

# Magnetic flux avalanches in Nb/NbN thin films

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## Abstract

Technological applications of NbN thin films may be threatened by the development of magnetic flux avalanches of thermomagnetic origin appearing in a large portion of the superconducting phase. In this work we describe an approach to substantially suppress the magnetic flux avalanche regime, without compromising the upper critical field. This procedure consists of depositing a thin Nb layer before the reactive deposition of NbN, thus forming a bi-layered system. We use AC susceptibility and DC magnetometry to characterize both the single layer films, Nb and NbN, and the bi-layered specimen, as well as calibrated Magneto-Optical Imaging to map the instability regime of the studied samples. Magnetic flux imaging reveals interesting features of the dendritic flux avalanches in the bi-layer system, including halo-like patterns and crossing avalanches.

**keywords:** Bi-layer, proximity effect, anti-avalanches, halo-like structure, crossing avalanches.

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# 1 Introduction

According to the Bardeen-Stephen model [1], moving flux quanta dissipate energy due to the existence of an electric field through the normal core of each flux tube. The electric field accelerates the quasiparticles at the core thereby increasing their energy. This energy raising process is normally compensated by the energy relaxation rate of quasiparticles by means of their inelastic scattering [2]. Larkin and Ovchinnikov [3] realized that at high vortex velocities and correspondingly to high electric fields, the energy of quasiparticles can reach the superconducting gap, and diffuse into the superconducting phase surrounding the vortex core. As a consequence, the quasiparticle density in the vortex core is reduced and the vortex shrinks. The higher the vortex speed, the larger the deficit of quasiparticles at the core, the smaller its size and therefore the lower the damping coefficient  $\eta$ . If  $\eta$  decreases with increasing  $v$ , an instability point in the viscous flux flow is reached when the damping force  $\eta(v)v$  starts to decline as  $v$  increases. A single vortex moving at such high velocities will then leave a wake of quasiparticles behind its path which can be regarded as a trail of depleted order parameter. Naturally, other moving vortices will find energetically favorable to follow the same path, and therefore a rearrangement of the Abrikosov vortex lattice is expected [4]. Eventually, these rivers of rapidly moving vortices, directly observed in Ref. [5, 6], can transform into a phase slip line [7]. Note that the mechanism described above involves a non-thermal change of the distribution function of quasiparticles trapped in the vortex cores. Bezuglyj and co-authors [8, 9] theoretically demonstrated that for magnetic fields above a certain threshold, the Larkin-Ovchinnikov instability switches to a pure thermal instability of flux flow resulting from the heating of quasiparticles.

The scenario described above corresponds to a bulk three-dimensional superconductor. When treating thin film geometries, where the penetration depth  $\lambda$  is larger than the thickness  $d$  of the film, an additional complication arises due to the fact that magnetic flux diffusion becomes strongly nonlocal and the vortex interaction is, to a large extent, mediated by the magnetic stray field and the screening of the in-plane supercurrents [10]. In this case, fastly moving flux quanta triggered either by a bias current [11, 12] or by magnetic field changes [13] will then act as precursors of thermomagnetic flux avalanches [14].

42 These events consist of abrupt bursts of magnetic flux rushing into the sample, usually  
43 taking the form of dendrite-like structures. Their branches avoid each other during their  
44 growth [15], as reported both in experiments and in simulations using the thermomagnetic  
45 (TM) model [16, 17]. This model describes the TM instabilities as the source of the flux  
46 avalanches, and predicts the existence of a threshold flux penetration depth ( $l^*$ ) needed to  
47 trigger them. Once the penetration depth reaches  $l^*$ , the first burst takes place and the  
48 instability regime lasts until  $l^*$  is equal to half of the sample size [18, 19, 20]. Therefore, an  
49 upper and lower threshold fields [20] can be identified as the borders of the instability region  
50 for a isothermal field ramping. By means of magnetization measurements, Colauto *et al.* [21]  
51 were able to delineate a region on the magnetic field-temperature diagram where instabilities  
52 occur in a 200 nm thick Nb film. Among other materials, dendritic flux avalanches were also  
53 observed in NbN films, spanning over a large window of fields and temperatures [22].

54 Niobium nitride has higher critical field ( $H_{C2}(T)$ ) and critical temperature ( $T_C$ ) than  
55 pure Nb [23], what makes it NbN more suitable for superconducting devices [24], such as  
56 hot-electron bolometers [25], high-frequency superconducting circuits [26, 27], single-photon  
57 detectors [28], and qubits for quantum computers [29], to name a few. Some applications  
58 may suffer from unwanted TM instabilities. Yurchenko *et al.* [30] have shown that such  
59 abrupt phenomena can be prevented by covering the NbN film with a copper coating. In-  
60 dependently, the idea of depositing a superconducting coating to enhance the applicability  
61 range was investigated in references [31, 32, 33]. In these works it was shown that the sta-  
62 bility improvement in superconducting wires covered by a thin superconductor depends on  
63 the electrical and thermal properties of both. Such a system becomes more stable if the  
64 capping layer has lower critical current density and higher heat capacity. Moreover, Ivry  
65 and co-workers [34] have proposed to use a thin proximitized bi-layer structure NbN/WSi to  
66 optimize the performance of superconducting single photon detectors.

67 Stacks of overlapping but electrically disconnected superconducting thin films is another  
68 way to affect the avalanche regimes, as demonstrated by Tamegai *et al.* [35]. They studied the  
69 critical states and thermomagnetic avalanche activities in three-dimensional nanostructured  
70 superconductors, i.e., stacks of Nb strips, insulated from each other by SiO<sub>2</sub> layers. Here,  
71 it was shown that flux avalanches can start in one layer and end at another. Although the

72 avalanche activity has been shown to be reduced in superconducting films capped with a  
73 normal metal [30, 36], thermomagnetic instabilities were not reported so far in bi-layered  
74 systems where NbN is in intimate contact with other superconducting layer. In this context,  
75 we present in this paper an approach capable of enhancing the potential applicability of NbN  
76 thin films by systematically exploring the avalanche regime in a bi-layer system composed of  
77 Nb and NbN thin films.

## 78 2 Experimental Details

79 In order to perform a comparative analysis of the magnetic flux avalanche regime, 15 nm thick  
80 Nb (Nb15) and 60 nm thick NbN (NbN60) single films and hybrids (NbN/Nb) were deposited  
81 on Si(100) substrates at room temperature in a UHV DC diode magnetron sputtering system  
82 with a base pressure in the low of  $10^{-8}$  mbar range. The Ar pressure, during the deposition of  
83 the Nb layer, was  $P_{Ar} = 2.5 \cdot 10^{-3}$  mbar, while NbN was reactively sputtered in an atmosphere  
84 of Ar and  $N_2$ , with  $P_{Ar} = 2.5 \cdot 10^{-3}$  mbar and  $P_{N_2} = 0.7 \cdot 10^{-3}$  mbar. The deposition rates  
85 were  $r_{Nb} = 0.26$  nm/s for Nb and  $r_{NbN} = 0.17$  nm/s for NbN, as measured by a quartz  
86 crystal monitor previously calibrated by measuring the step height of photolithographically  
87 patterned films with a Bruker DektakXT stylus profiler. Samples with different structures  
88 were deposited by keeping the thickness of the Nb and NbN individual layers constant, namely  
89  $d_{Nb} = 15$  nm and  $d_{NbN} = 60$  nm.

90 All the studied samples, having approximately the same area of  $4 \times 4$  mm<sup>2</sup>, were char-  
91 acterized by AC susceptometry and DC magnetometry in a commercial MPMS 5 Quantum  
92 Design magnetometer. Magneto-Optical Imaging (MOI) experiments, based on the Faraday  
93 effect [37], was carried out by placing a  $Bi_x Y_{1-x} FeO$  indicator [38] on top of the supercon-  
94 ducting film. More details about the MOI setup can be found elsewhere [39, 40]. In all  
95 measurements, the field  $H$  was applied perpendicular to the film surface. We also performed  
96 a numerical conversion from pixel intensity of magneto-optical (MO) images to the local  
97 magnetic flux density ( $B$ ), mapping  $B_z$  all over the sample and its neighborhood, following  
98 the protocol reported in Ref. [41].

### 99 3 Results and Discussion

#### 100 3.1 Upper critical field

101 The critical temperature of superconducting thin films is thickness dependent and usually  
 102 lower than the bulk values [42]. The onset critical temperature ( $T_C$ ) of the samples were de-  
 103 termined by AC susceptibility measurements, presented in Figure 1(a), showing the following  
 104 values:  $(6.90 \pm 0.05)$  K for Nb15;  $(10.50 \pm 0.05)$  K for the NbN60, and  $(10.00 \pm 0.05)$  K for the  
 105 hybrid sample. The critical temperature for the NbN film is close to values reported in the  
 106 literature for similar thicknesses [43, 44]. It is also important to mention that  $T_C$  of the single  
 107 NbN layer is 0.5 K above of that the bi-layer, which is assumed here to be a consequence of  
 108 growing the NbN film on top of the Nb layer already deposited on the substrate.

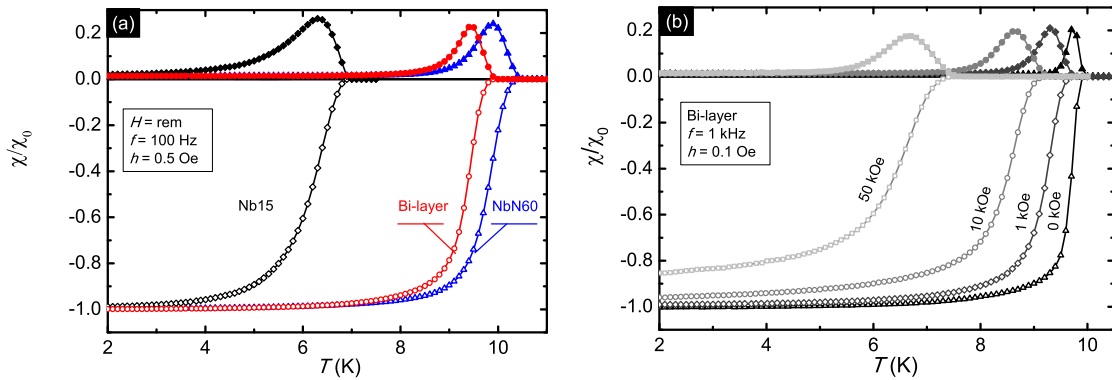


Figure 1: AC susceptibility *vs* temperature for all the investigated films (a) and for different applied magnetic fields  $H$  for the bi-layer system (b). The  $y$  axis is normalized by  $\chi_0$ , which is the Meissner plateau value for the in-phase component of the AC susceptibility for each sample. The frequency ( $f$ ) and the amplitude ( $h$ ) of the AC excitation are indicated in each panel.

109 The fact that we do not observe a double or a broader transition in the bi-layer with  
 110 respect to the single layer films is a relevant hint suggesting highly transparent proximity  
 111 effect [45] in between the layers. The absence of a double transition in the bi-layer sample  
 112 remains for applied magnetic fields up to  $H = 50$  kOe, as shown in Figure 1(b). The existence  
 113 of the proximity effect in the bi-layer specimen was also confirmed by a double transition in

114 a sample with an additional 5 nm thick Nb<sub>2</sub>O<sub>5</sub> insulating layer between the superconducting  
 115 films (data not shown here).

116 By performing susceptibility measurements as a function of temperature for  $H$  up to  
 117 50 kOe, we determined the  $H_{C2}$  vs  $t$  diagram presented in Figure 2,  $t$  being the reduced  
 118 temperature,  $t = \frac{T}{T_c}$ . We estimate the  $H_{C2}(0)$  values by fitting to the data the expression  
 119  $H_{C2}(t) = H_{C2}(0) \cdot (1 - t^2)$ , plotted as dashed lines in the same graph. For both the bi-layer  
 120 and NbN60 films,  $H_{C2}(0)$  is close to 110 kOe, whereas for the Nb15 film it is approximately  
 121 27 kOe. Based on the derivative of the upper critical field versus temperature near  $T_C$ , we  
 122 determined the superconducting coherence lengths at 0 K ( $\xi(0)$ ) of 10.7 nm, 4.7 nm, and  
 123 4.6 nm, for Nb15, NbN60, and the bi-layer, respectively. We did not detect flux avalanches  
 124 in AC susceptibility measurements using driving fields up to 3.8 Oe, consistently with the  
 125 existing literature [46, 47], since avalanches occur only at higher fields.

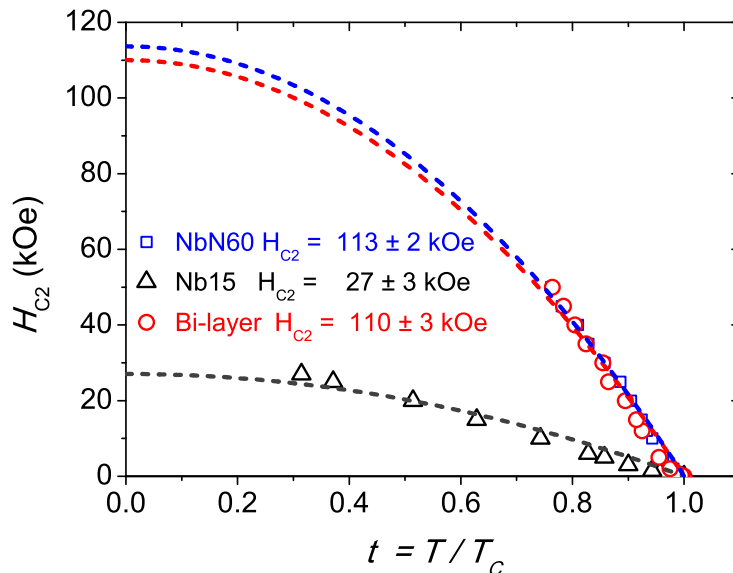


Figure 2:  $H_{C2}$  versus reduced temperature  $t = T/T_c$  for Nb15, NbN60 and the bi-layer, showing the extrapolated upper critical field at  $T = 0$  K.

### 126 3.2 Flux jumps regime

127 In order to identify the instability regimes of these systems, we measured the DC magnetiza-  
 128 tion as a function of the applied magnetic field at the same reduced temperature  $t = 0.3$ . The

129 result is presented in Figure 3. The presence of magnetic flux jumps is clearly identified, for  
 130 all samples as a noisy magnetic response, being particularly prominent for the NbN sample.  
 131 Note, however, that avalanche activity is strongly suppressed by the proximitized Nb layer  
 132 (bi-layer sample).

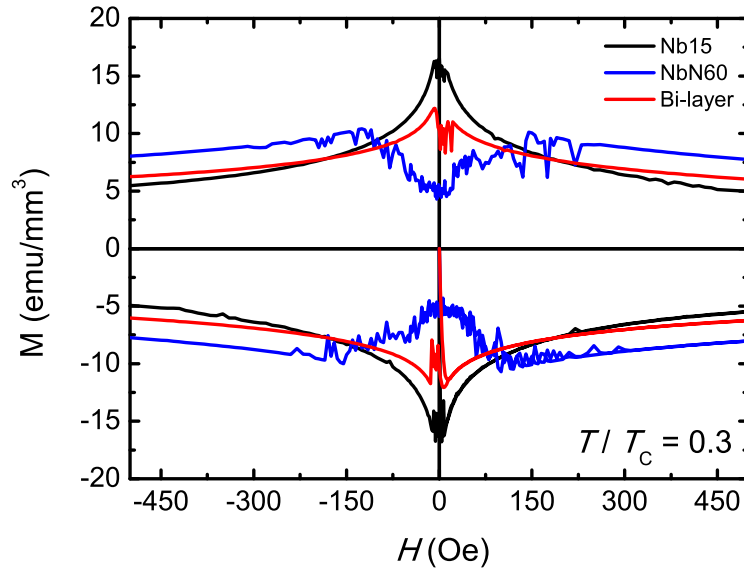


Figure 3: DC magnetization as a function of the applied magnetic field at  $t = \frac{T}{T_C} = 0.3$ . The noisy response at low fields observed in all samples corresponds to the presence of magnetic flux jumps.

133 The critical current density ( $J_C$ ) is a crucial parameter determining whether the ther-  
 134 momagnetic avalanches will take place [48]. The higher the  $J_C$  the larger the probability of  
 135 observing flux jumps. Based on the Bean critical state model [49], one can roughly estimate  
 136  $J_C$  by the difference between the increasing and decreasing branches of the magnetization  
 137 loop. This approach is acceptable in the smooth part of the magnetization loop (i.e. with-  
 138 out flux jumps). A direct inspection of Figure 3 shows that the critical current densities  
 139 are rather similar for all samples, whereas the avalanche activity in the bi-layer sample has  
 140 decreased as compared to that in the NbN single layer.

### 141 **3.3 Magneto-Optical Imaging**

142 Flux avalanches disrupting the smooth penetration after a zero field cooling (ZFC) procedure  
143 can be visualized in the MO images of Figure 4(a), for each of the investigated films. In all  
144 those MO images, the brighter the pixel, the higher perpendicular flux density. While large  
145 dendritic flux avalanches are observed in both NbN60 and Nb15 films, the bi-layer system  
146 exhibits much less activity, only some small finger-like avalanches occurring from the left and  
147 right edges. By changing the temperature, magnetization loops allow one to delineate the  
148 instability region in the applied magnetic field versus reduced temperature ( $H-t$ ) diagram  
149 shown in Figure 4(b). This figure presents one of the main messages from this work, namely  
150 a substantial enlargement of the stability regime, i.e. where only smooth flux penetration  
151 occurs, of the bi-layer system as compared to the bare NbN film. In other words, the bi-layer  
152 instability regime (in green) shrinks toward that of the Nb15 one (in yellow).



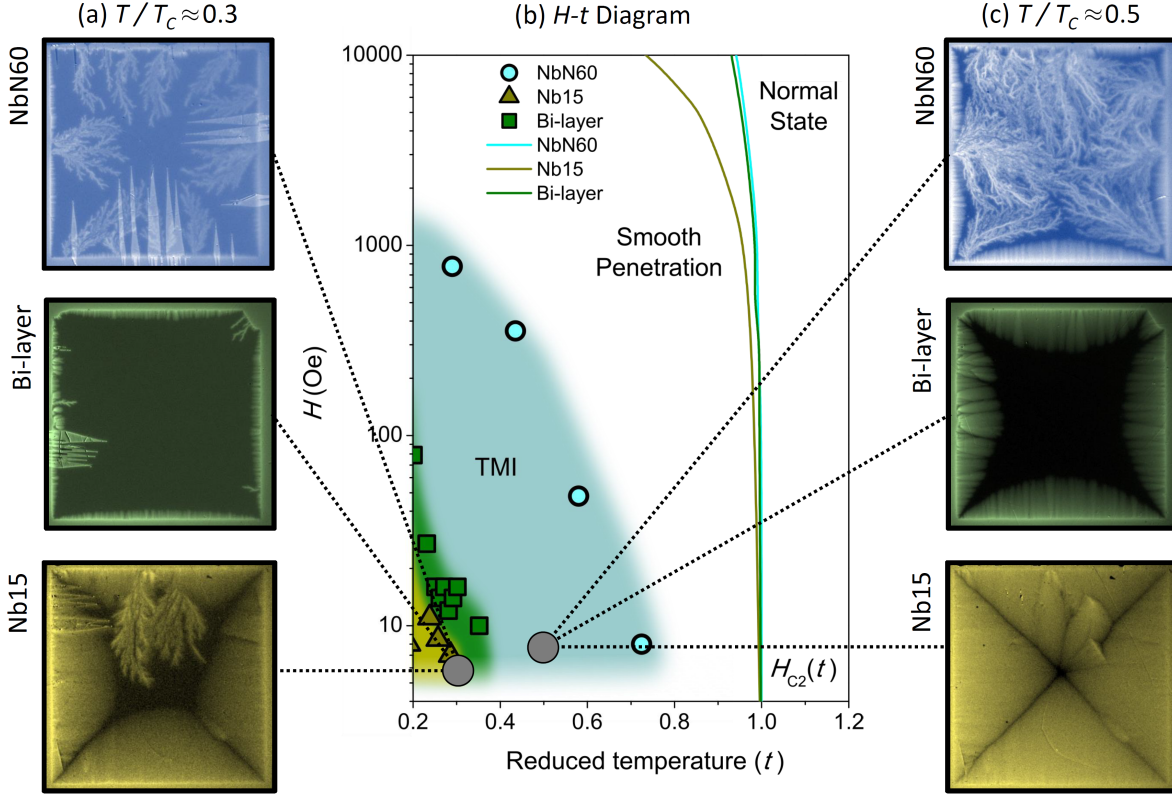


Figure 4: Column (a) shows MO images for each studied samples at  $t \approx 0.3$  and  $H = 4$  Oe. The zig-zag-like features in the images are related to domain walls in the indicator film. (b)  $H-t$  diagram showing the thermomagnetic instabilities regime (TMI) as a function of the reduced temperature; (c) MO images taken at  $t \approx 0.5$  and  $H = 8.5$  Oe.

153 When  $t \approx 0.5$ , the NbN60 sample exhibits avalanches, as presented in the top MO image  
 154 in Figure 4 (c). Both the bi-layer and the Nb15 films show smooth penetration, with the  
 155 latter one in the fully flux penetrated state. All the flux avalanches presented in Figure 4  
 156 show positive flux only, i.e., they were created following the virgin curve of the magnetization  
 157 loop by increasing the applied field from zero. By decreasing the applied magnetic field in  
 158 a superconducting film, after keeping flux trapped into the sample, negative field-polarity  
 159 avalanches, or simply anti-avalanches, can occur. Anti-avalanches can show an annihilation  
 160 zone [15, 50], i.e., a boundary of zero flux density separating the regions of flux and antflux,  
 161 which coexists due to the application of moderate reverse fields in a sample with flux already  
 162 trapped by the pinning centers. This terminology has been used to describe the contour

163 of anti-avalanches in the early stage of MO investigations of the abrupt flux penetration in  
164 superconducting thin films [15].

165     Once anti-avalanches are created by decreasing the applied magnetic field, their onset  
166 depends on the previous magnetic history of the system. Figure 5(a)-(e) presents quantita-  
167 tive MO images obtained at certain magnetic fields along the hysteresis loop of the NbN/Nb  
168 bilayer sample at  $T = 3.5$  K. The spatial profile of the induction component  $B_z(r)$  at the bot-  
169 tom of each image has been obtained from an average of 40 lines as shown by the translucent  
170 yellow bar in panel (a).

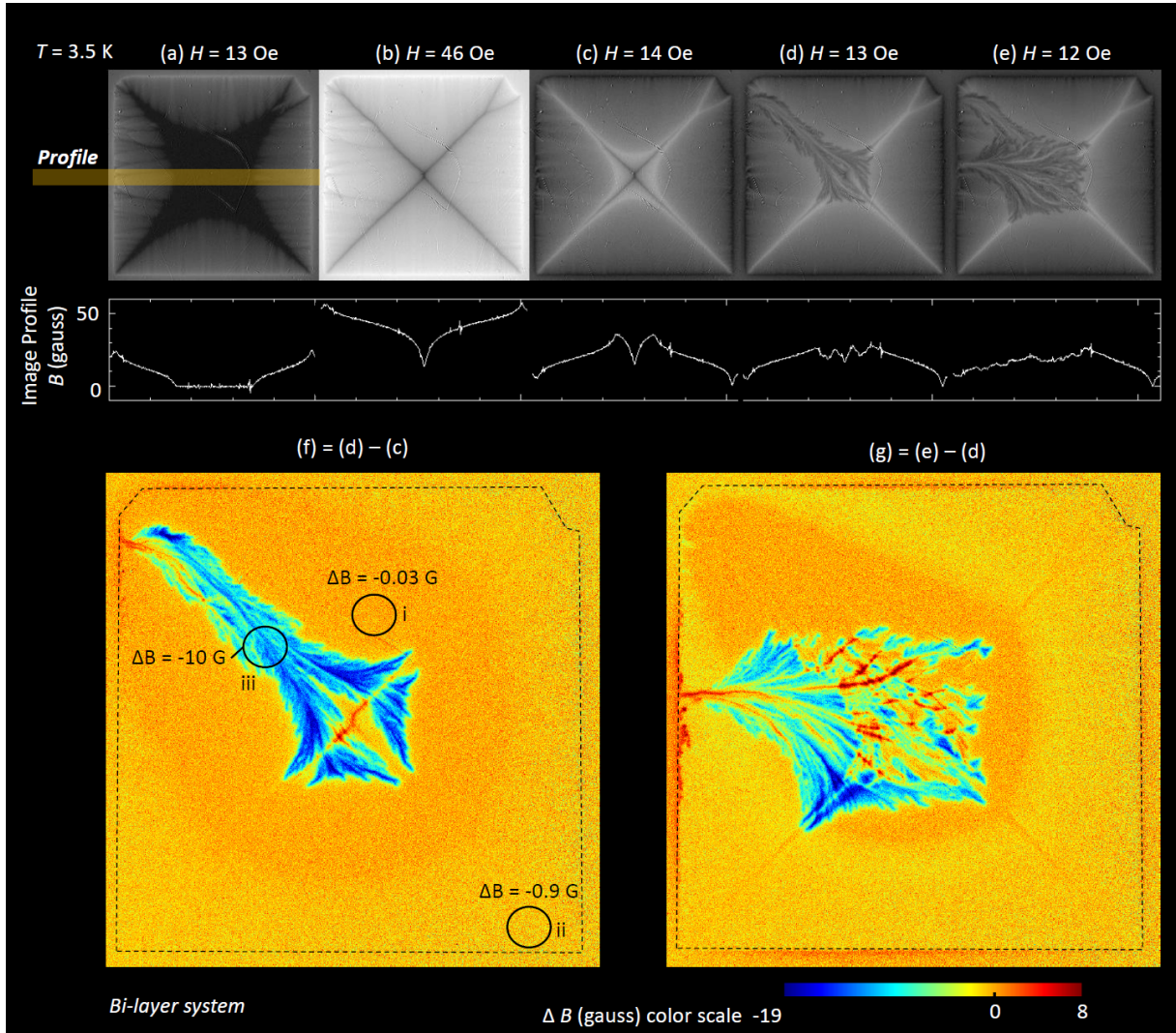


Figure 5: (a)-(e) A sequence of quantitative MO images of the bi-layer system, for different applied magnetic fields, after a ZFC procedure, at  $T = 3.5 K$ . At the bottom of each panel the field profile is shown, averaged from 40 lines delimited by the translucent yellow bar identified in panel (a). (f) Differential image obtained by subtracting (d) and (c) panels, which shows the first anti-avalanche and an orange halo surrounding it. The field variation  $\Delta B$  is indicated by black circles, (i) inside the halo, (ii) outside the halo, and (iii) inside the avalanche. (g) Differential image between panels (e) and (d), where the second anti-avalanche crosses the first one. The color scale indicates that the trapped field increased in some regions where the avalanche branches cross. The dashed lines in panels (f) and (g) are indications of the sample edges.

171 Figure 5(a) shows a typical critical state-like field profile for the virgin curve in a magne-  
 172 tization loop where the inner part of the film is still in the Meissner state (dark inner area),  
 173 i.e.,  $B = 0$ . In panel (b), the applied field reaches its maximum value ( $H = 46$  Oe), and  
 174  $B > 0$  at the center of the sample. The diagonal dark lines forming an X shape pattern  
 175 are named discontinuity lines (d-lines), and delineate the locations where the supercurrent  
 176 undergoes an abrupt change of direction. Panel (c) shows the flux density landscape after  
 177 decreasing  $H$  down to 14 Oe starting from its maximum value, and just before the occurrence  
 178 of the first anti-avalanche in the system. The field profile in panel (c) reveals a large quantity  
 179 of positive flux trapped in the sample. The first anti-avalanche (d) starts from the top left  
 180 corner into the positive upper left d-line. This preferential track suggests that most likely  
 181 this avalanche is driven by the flux-antiflux annihilation process. The magnetic profile at the  
 182 bottom of panel (d) shows the recorded imprint of this anti-avalanche, and it does not change  
 183 the polarity of the induction field  $B$  along its path, but strongly decreases the local field as it  
 184 passes. By decreasing the applied magnetic field by 1 Oe, another anti-avalanche is triggered  
 185 from the left edge, transpassing the center of the sample and then crossing the first avalanche  
 186 of anti-flux. The second anti-avalanche does not change the local field to negative values,  
 187 although it decreases further the average  $B$  in the whole sample. Nevertheless, these two  
 188 anti-avalanches exhibit particular features that can be better emphasized by implementing  
 189 differential MOI [51], i.e. by subtracting consecutively recorded images. The result of this  
 190 procedure is presented in panels (f) and (g) of Figure 5.

191 Note that the first anti-avalanche running along the d-line produces end branches directed  
 192 along the crossing d-line. A remarkable feature is the appearance of a halo surrounding the  
 193 anti-avalanche, a feature that, to the best of our knowledge, has not been reported so far.  
 194 To describe this halo in a quantitative form, we measured the average variation of  $B$  in three  
 195 circular regions with 25,000 pixels each in different regions throughout the sample. The  
 196 result is marked by the black circles seen in panel (f). The circle (i), inside the halo itself, is  
 197 the region where the local field decreased less ( $\Delta B_i = -0.03$  G). This procedure was done  
 198 in other points across the halo (not shown), to confirm this observation. Outside the halo,  
 199 the circle (ii) results in  $\Delta B_{ii} = -0.9$  G, and inside the avalanche (circle (iii)), the average  
 200 flux density variation from this area was  $\Delta B_{iii} = 16$  G  $-$   $26$  G  $= -10$  G. According to the

201 color scale, one can see that there are regions in the anti-avalanche branches where the field  
 202 variation is as high as  $-19$  G. The differential MO image in panel (f) allows one to state that  
 203 the average field in the sample decreased. The trapped flux in the system seems to lead to  
 204 this unexpected halo. More details on the halo structure and its surroundings are provided  
 205 in the supplementary material. The sample Nb15 also shows a halo-like structure around  
 206 its first anti-avalanche, but this halo was not detected in the film NbN60. The halo is not  
 207 a thick annihilation zone, as one can see in the  $B$  profile of Figure 5 (d) and (e) – there  
 208 is no crossover between positive to negative flux there, and thus, no zero-field region. This  
 209 halo refers to the absence of rearrangement of the flux distribution in the region around the  
 210 abrupt penetration during the first anti-avalanche.

211 Another intriguing aspect of this set of images is that the second anti-avalanche crosses  
 212 the first one. The color scale in Figure 5(g) allows one to highlight the fact that the branches  
 213 of the first avalanche transpassed by the second one, undergo a positive variation of the local  
 214 magnetic field as high as 8 G. Flux avalanches triggered during a ZFC procedure are known  
 215 for avoiding each other during their propagation into the sample [15], no matter whether they  
 216 are small and fingerlike or large and highly branched. However, avalanches may cross each  
 217 other in descending fields because there is still positive flux where the prior anti-avalanche  
 218 passed. Although the halo of the first anti-avalanche has changed after the advent of the  
 219 second avalanche, this last one does not have a halo surrounding it.

### 220 3.4 Halo definition

221 What we call halo is a region of extra brightness (in our case,  $\Delta B$ ) surrounding the first  
 222 anti-avalanche for the bi-layer system. Figure 6 (a) is a differential image, as presented in  
 223 Figure 5 (f). Panels (b), (c) and (d) are the averaged  $\Delta B$  profiles for three regions of the  
 224 sample, indicated by translucent gray bars. Panel (b) shows the  $\Delta B$  profile passing through  
 225 the avalanche trunk, where there is an intense negative variation of  $B$  ( $\Delta B < 0$ ), as well as a  
 226 smooth variation close to the sample edges (outside the halo). In (c), the halo region presents  
 227 the highest brightness in the whole image ( $\Delta B = 0$ ). Panel (d) present a region outside the  
 228 halo where  $\Delta B$  is negative and constant. Therefore, the halo is suitable to describe such a  
 229 region in the framework of differential images.

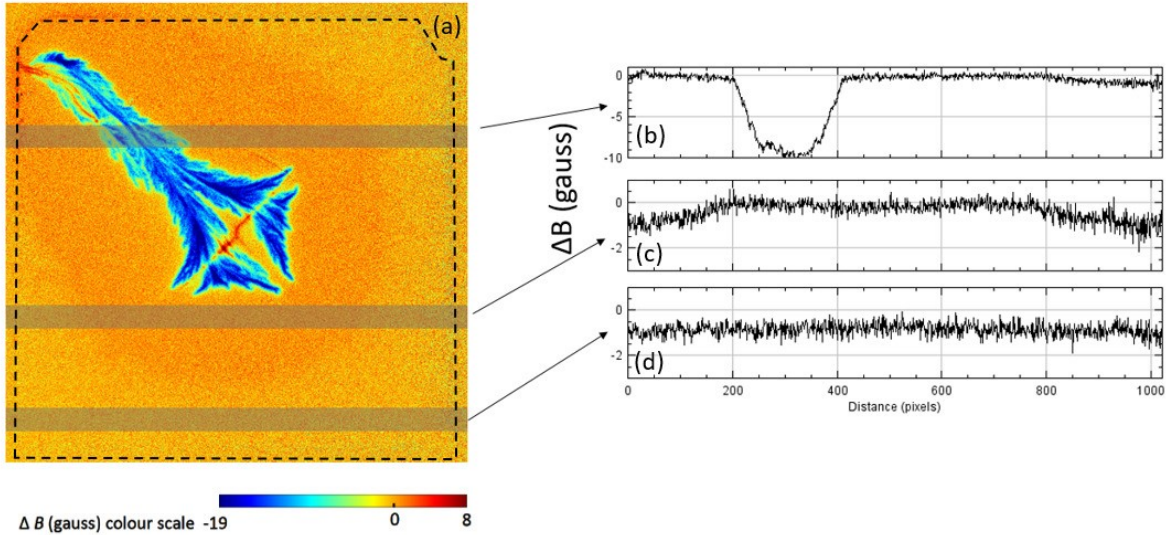


Figure 6: Description of the halo structure. (a) Differential image obtained by subtracting (d) and (c) panels of Figure 5, presented as panel (f). Averaged  $\Delta B$  profiles taken from the translucent gray bars, passing through (b) the anti-avalanche, (c) the halo, and (d) outside the halo.

## 230 4 Conclusions

231 In this paper we propose an approach to reduce the flux avalanche activity in NbN films  
 232 by coating them with a Nb layer. This measure may improve the applicability of thin  
 233 films of NbN without changing its upper critical field at the same reduced temperature.  
 234 A similar effect has been reported when the superconductive film is coated with a normal  
 235 metal [30, 36, 52, 53, 54, 55]. The region where the avalanches take place in the field-  
 236 reduced temperature diagram decreases for the hybrid system as compared to the single NbN  
 237 layer, becoming closer to the single Nb film. In other words, there is a suppression of the  
 238 occurrence of flux avalanches in the hybrid NbN/Nb system without considerably depreciating  
 239 its other properties. In addition to that, quantitative MOI allows one to unveil anti-avalanches  
 240 crossing, as well as the lack of vortex rearrangement in a large region surrounding the first  
 241 anti-avalanche. This latter effect manifests itself as a halo of nearly unperturbed magnetic  
 242 field intensity.

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