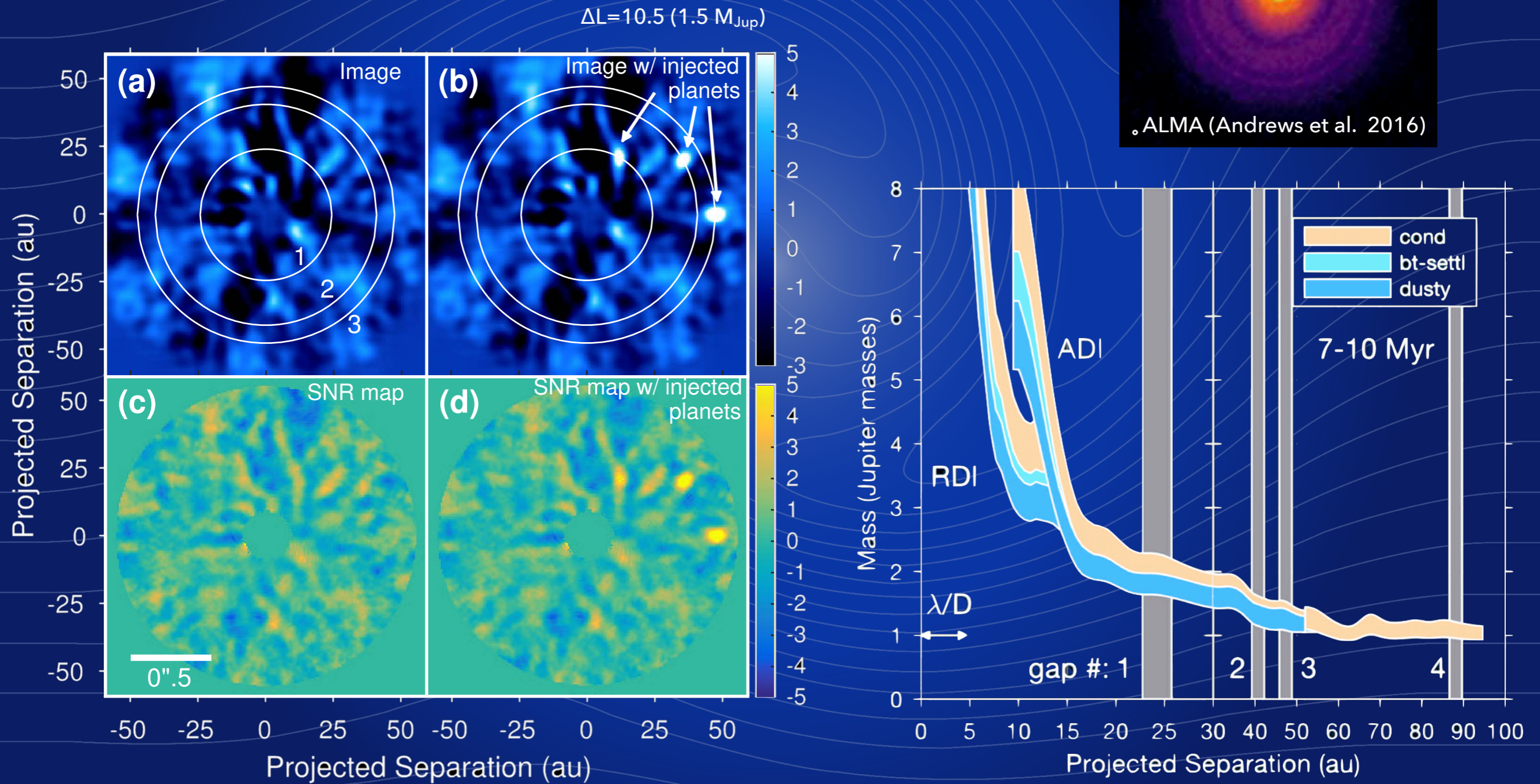


MWC758B: YET ANOTHER PROTOPLANET CANDIDATE?

- ▶ 0.1'' separation (20 au),
 $\Delta L = 7$
- ▶ movement consistent with Keplerian orbit
- ▶ if photospheric emission, would be $\sim 50 M_{\text{Jup}}$
 - * not consistent with structure of inner disk
- ▶ accreting protoplanet or disk feature?

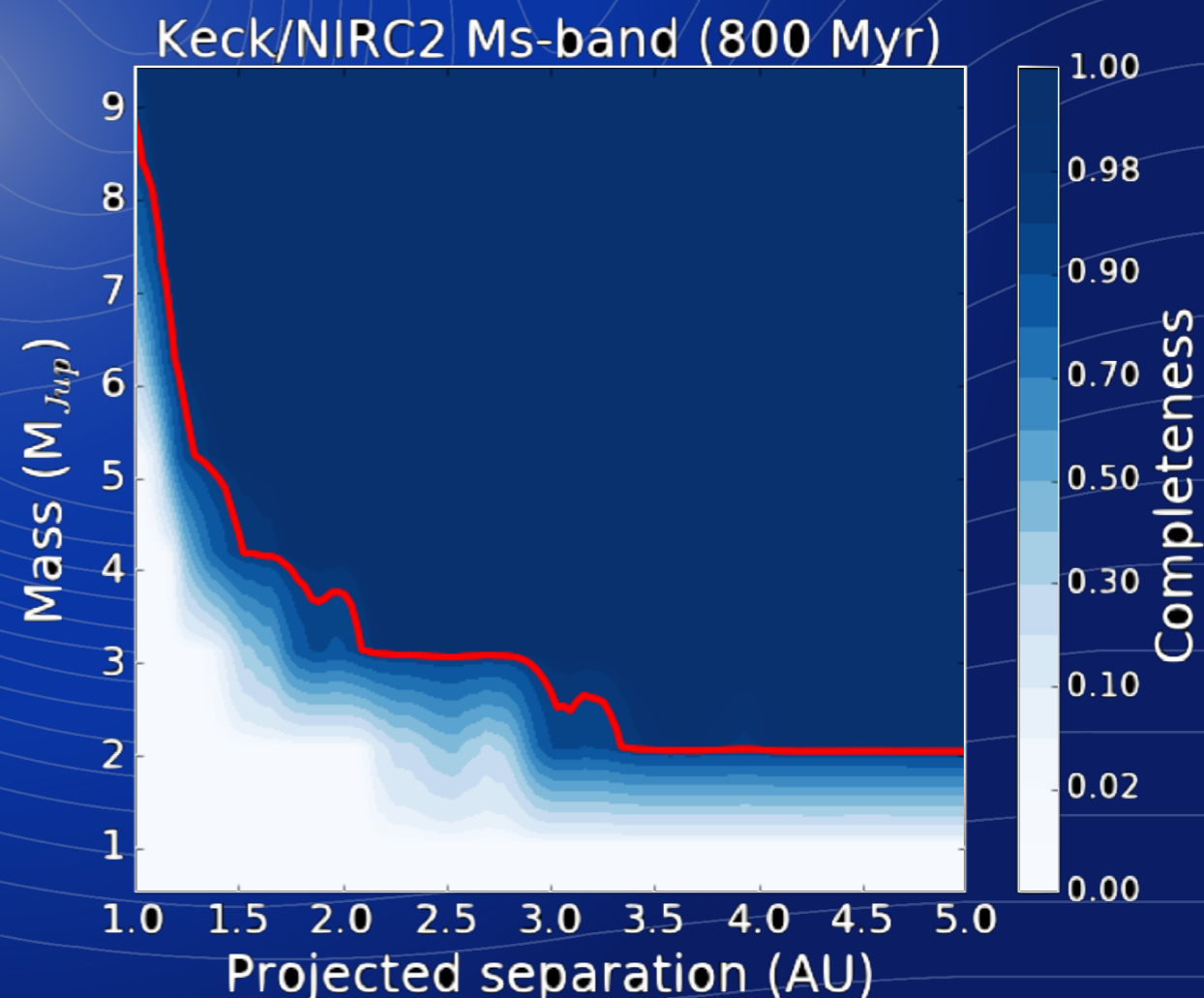
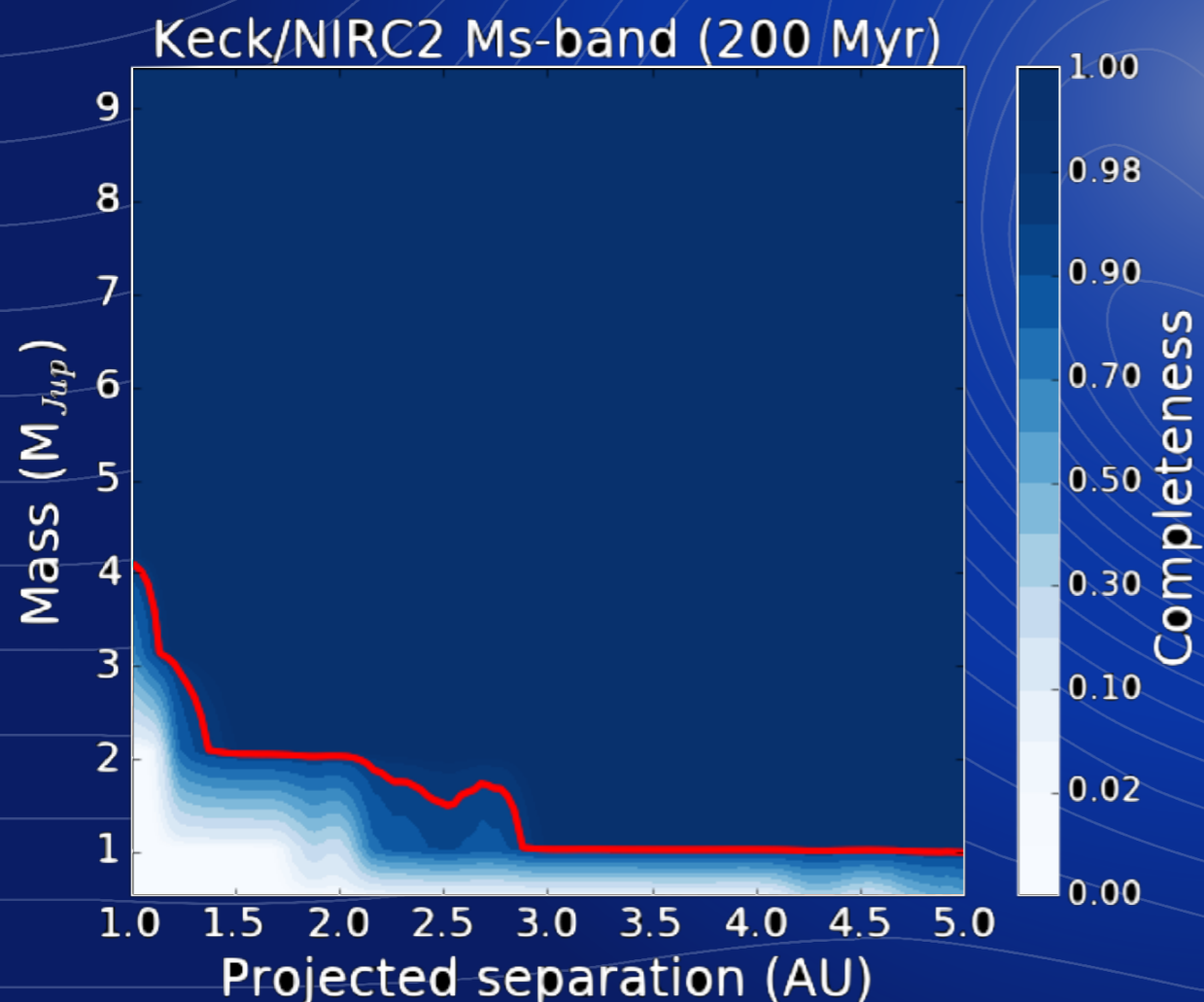
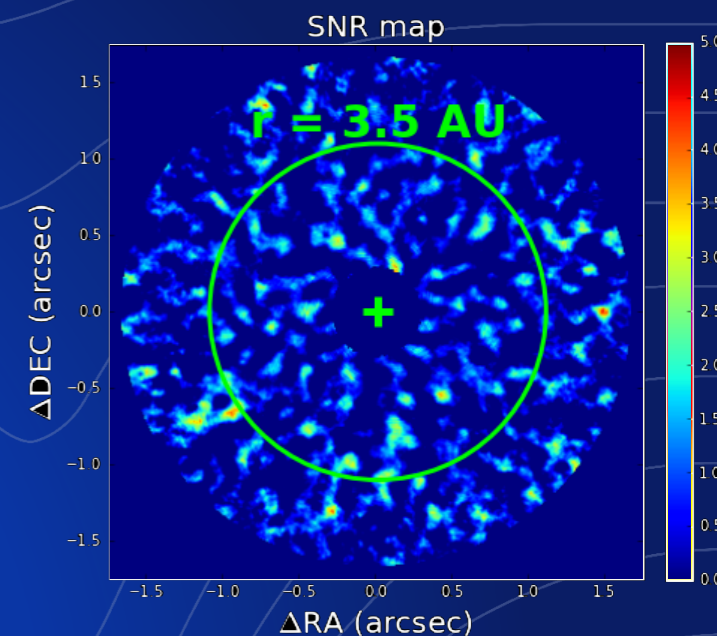


KECK CORONAGRAPHIC DEEP FIELD: TW HYA



KECK CORONAGRAPHIC DEEP FIELD: EPS ERI

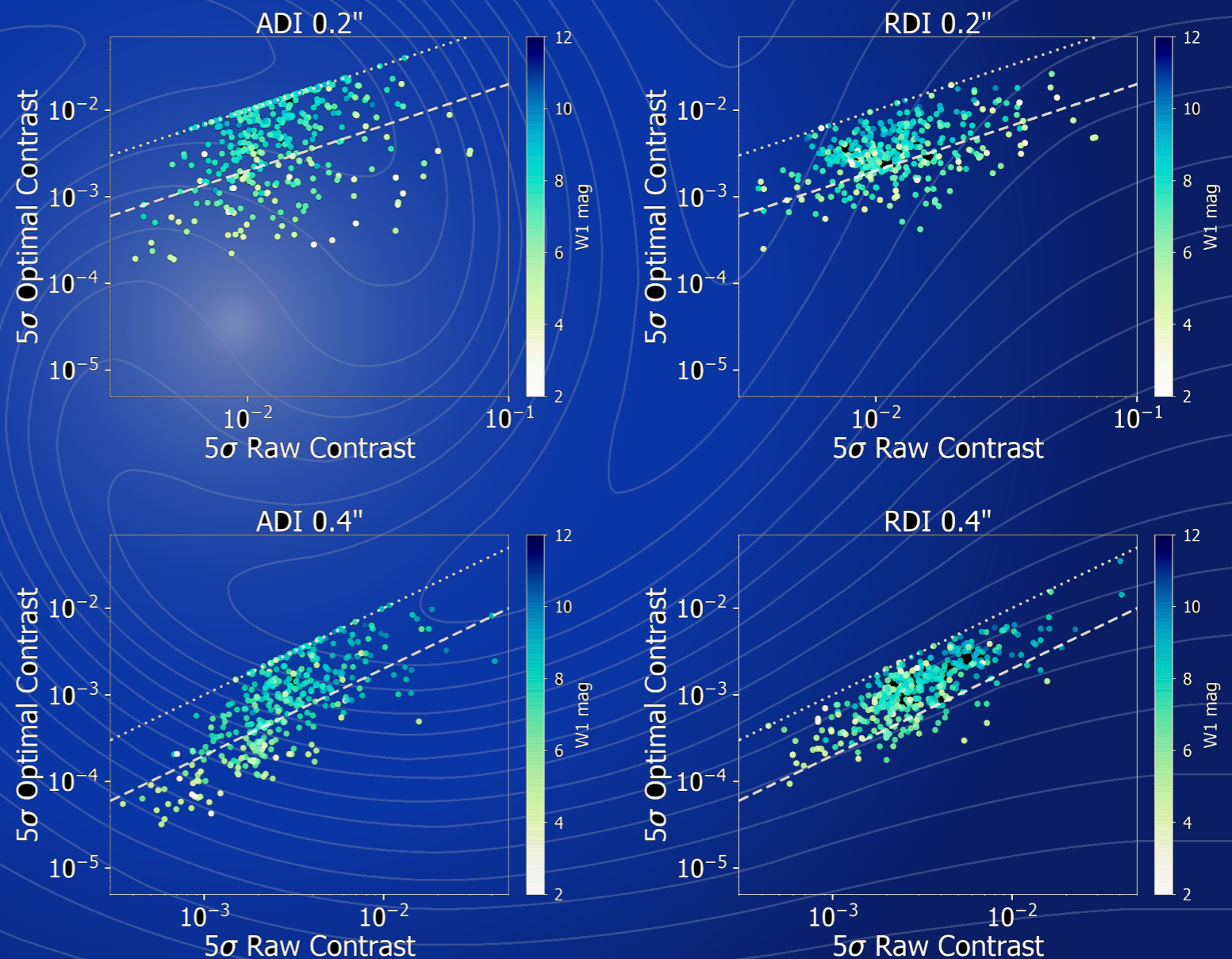
- ▶ deepest detection limits around an adolescent Sun-like star



ON SKY PERFORMANCE: THREE YEARS OF VORTEX @ NIRC2

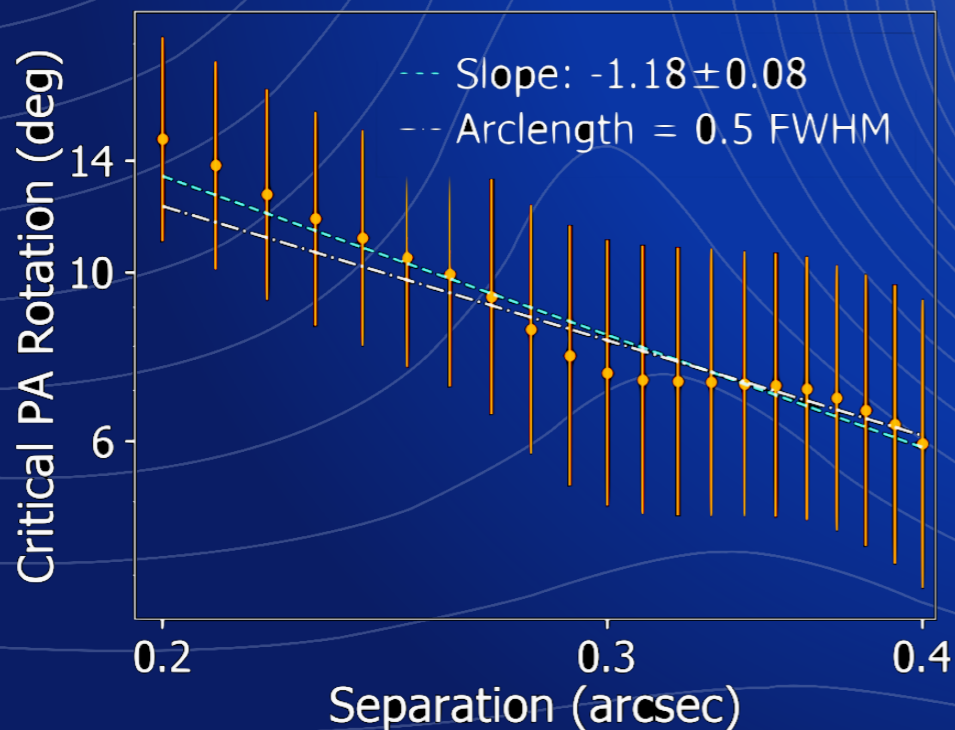
- ▶ 359 vortex observations in 2016-2017, using QACITS focal-plane pointing control
- ▶ automatic pipeline provides raw and post-processed contrast curves for ADI and RDI

gain in contrast from post-processing

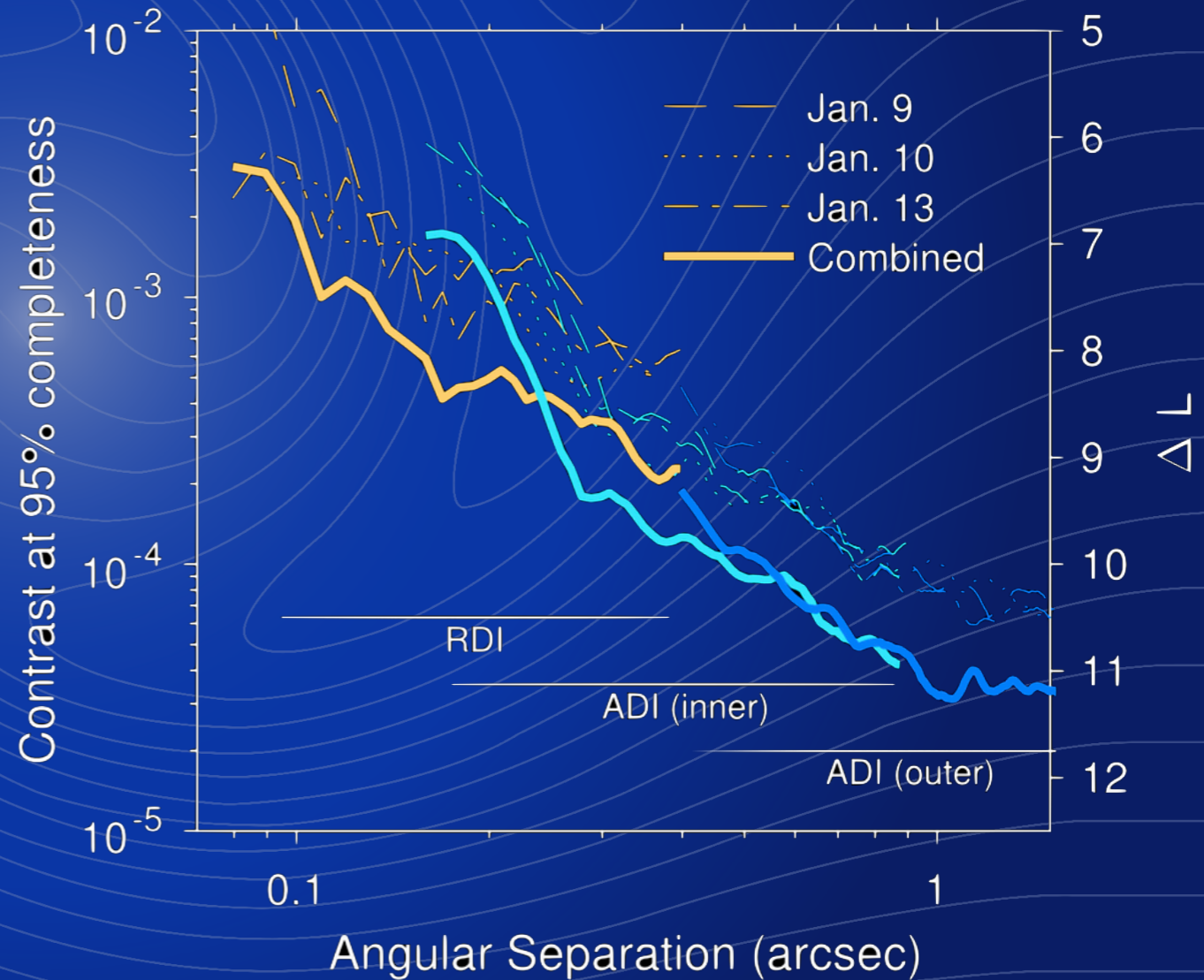


ON SKY PERFORMANCE: ADI VS RDI

- ▶ RDI better at small separations
- ▶ critical separation depends on amount of field rotation



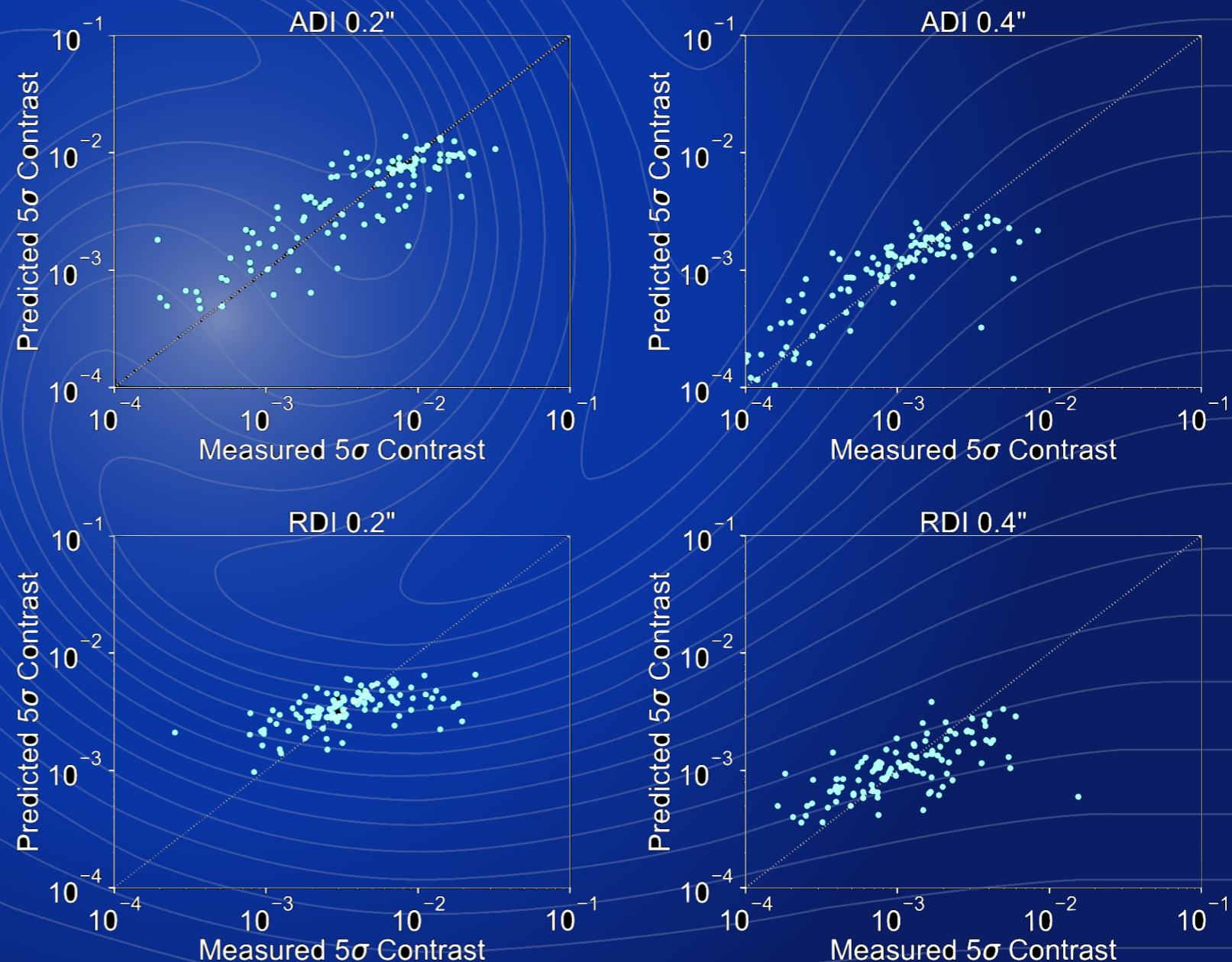
Ruane et al. 2017



CONTRAST PREDICTIONS BASED ON RANDOM FORESTS

- ▶ random forests used to identify main explanatory variables and make predictions
- ▶ ADI performance mostly explained by (i) field rotation, (ii) integration time, (iii) magnitude
- ▶ for RDI performance, FWHM of PSF becomes as important

predicted vs measured contrast

RMSE ~ 0.3 dex

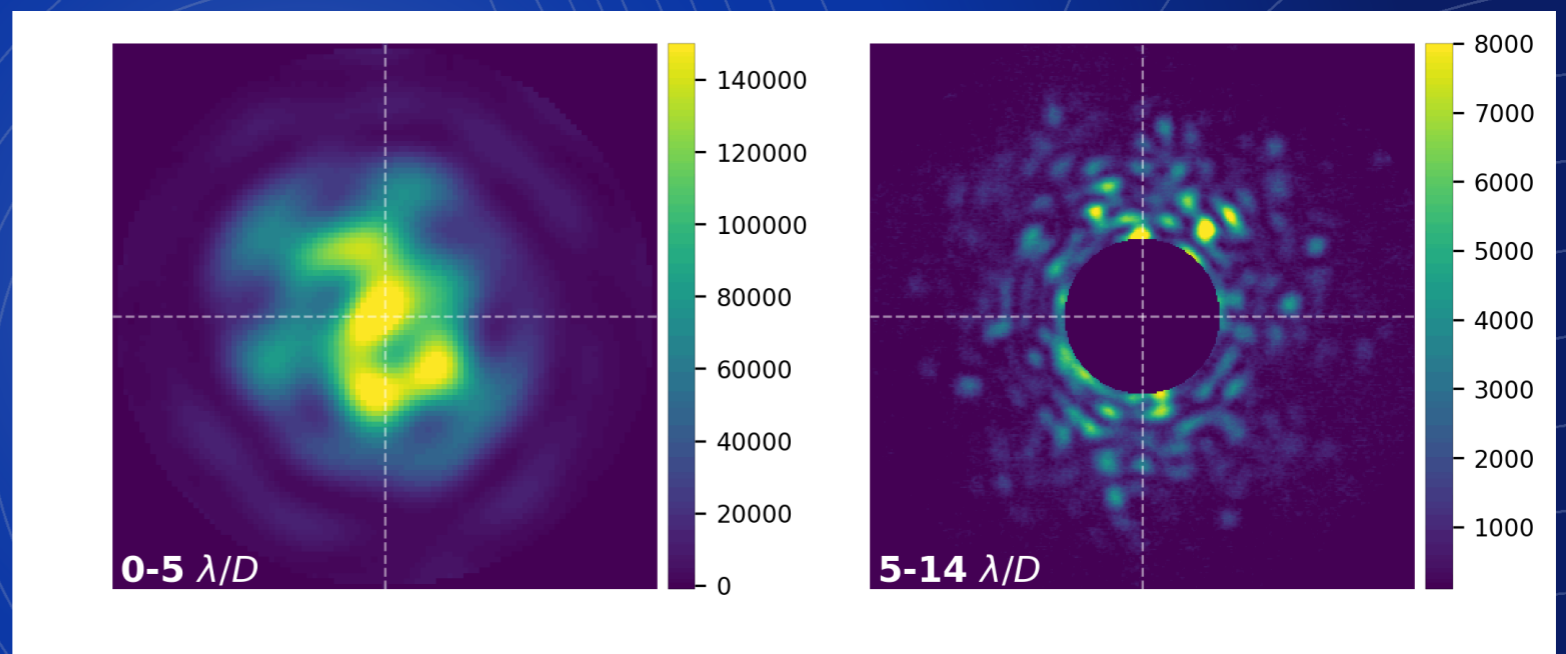
HOW TO BETTER EXPLOIT THE DATA?

- ▶ interesting science at $1-3 \lambda/D$

- * strongly affected by residual speckles
- * non-Gaussian noise
→ more false positives
- * hard to validate candidates

- ▶ ADI-based techniques produce SNR, but do not inform on nature of the source
- ▶ machine learning can help

NIRC2+vortex image sequence

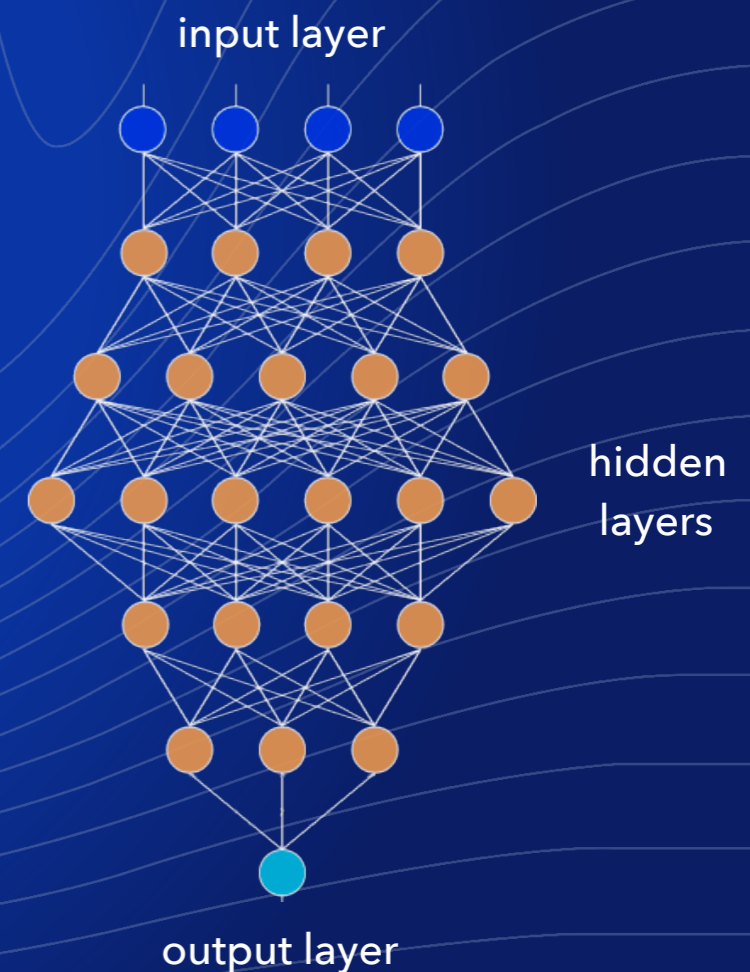


SUPERVISED LEARNING

- ▶ goal: learn function f mapping input samples \mathcal{X} to labels \mathcal{Y} given a labeled dataset $(x_i, y_i)_{i=1, \dots, n}$:

$$\min_{f \in \mathcal{F}} \frac{1}{n} \sum_{i=1}^n \mathcal{L}(y_i, f(x_i)) + \lambda \Omega(f)$$

- ▶ mapping function f based on (deep) neural network
 - * layers of neurons whose parameters can be tuned to approximate a complex function
 - * DNN can be trained with labeled datasets
- ▶ problem: need labels & large training sample!



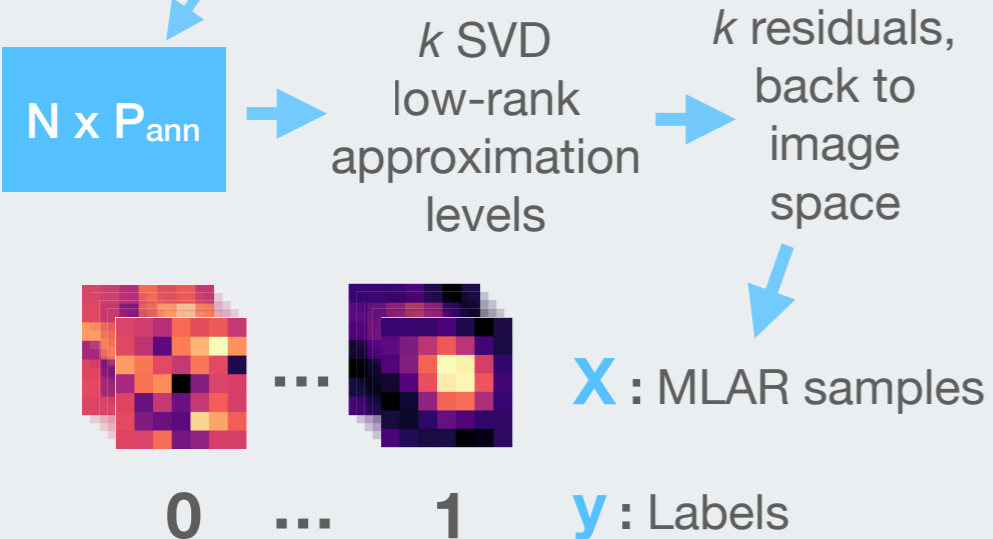
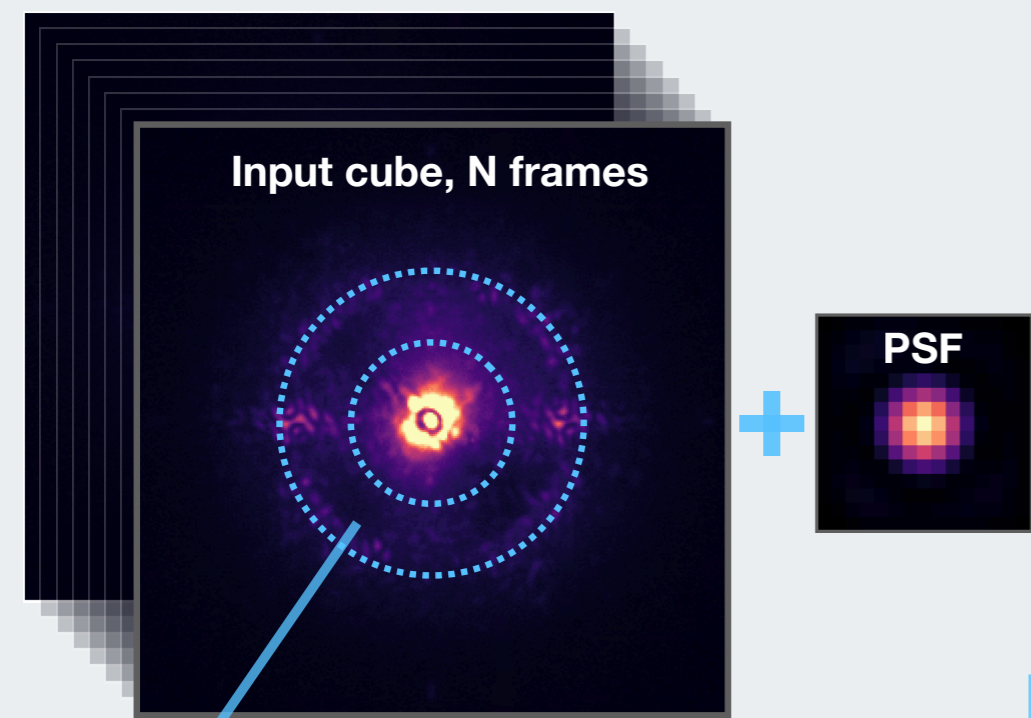
SUPERVISED DETECTION OF EXOPLANETS

Gomez Gonzalez et al. 2018

1. generation of labeled data

2. training the DNN

3. prediction



X and y to train/test/validation sets

Convolutional LSTM layer
kernel=(3x3), filters=40

3d Max pooling
size=(2x2x2)

Convolutional LSTM layer
kernel=(2x2), filters=80

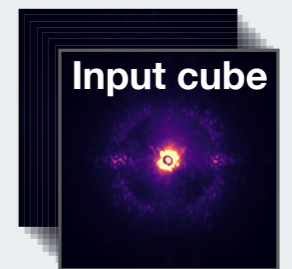
3d Max pooling
size=(2x2x2)

Dense layer
units=128

ReLU activation + dropout

Output dense layer
units=1

Sigmoid activation



MLAR patches

Trained classifier

Probability of positive class

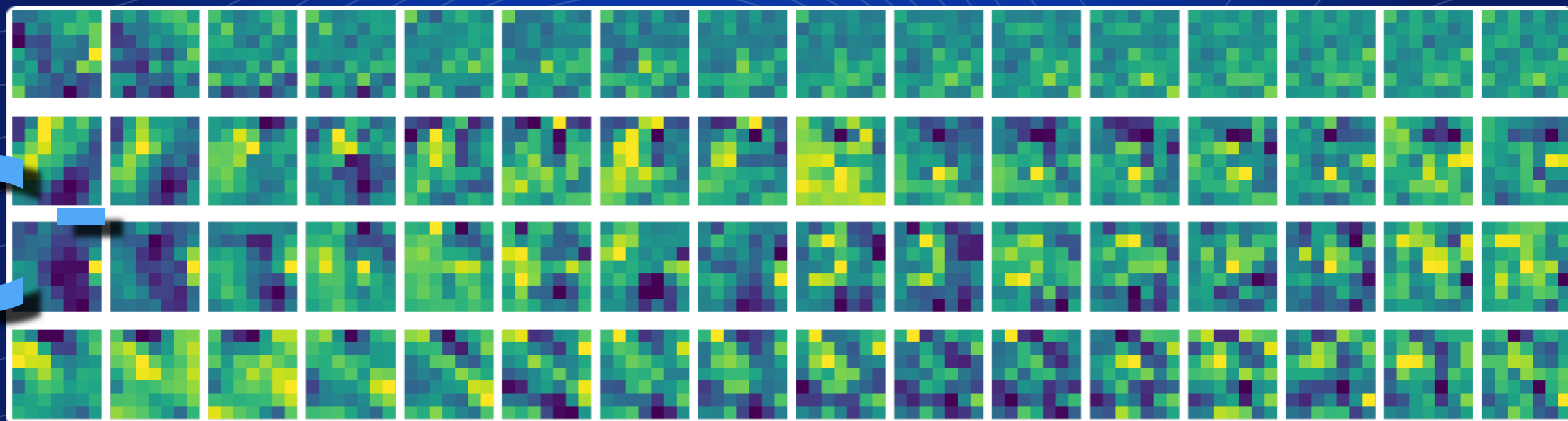
Binary map

probability threshold = 0.9

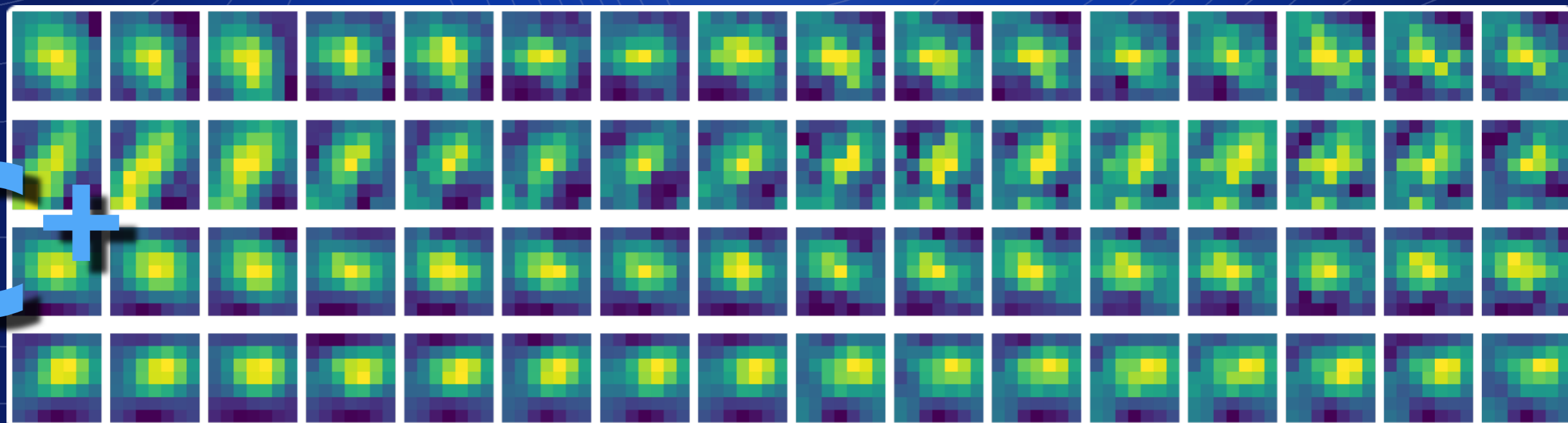
LABELED DATASET

Labels: $y \in \{c^-, c^+\}$

C⁻

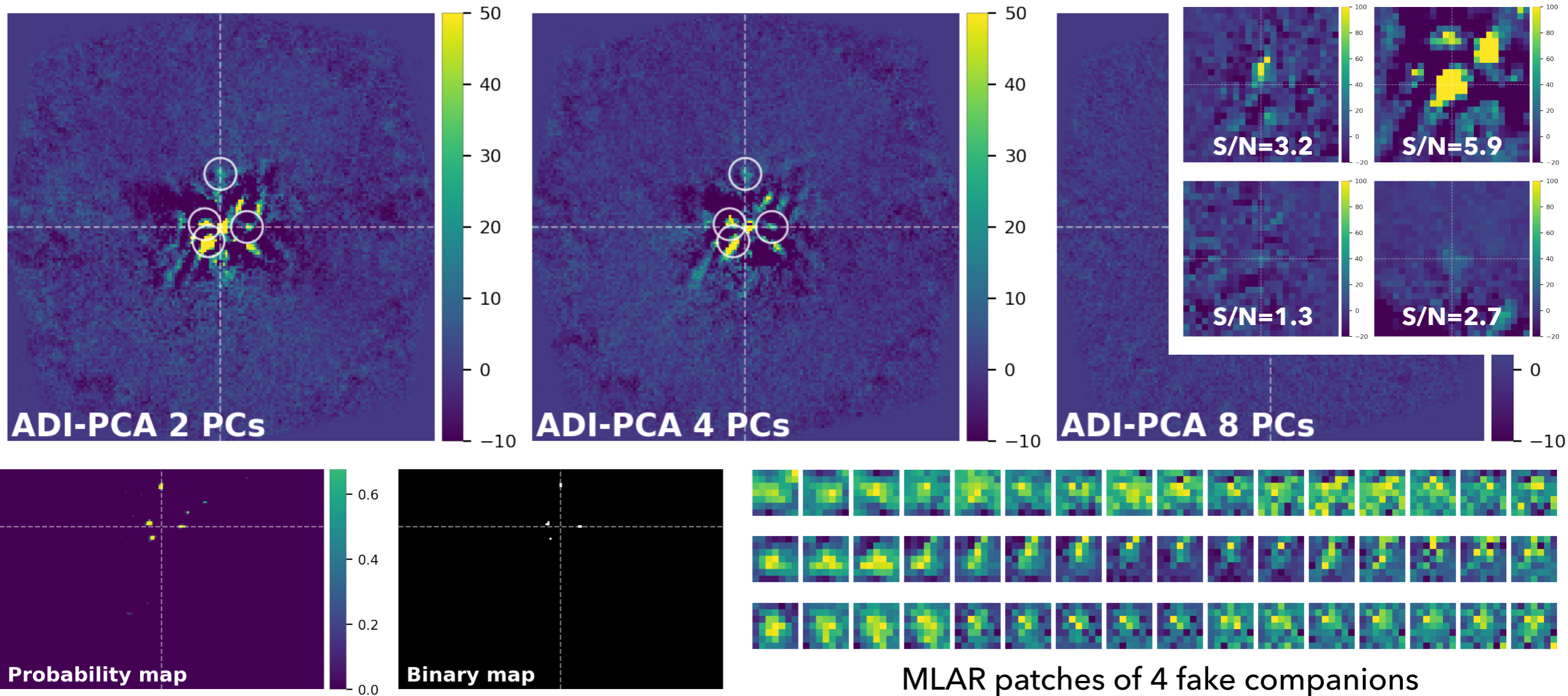


C⁺



TEST WITH INJECTED COMPANIONS (SPHERE/IRDIS DATA)

4 fake companions injected in data set ... can you spot them?

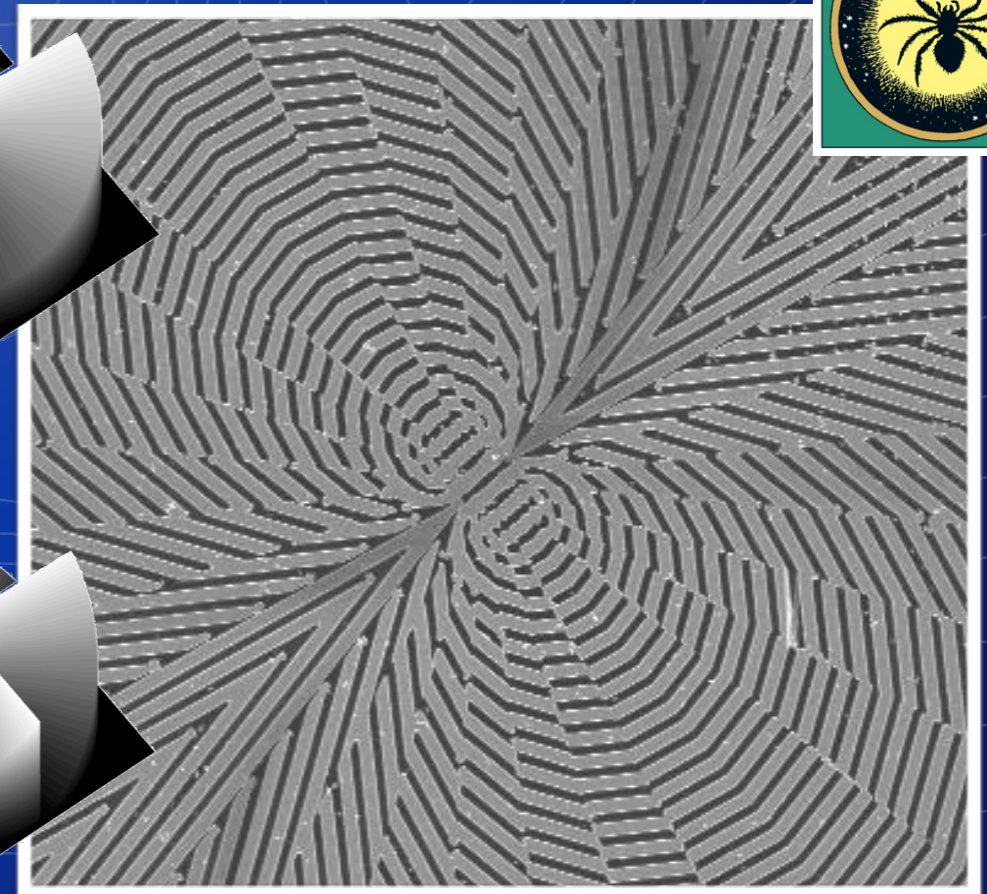
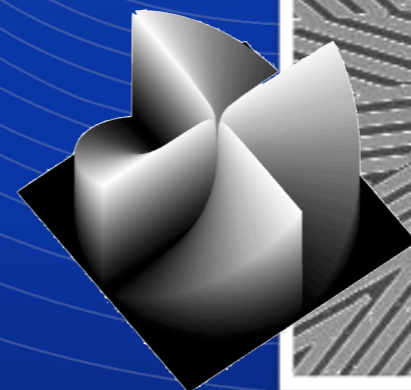




NEW RESEARCH PATHS AND FUTURE PROJECTS

EXTENDING THE AGPM CONCEPT

- ▶ AGPM first developed for thermal infrared (L, M, N bands)
 - * excellent performance on ~30% bandwidth
- ▶ shorter wavelengths
 - * science-grand K-band AGPM now available
 - * H-band AGPM development started
- ▶ higher topological charges
 - * less sensitive to tip-tilt, at the expense of larger IWA



charge-4 vortex, work in progress

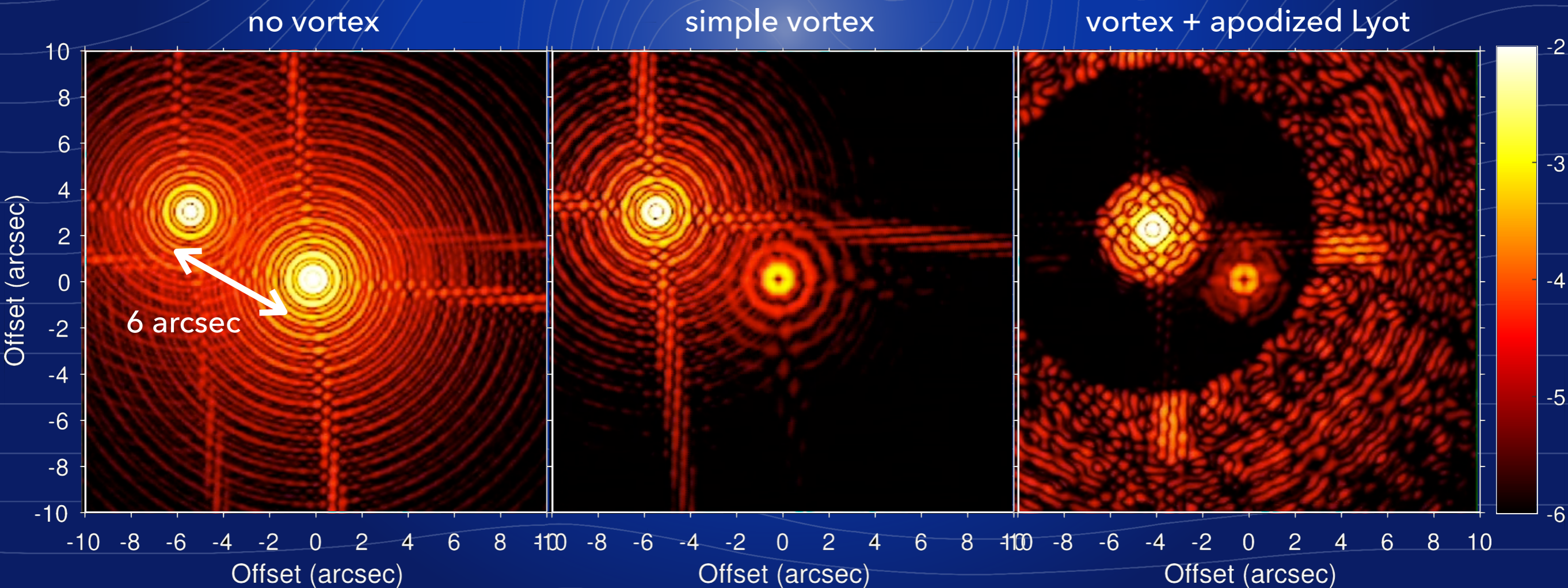
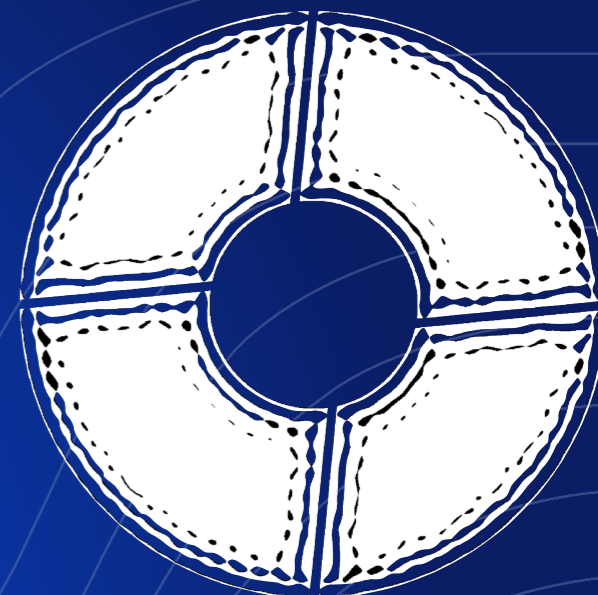
NEAR – NEW EARTH IN THE ALPHA CENTAURI REGION

10702-12 KÄUFL
(Sunday talk)

- ▶ ESO project funded by Breakthrough Initiatives
 - * what? search for rocky planets around α Cen A&B
 - * how? refurbish VISIR and put it behind UT4+AOF
 - * when? 100h observing campaign in mid-2019
- ▶ vortex team contribution
 - * provide optimized AGPM for 10-12.5 μ m filter
 - * design optimized Lyot stop
 - * develop closed-loop focal-plane pointing control (QACITS)

NOTIONAL IMAGES OF ALPHA CENTAURI SYSTEM

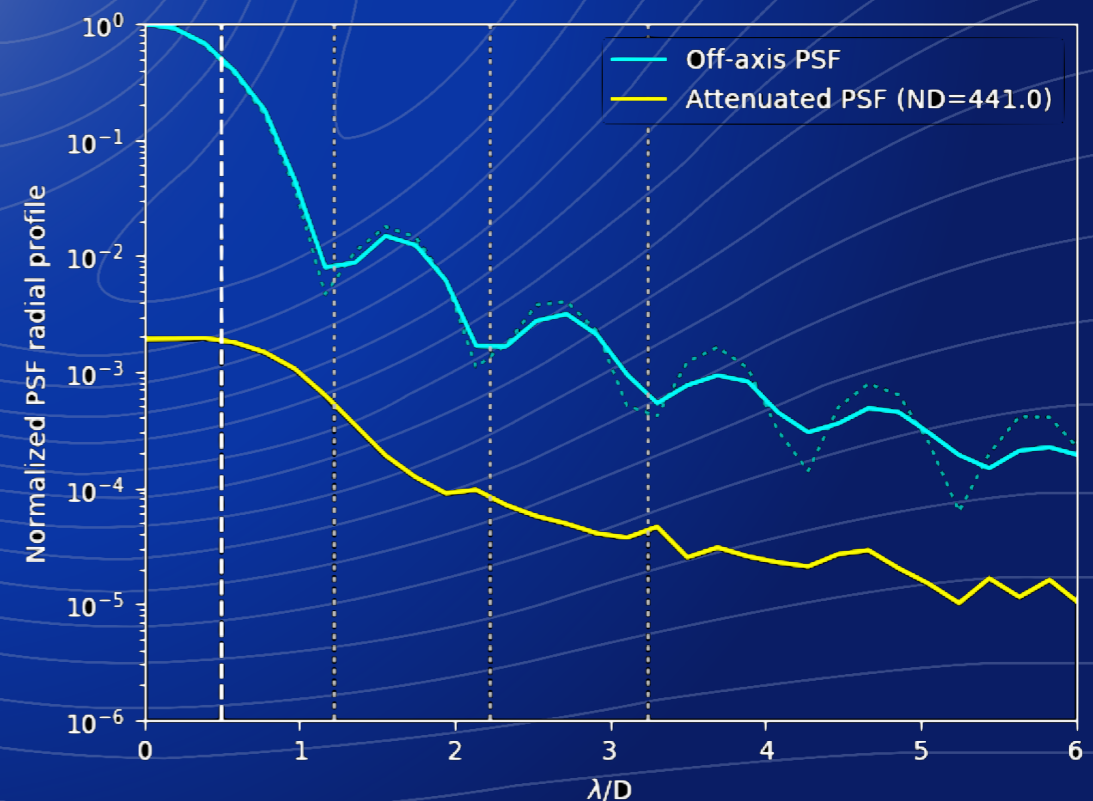
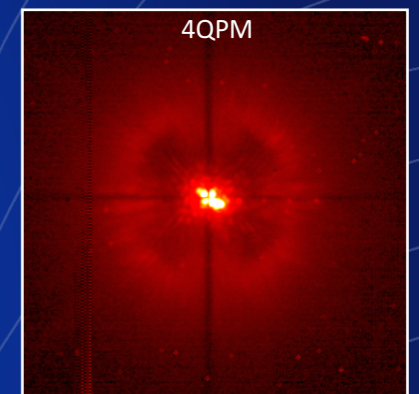
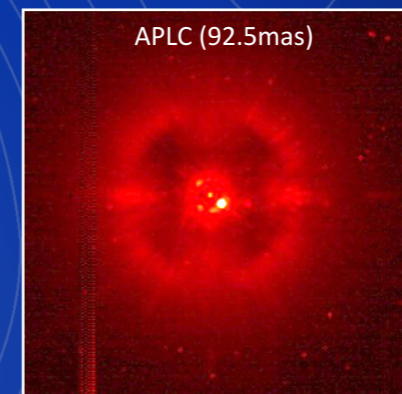
- ▶ habitable zone at 0.8" - 1.1" (A) or 0.5" - 0.65" (B)
- ▶ contrast $\sim 10^{-6}$ for $2 R_{\oplus}$ planet
- ▶ apodized Lyot stop to carve dark hole around secondary



A VORTEX UPGRADE FOR SPHERE?

- ▶ goal: open the $1-3 \lambda/D$ parameter space
 - * increase number of detections
 - * access a few RV planets
- ▶ need to identify main limitations to 4QPM performance
 - * component degradation?
 - * effect of dead actuators?
 - * low-order WF aberrations?
- ▶ K-band AGPM ready to go!

10702-145 HUBY
(Monday poster)



NEXT STEPS: VLT/ERIS AND ELT/METIS

10702-151 KENWORTHY
(Monday poster)

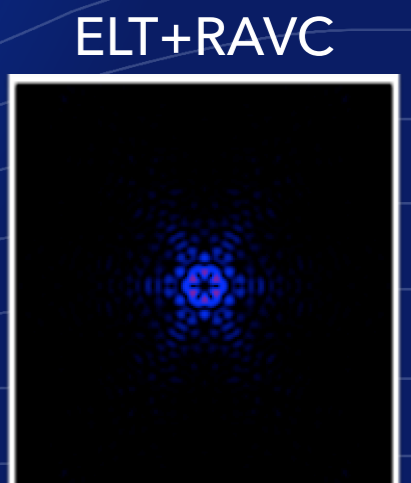
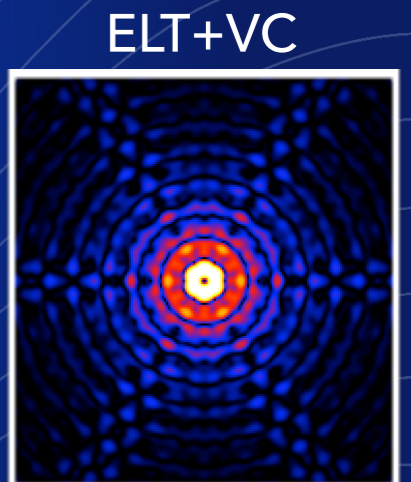
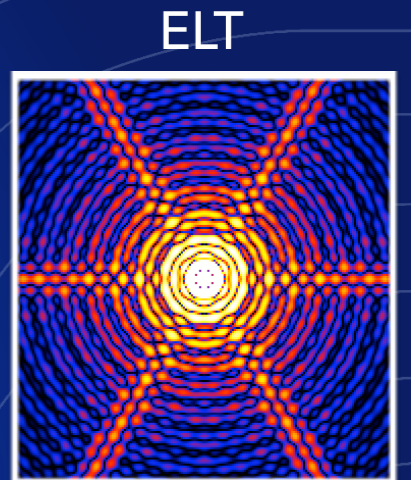
10702-369 KENWORTHY
(Thursday poster)

▶ ERIS: L & M band AGPMs

- * standard vortex coronagraph with simple Lyot stop

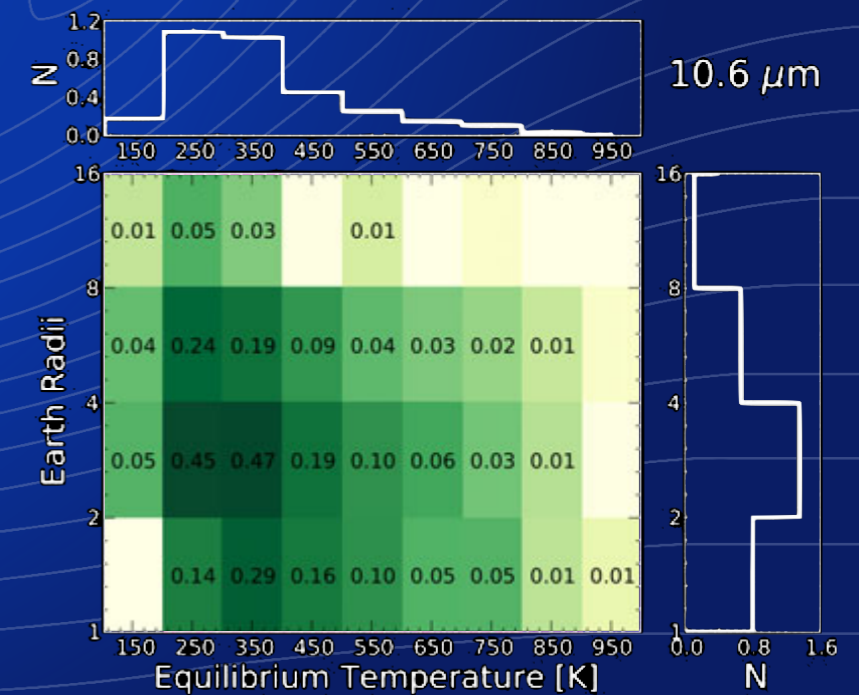
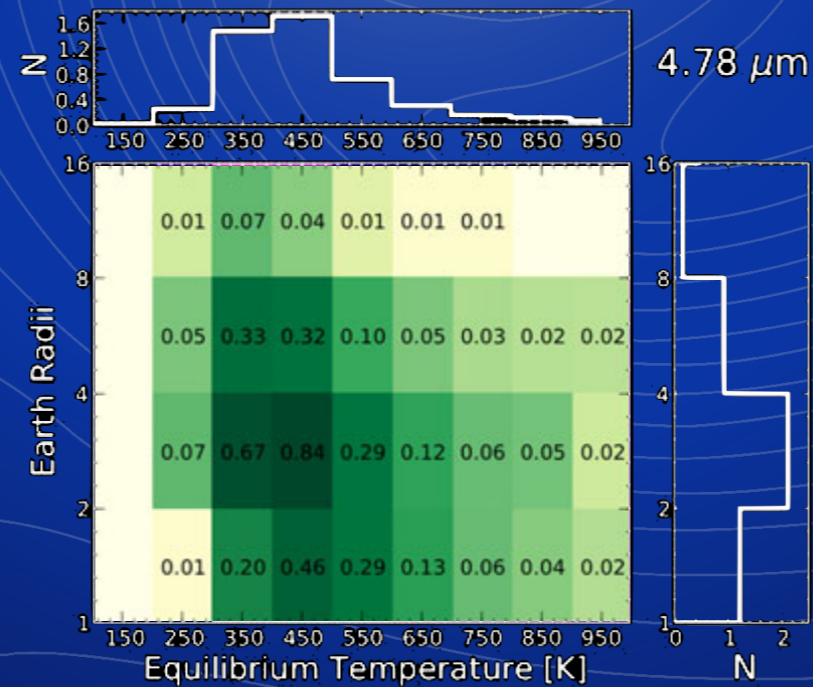
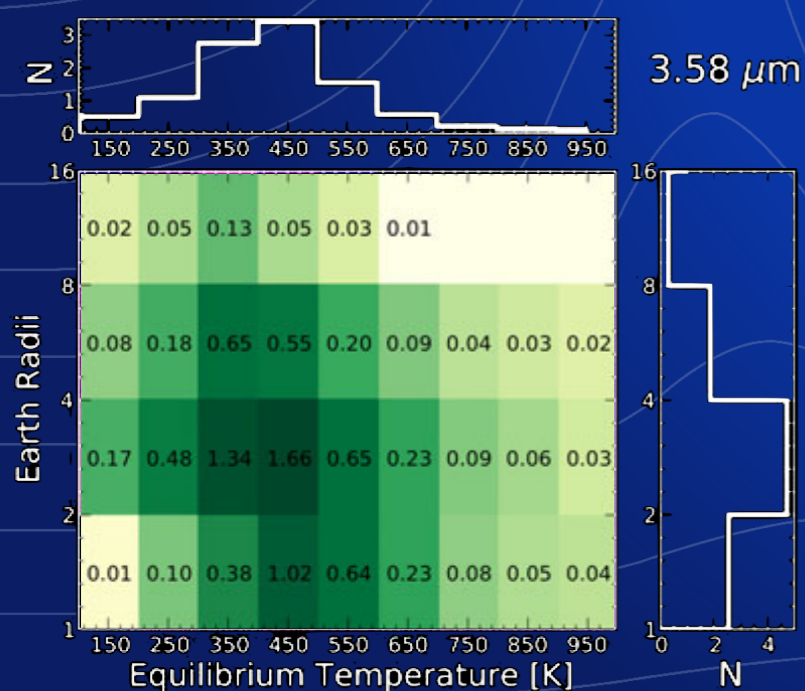
▶ METIS: L, M & N band AGPMs

- * ring-apodized vortex coronagraph: cancels diffraction from huge central obstruction



METIS SCIENCE HIGHLIGHTS

- ▶ direct imaging of several RV planets
- ▶ potential to detect temperate rocky planets
- ▶ exoplanet characterization with high-res LM-band IFS



AROUND THE WORLD WITH

VORTEX



keep light spinning!