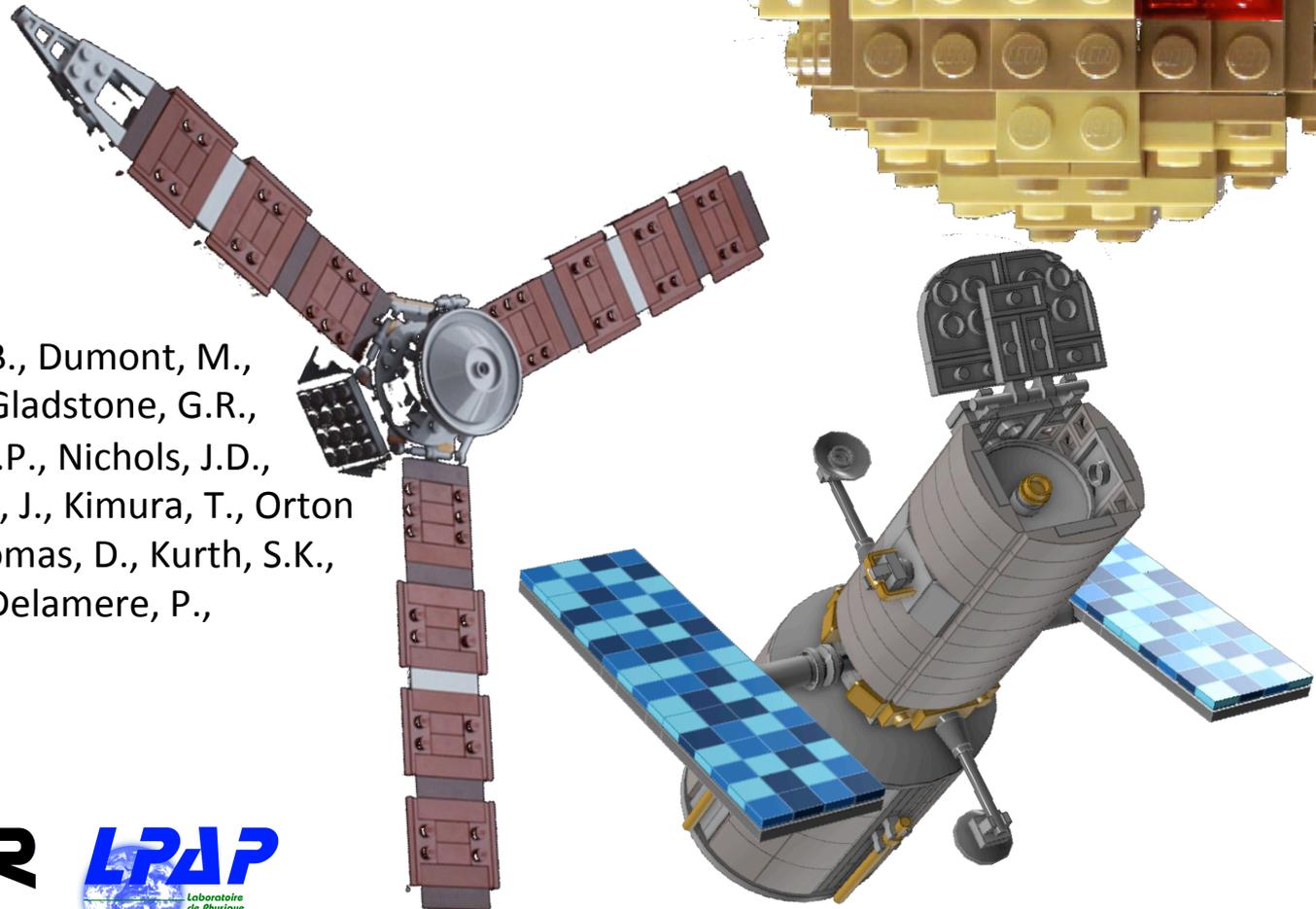


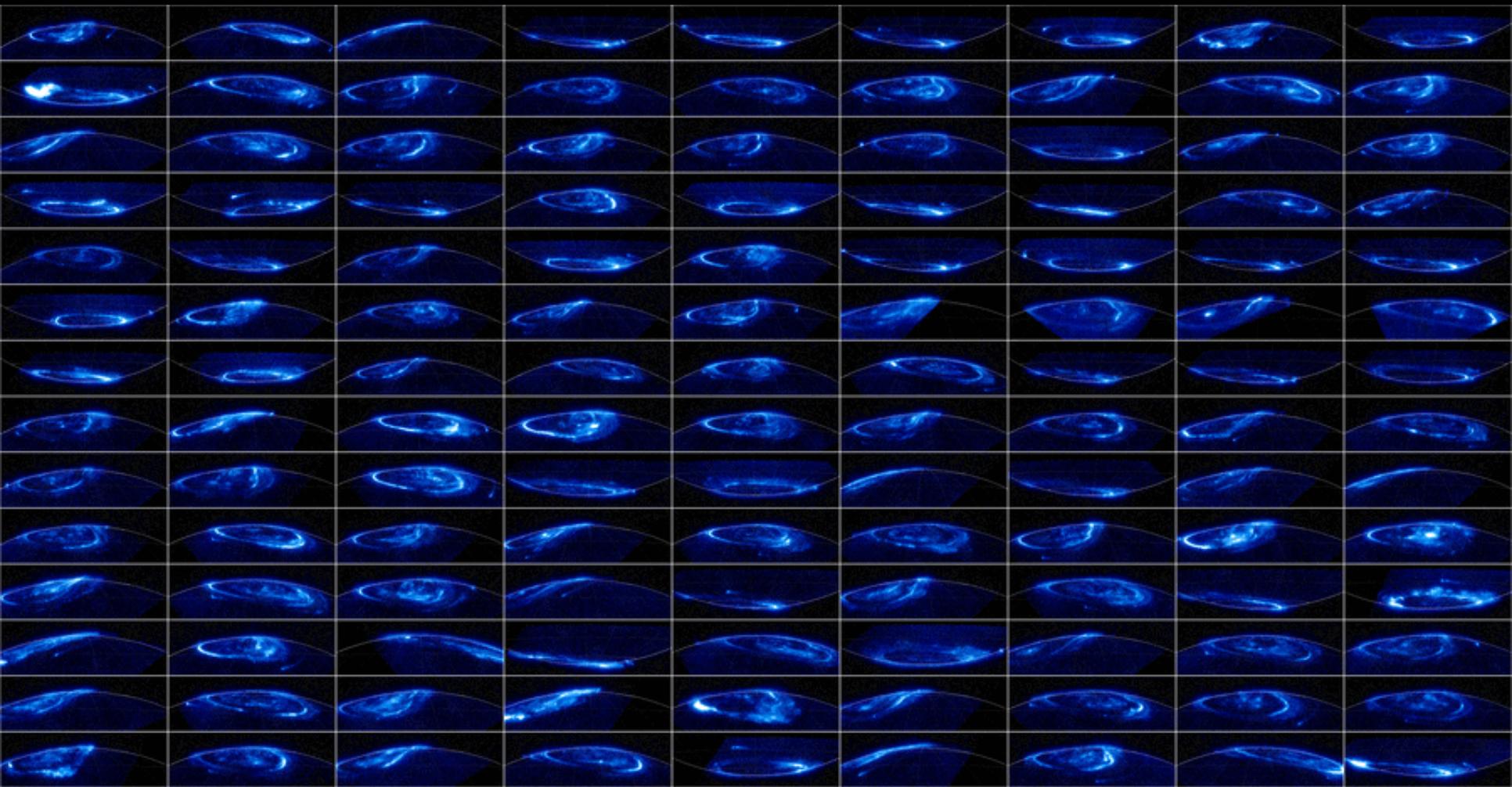
# Jupiter's mesmerizing auroral show (PJ13);

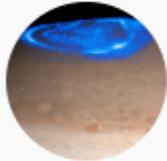
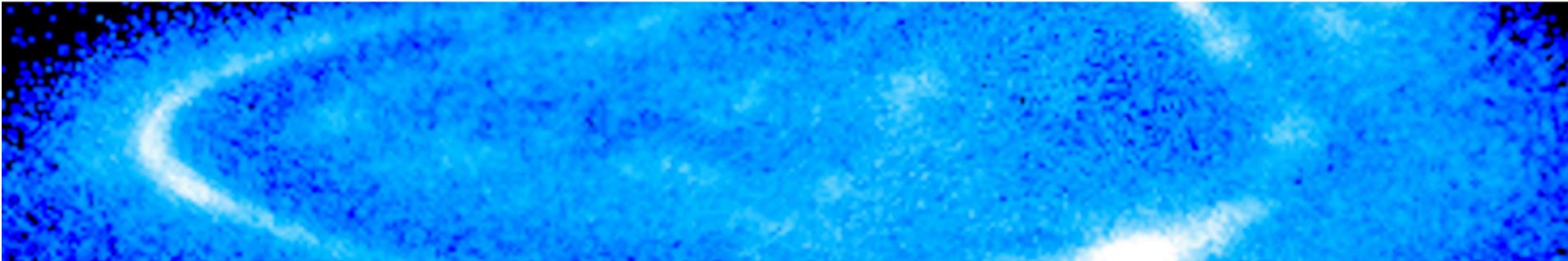
HST ultraviolet observations near and far from Juno perijoves

**Denis GRODENT,**  
Zhonghua YAO,

Bonfond, B., Palmaerts, B., Dumont, M.,  
Gérard, J.C., Radioti, A., Gladstone, G.R.,  
Mauk, B., Connerney, J.E.P., Nichols, J.D.,  
Bunce, E.J., Roth, L., Saur, J., Kimura, T., Orton  
G.S., Badman, S.V., McComas, D., Kurth, S.K.,  
Adriani, A., Valek, P.W., Delamere, P.,  
Bagenal, F.







Denis Grodent

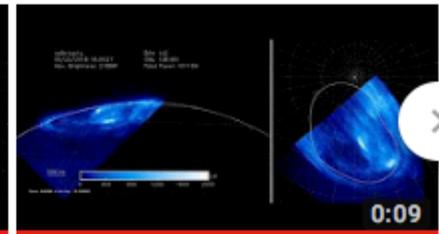
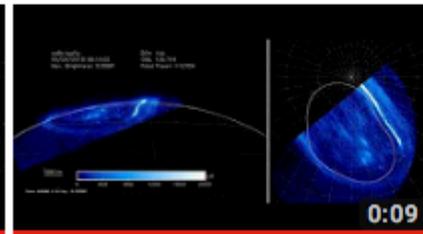
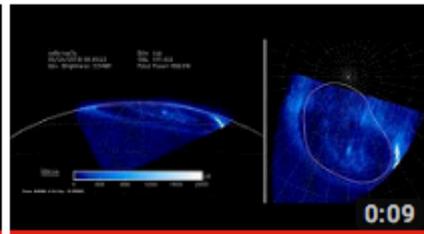
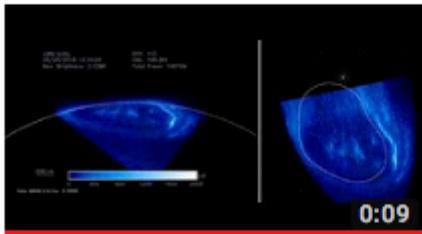
16 subscribers

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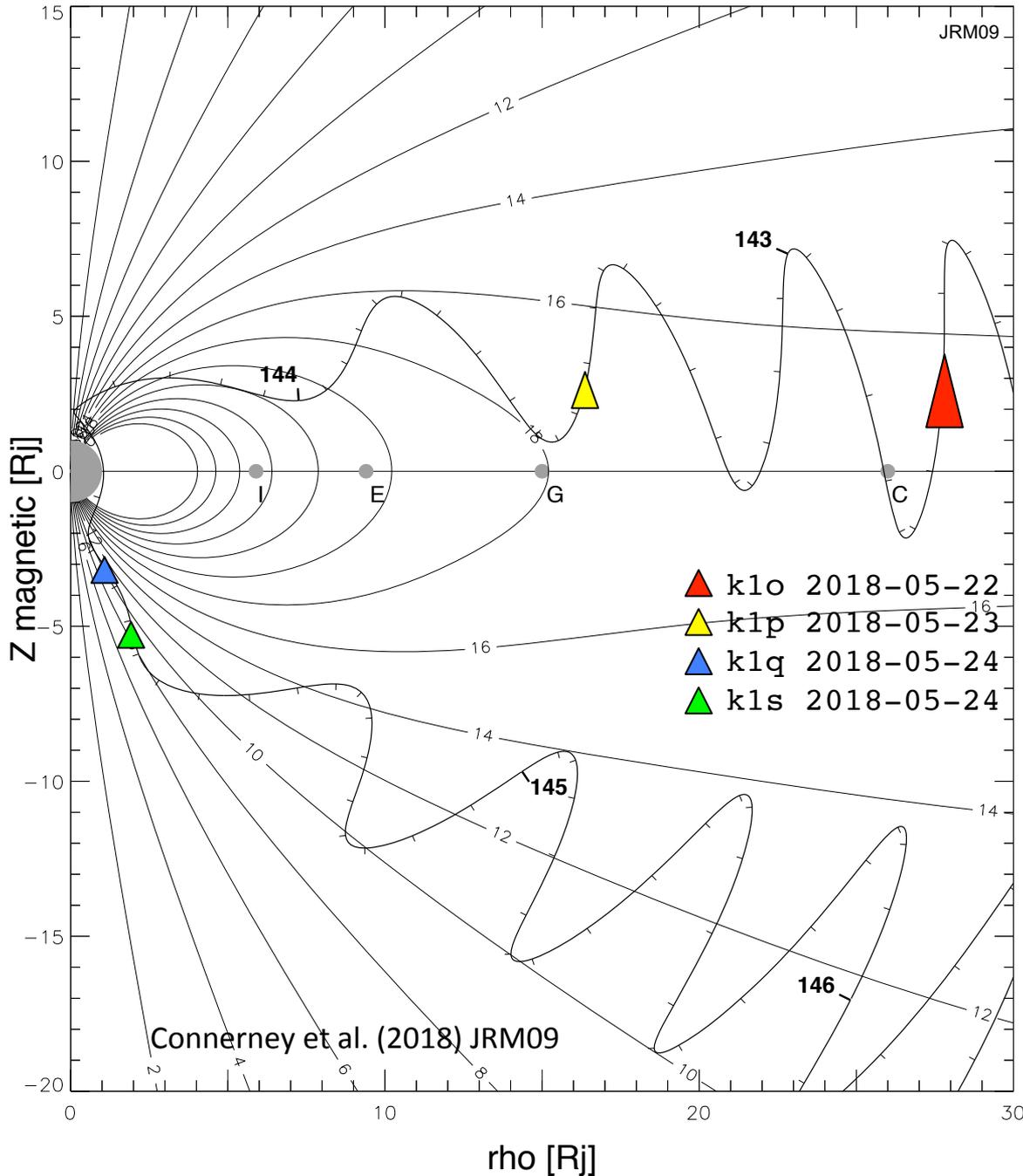
HST movie of Jupiter's ultraviolet aurora captured i...



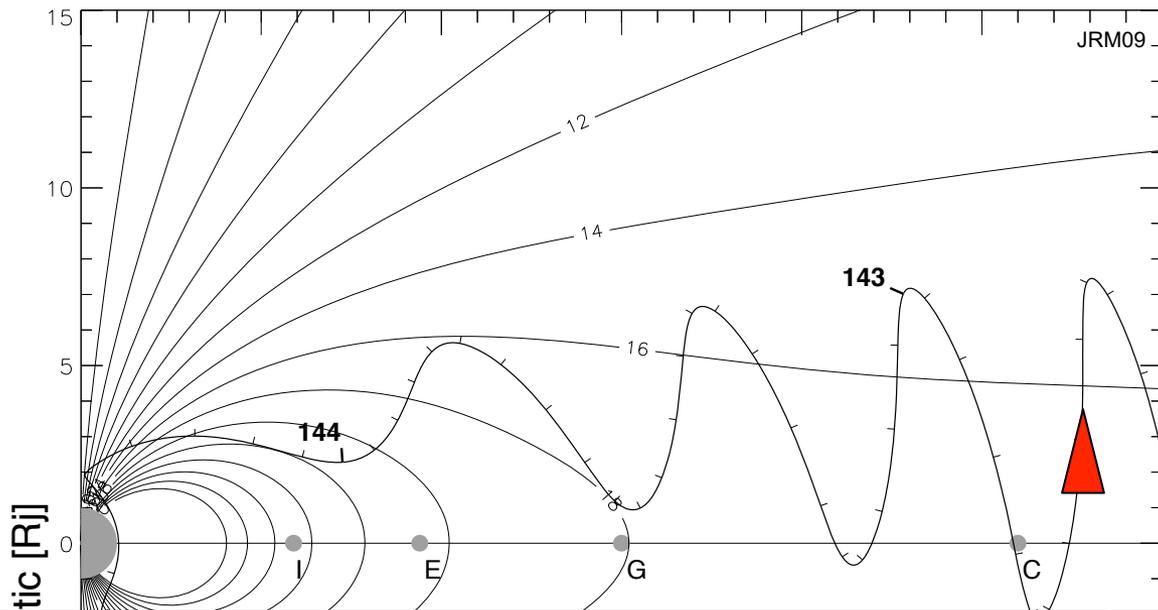
# PJ13

HST GO-14634  
 STIS ttag  
 Jupiter North  
 fUV Aurora

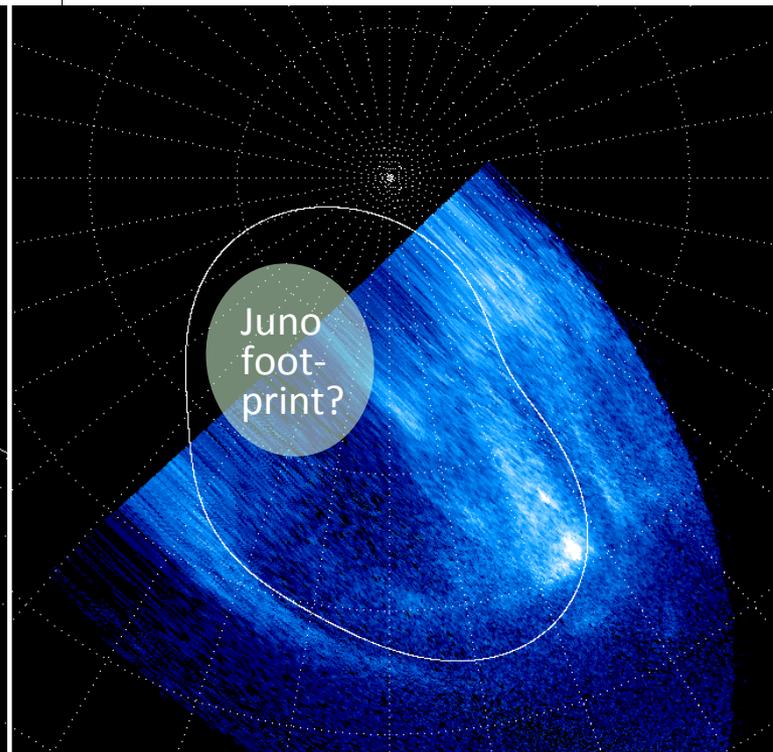
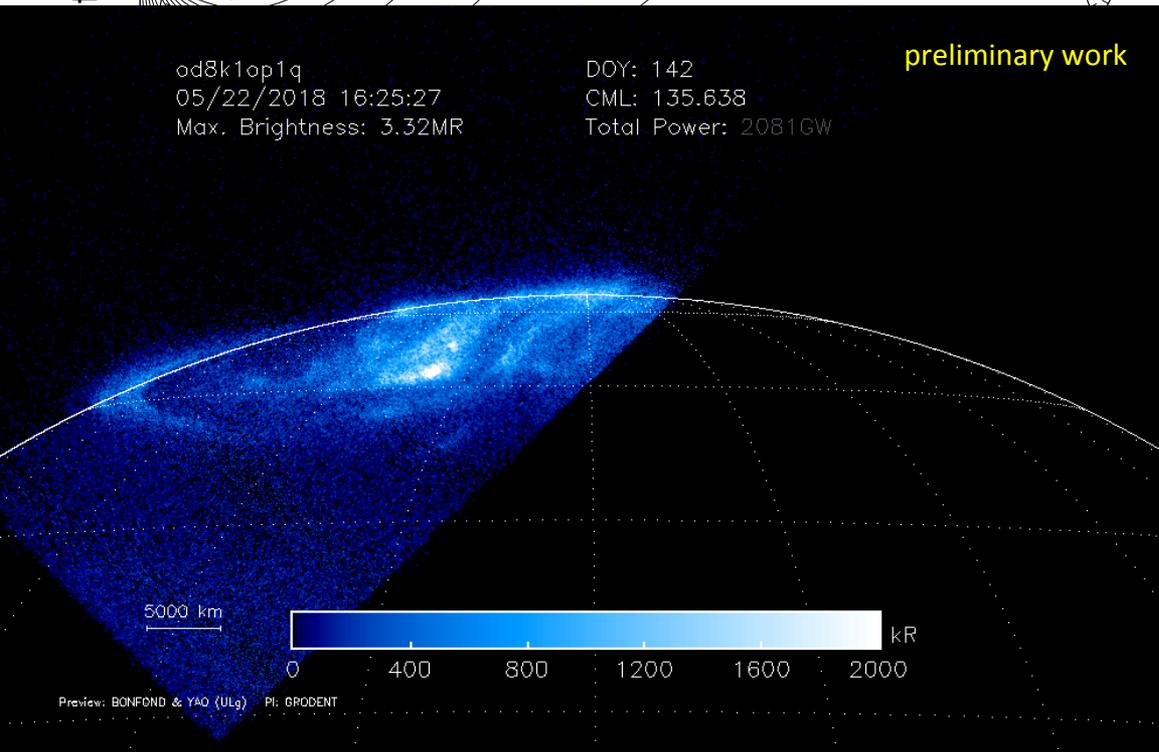
PJ13

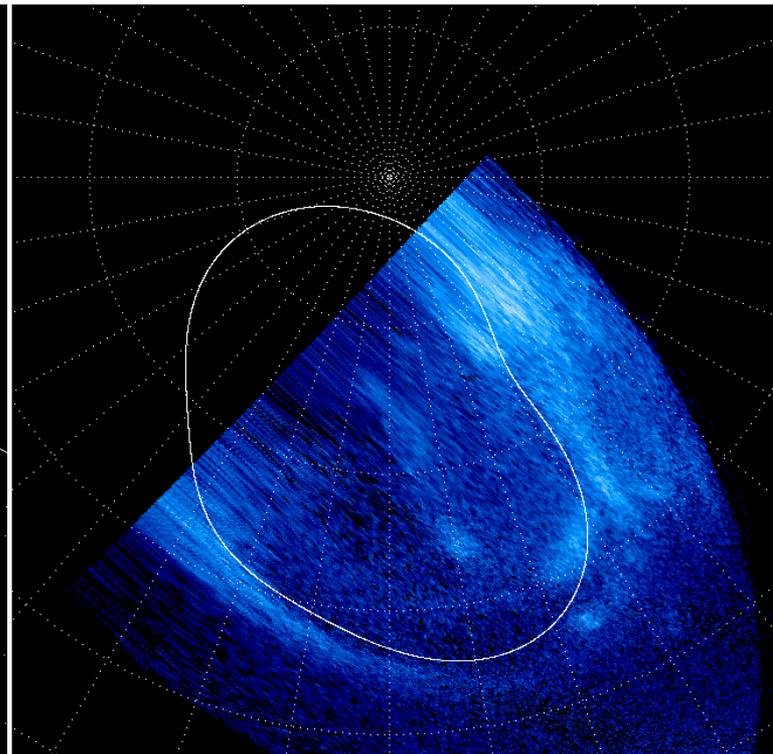
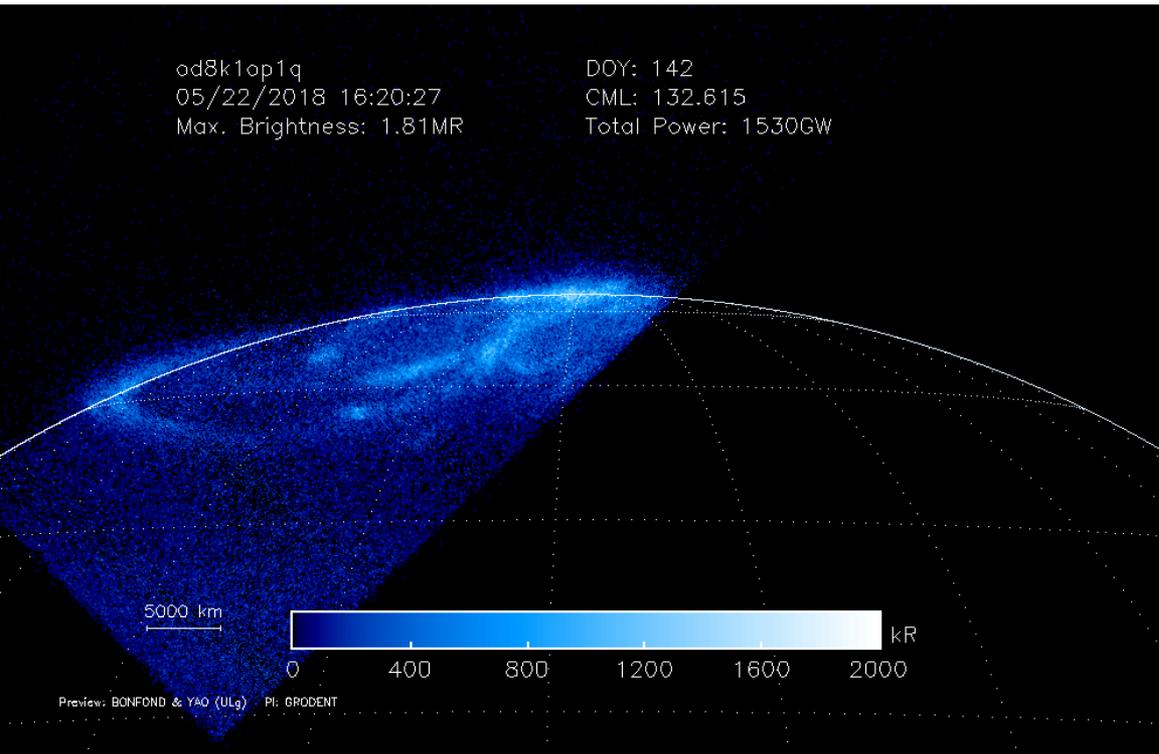


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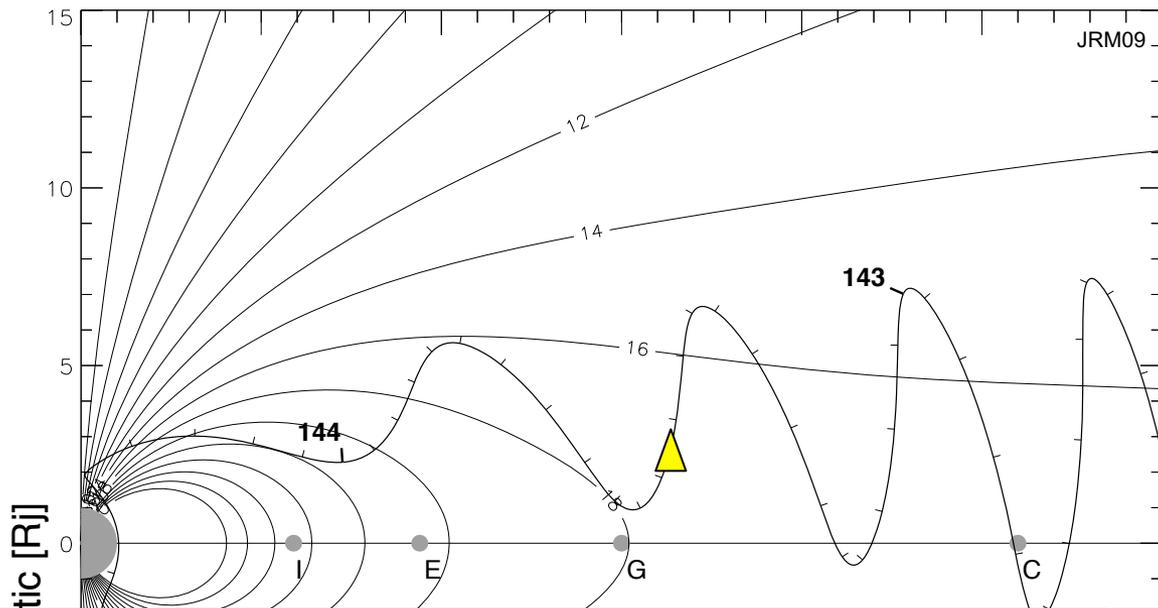


“QUIET” aurora  
probably resulting from a  
long period (days?) of low  
SW activity

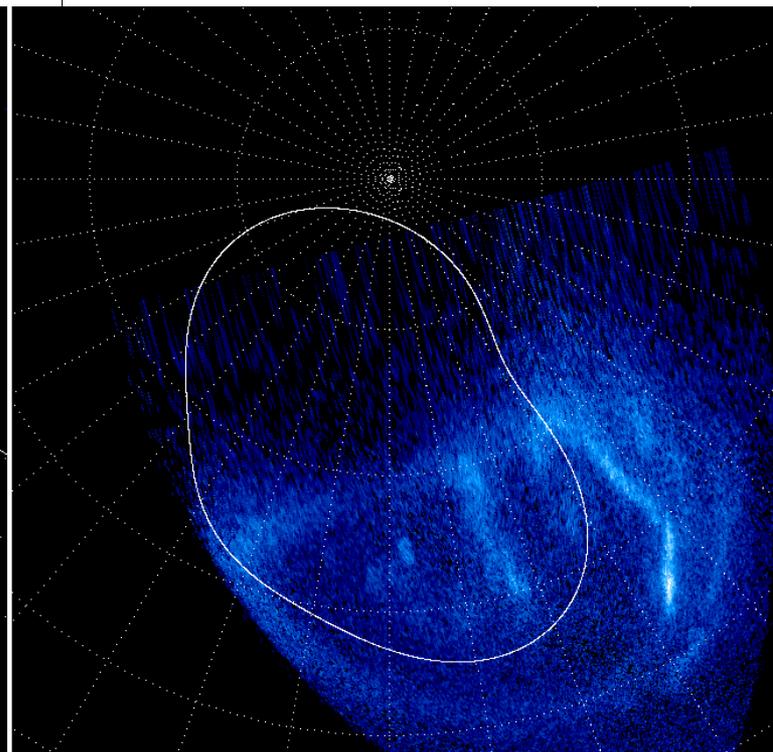
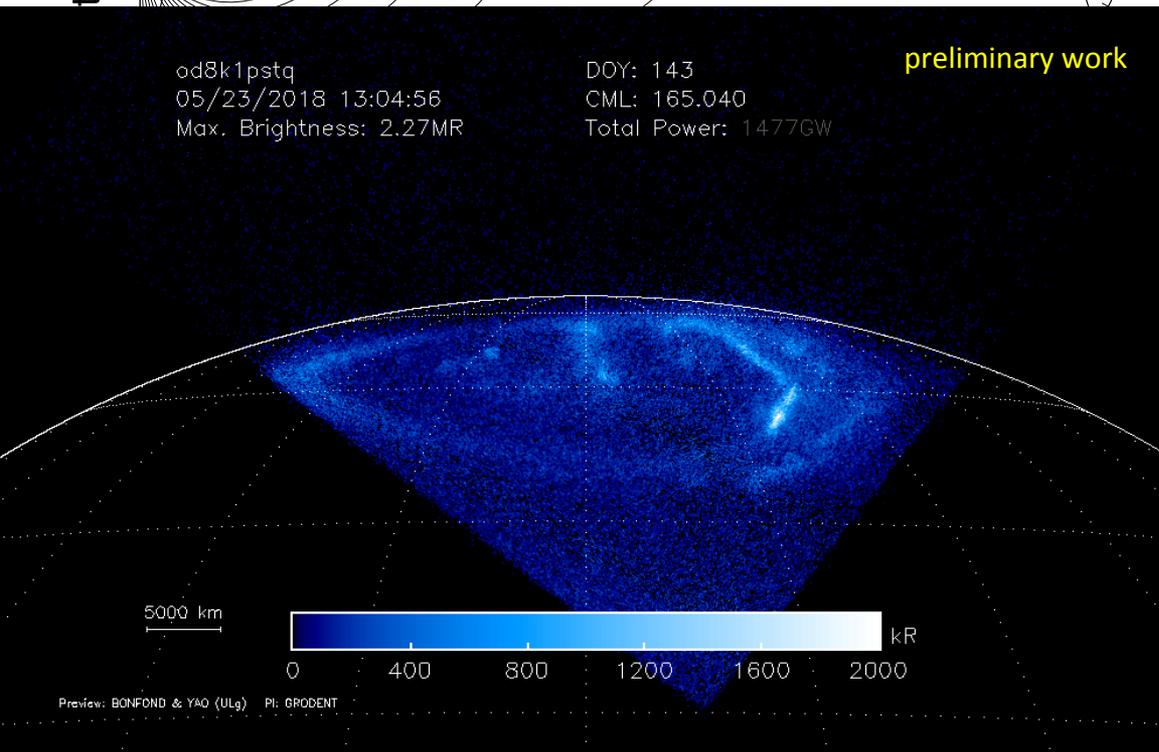




# PJ13

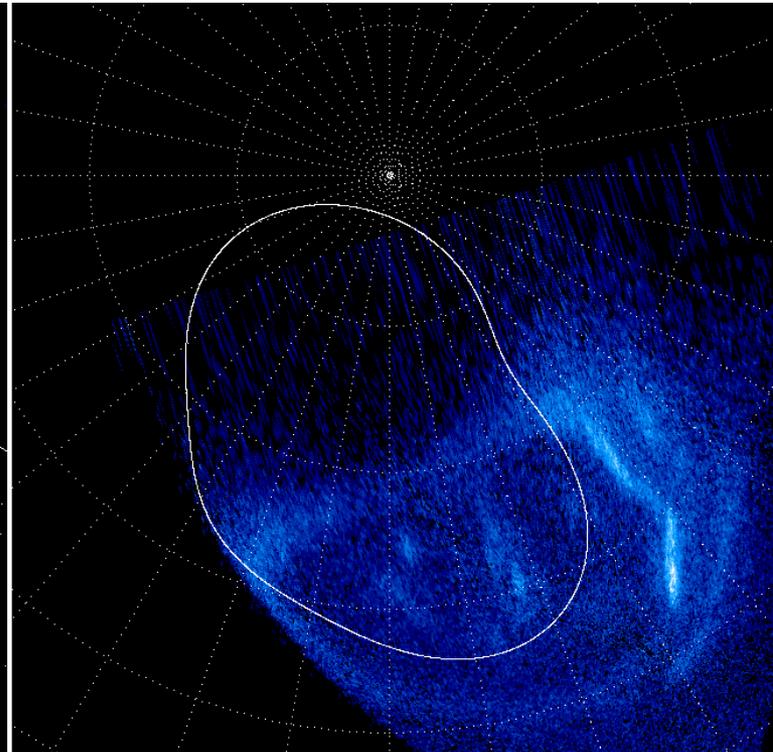
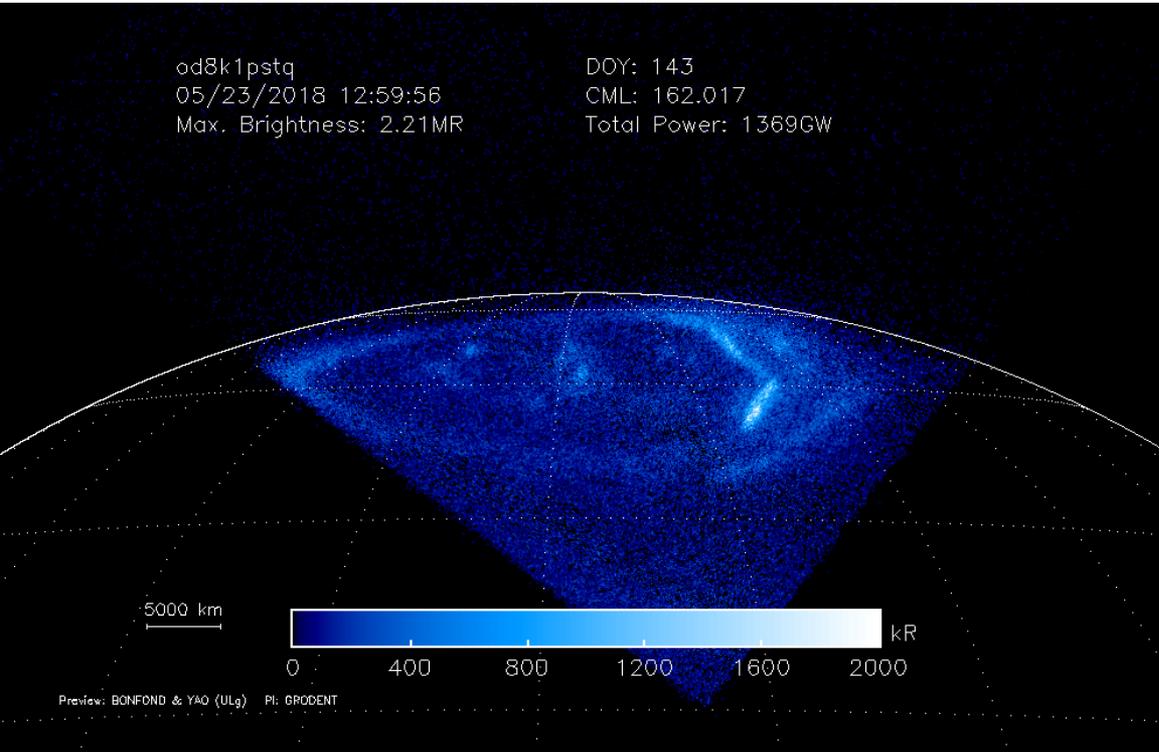


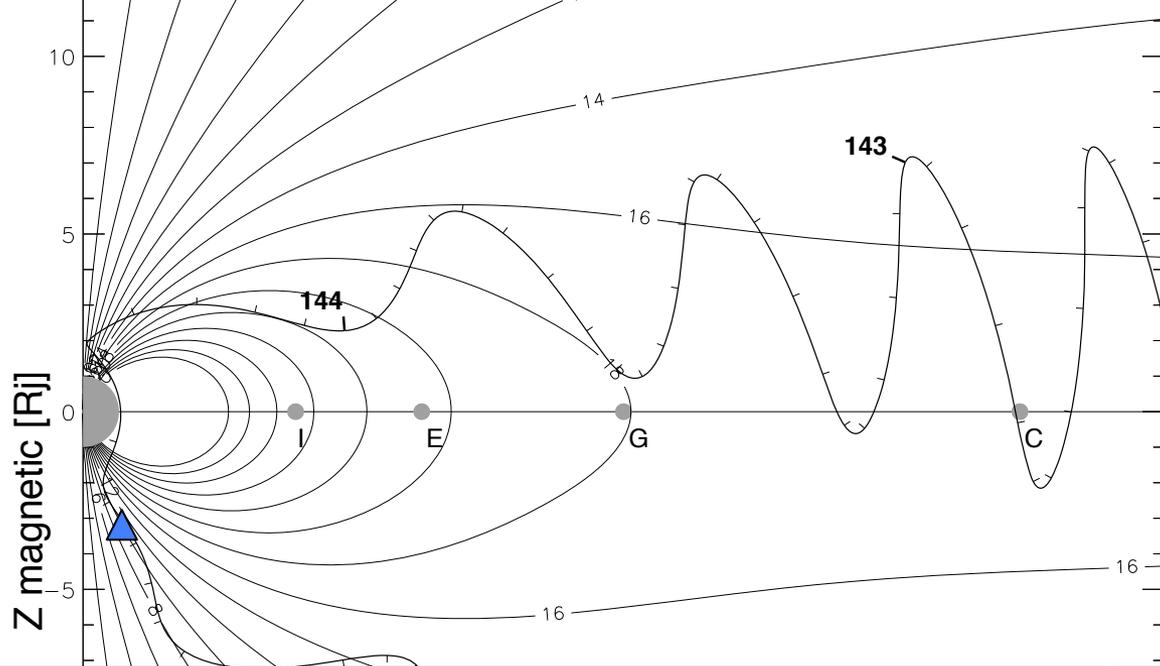
“QUIET” aurora



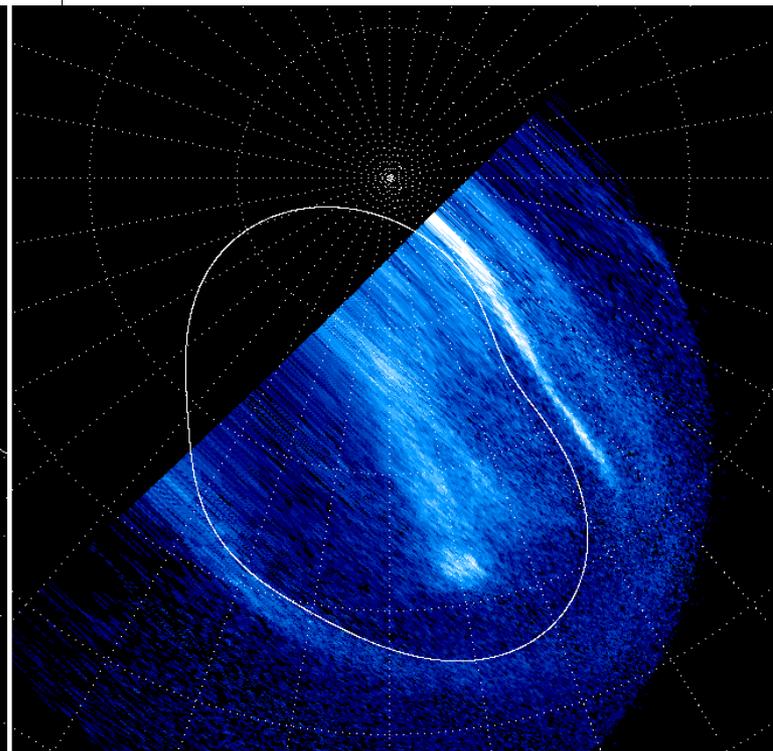
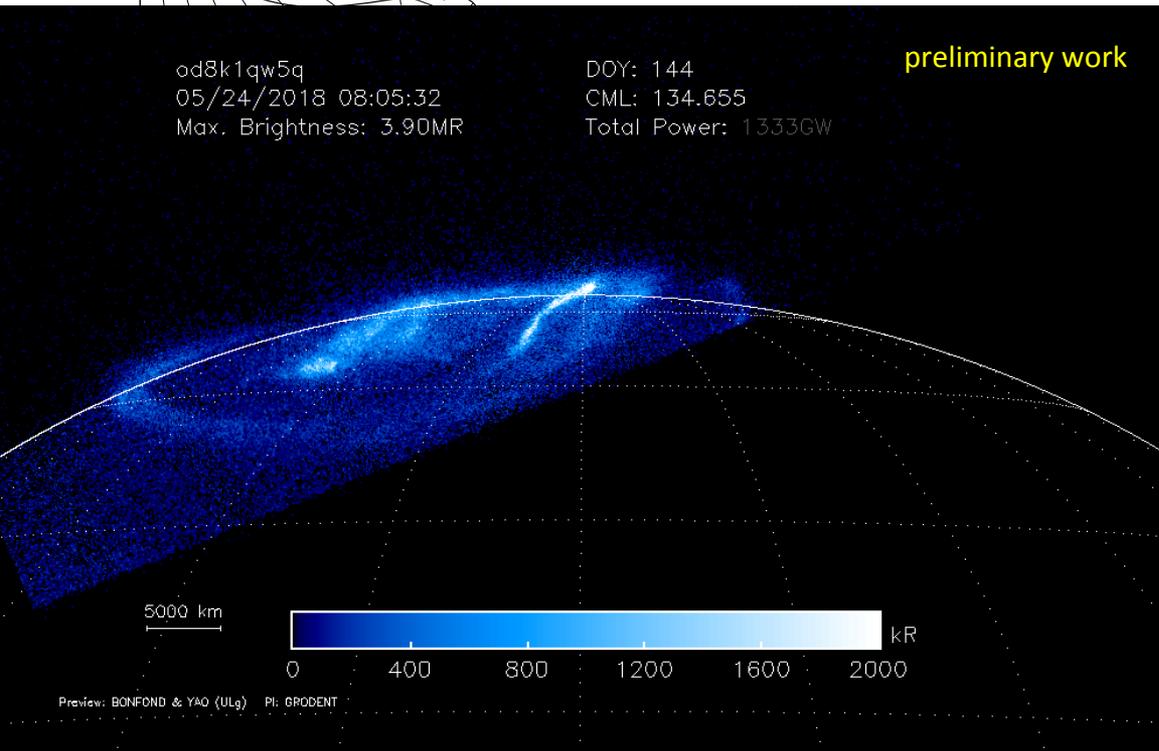
od8k1pstq  
05/23/2018 12:59:56  
Max. Brightness: 2.21MR

DOY: 143  
CML: 162.017  
Total Power: 1369GW



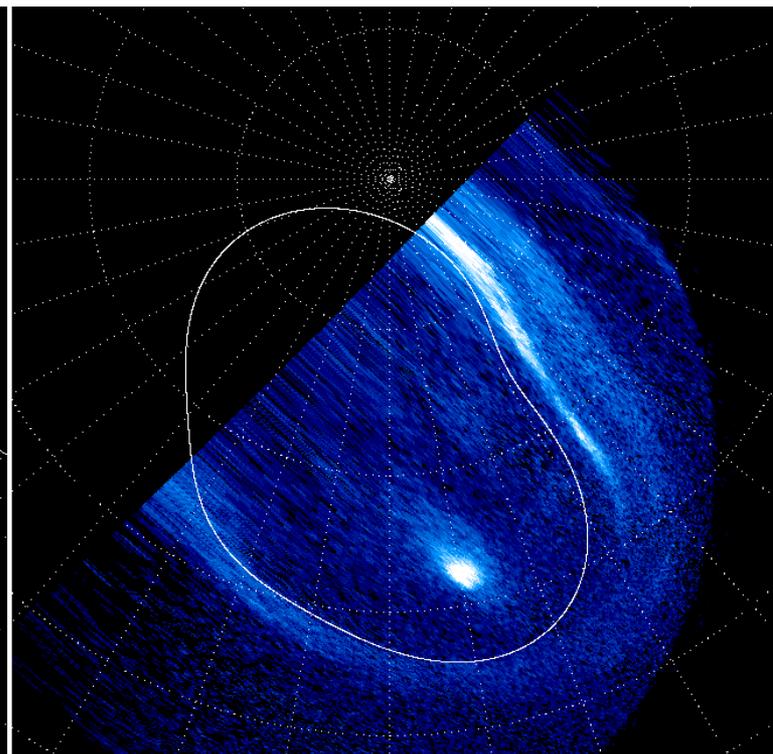
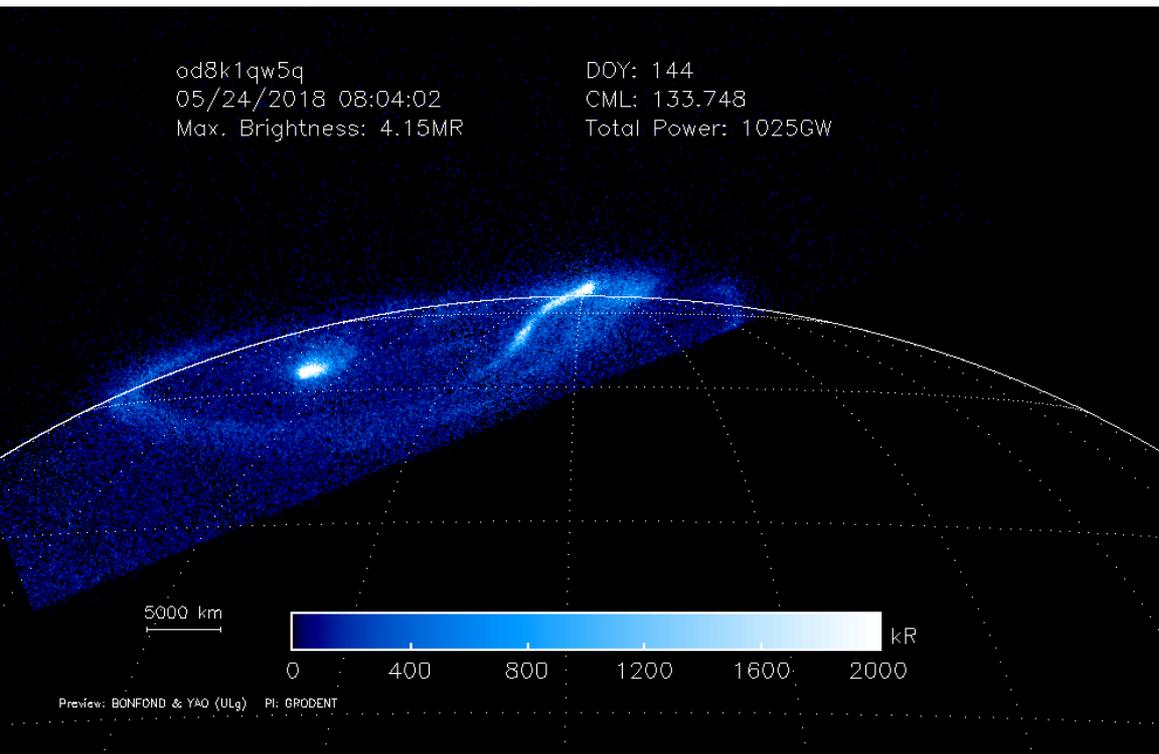


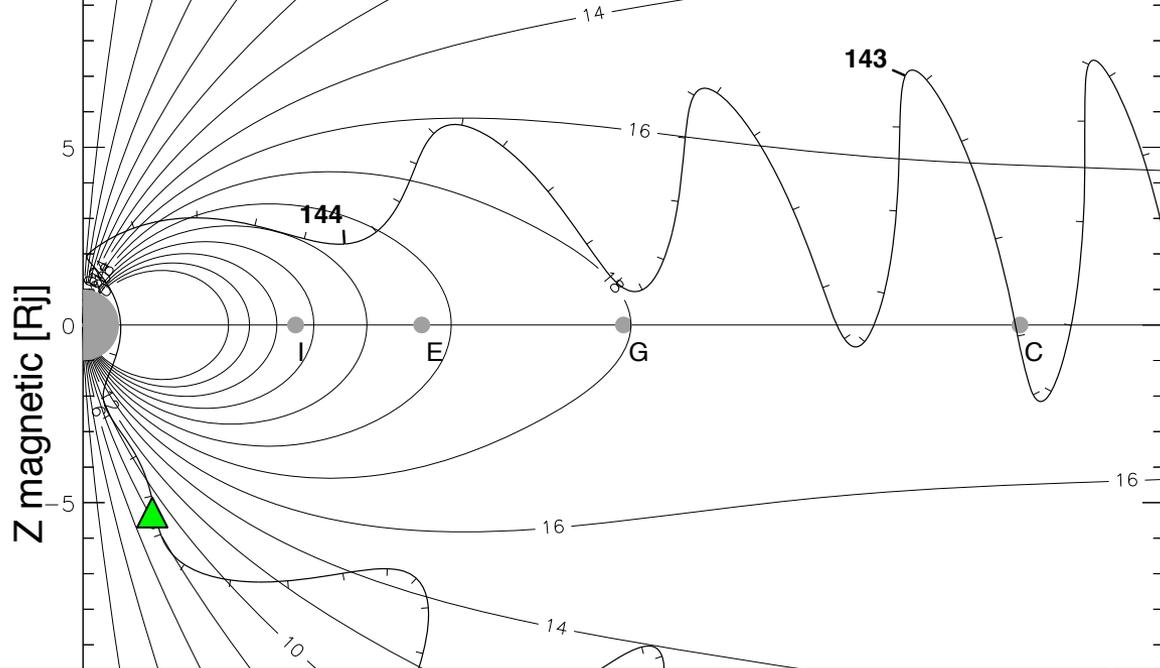
“QUIET” aurora



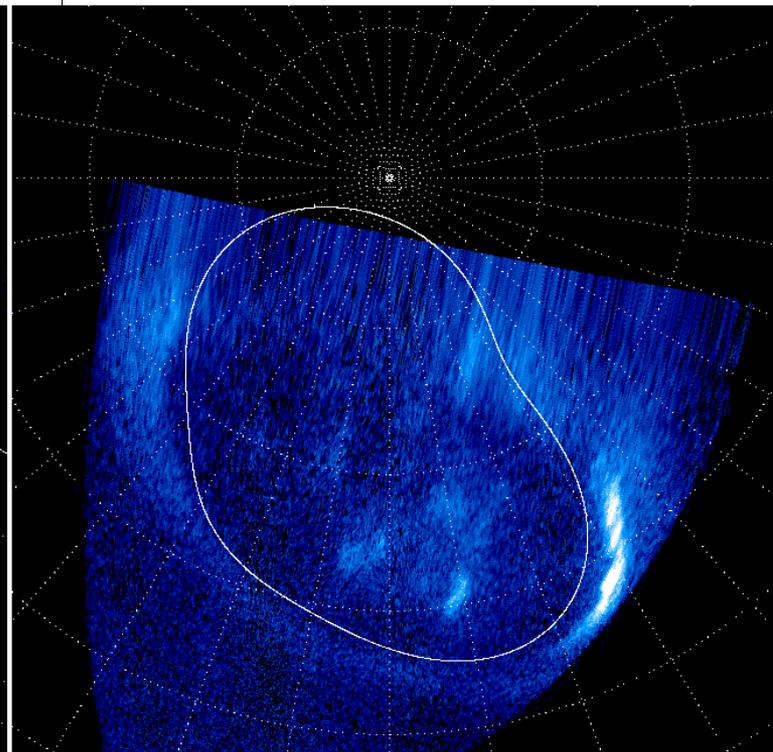
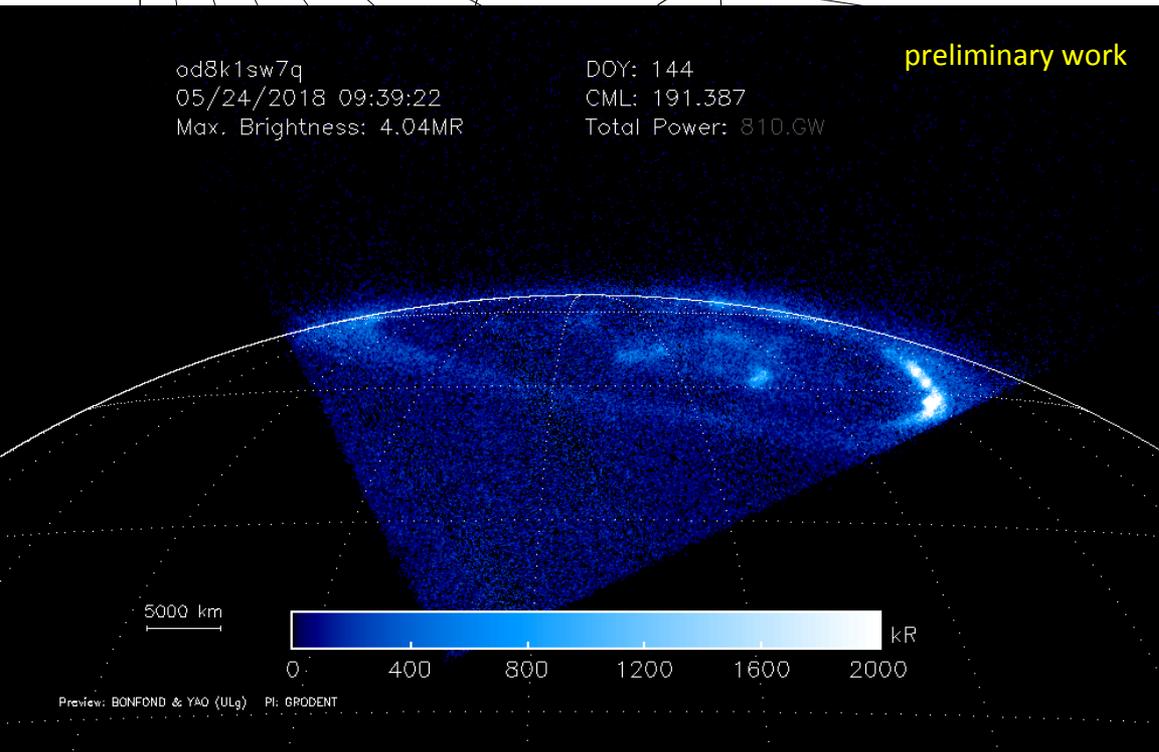
od8k1qw5q  
05/24/2018 08:04:02  
Max. Brightness: 4.15MR

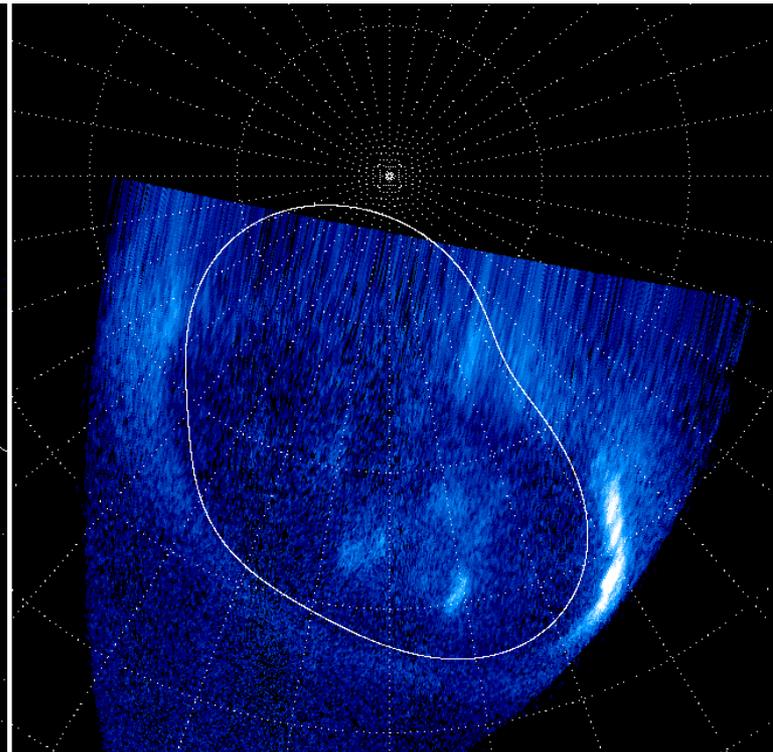
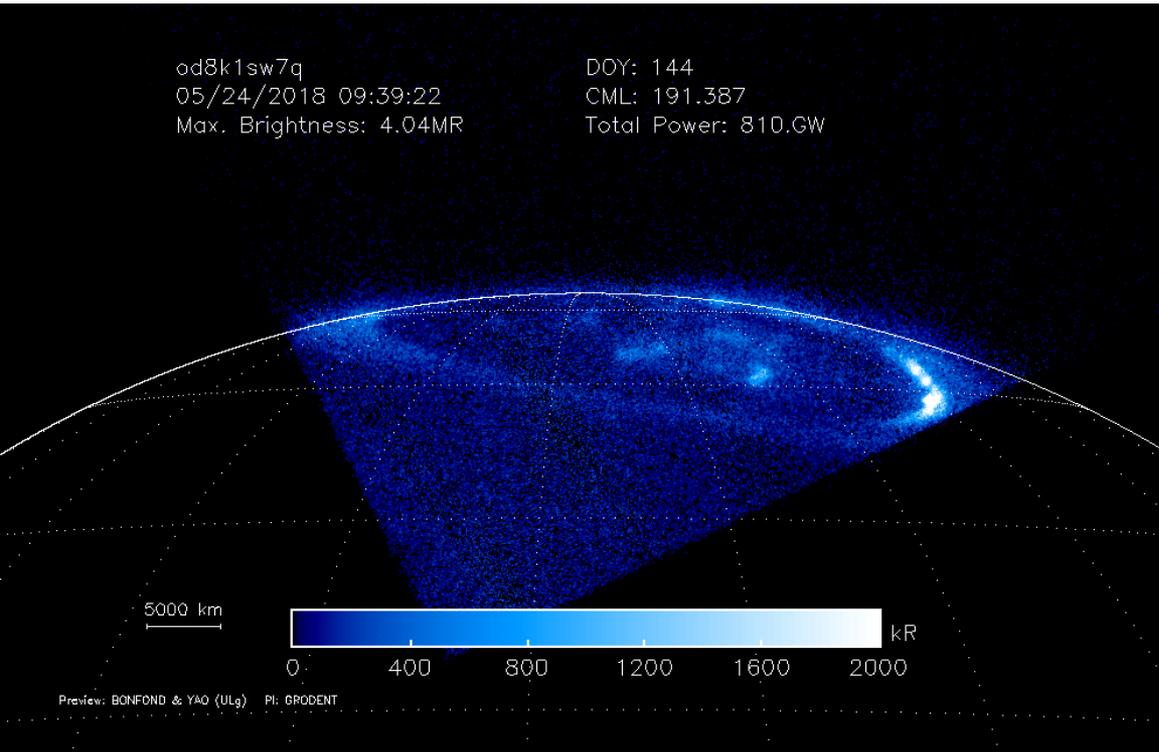
DOY: 144  
CML: 133.748  
Total Power: 1025GW

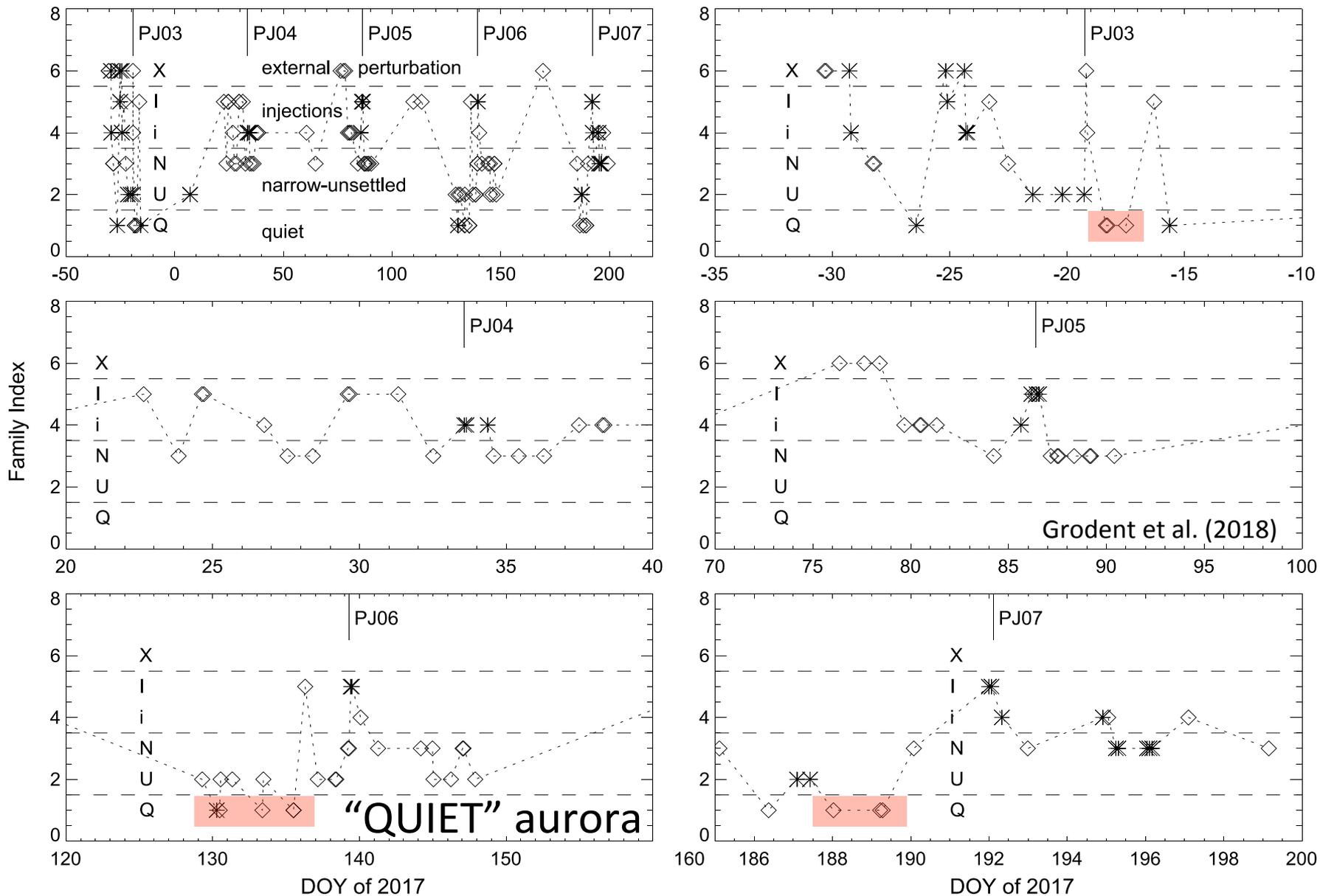




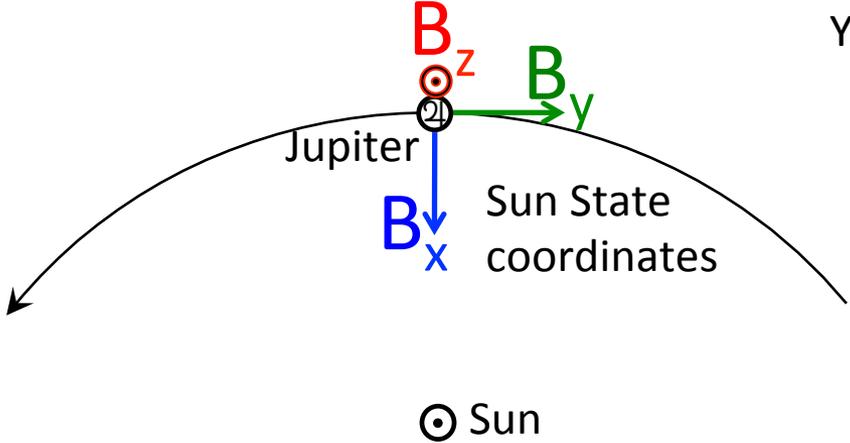
“QUIET” aurora





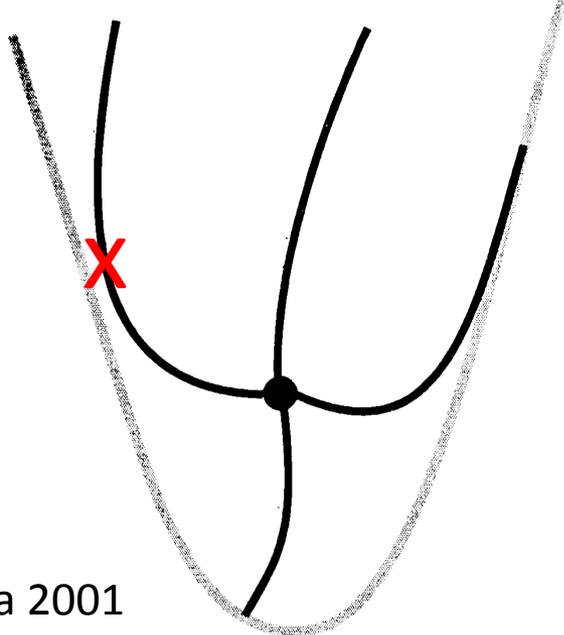


**Figure 4.** Graphical representation of the evolution of the auroral morphology over Juno orbits 3 to 7 (see also detailed parameters in Table 1). Like Figure



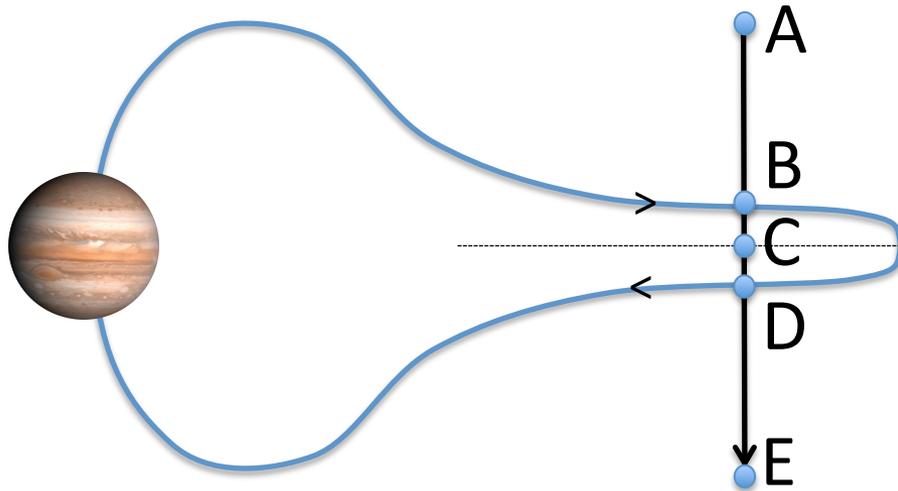
Dawn side CS

above	below	inside
$B_x < 0$	$> 0$	$= 0$
$B_y < 0$	$> 0$	$= 0$
$B_z < 0$	$< 0$	$\approx 0$

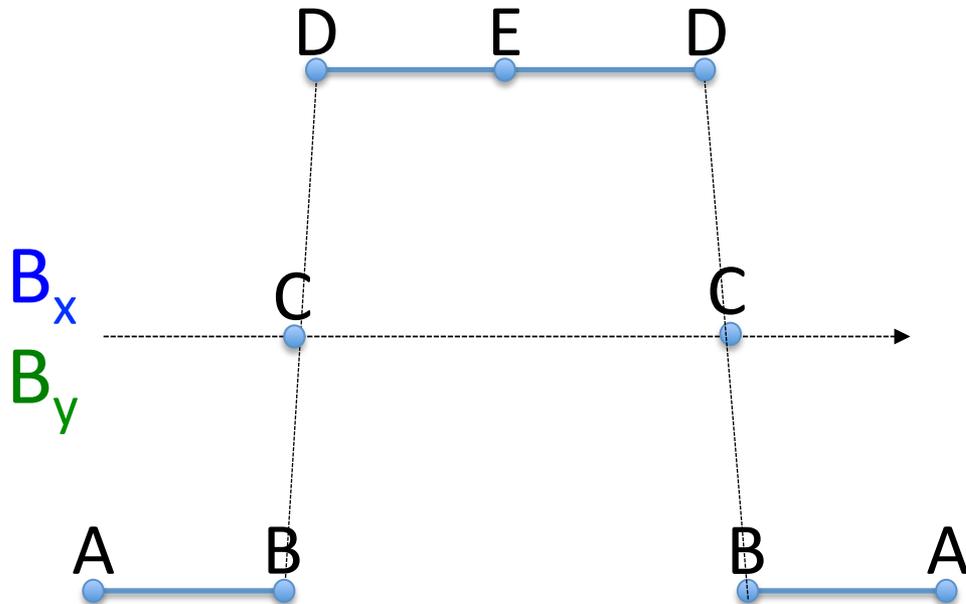


Khurana 2001

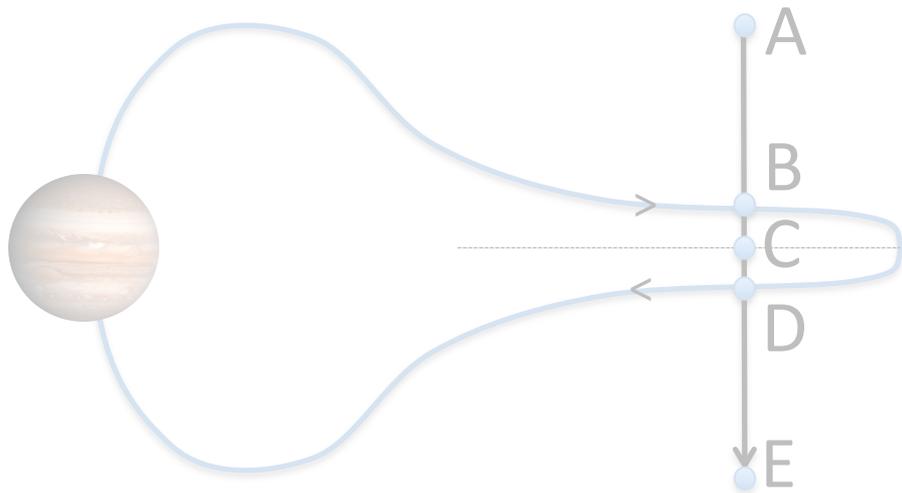
# THIN CS



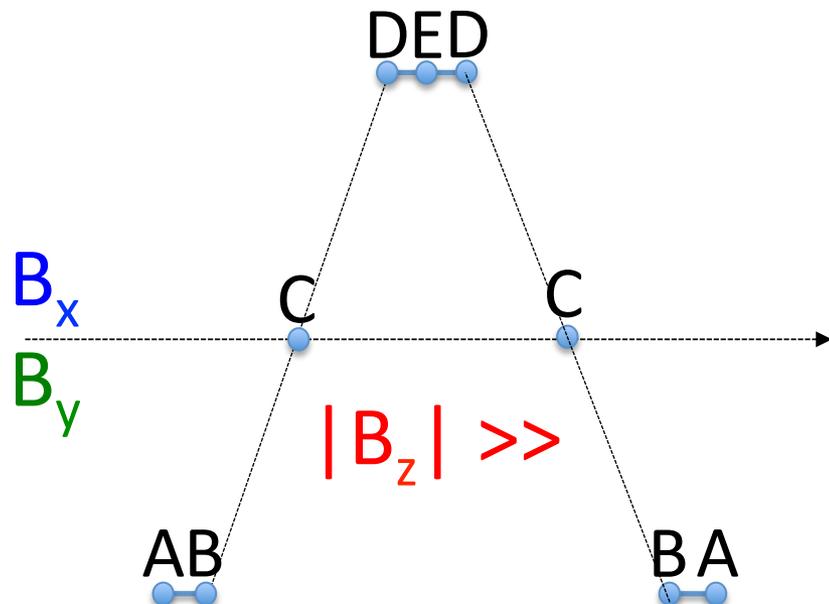
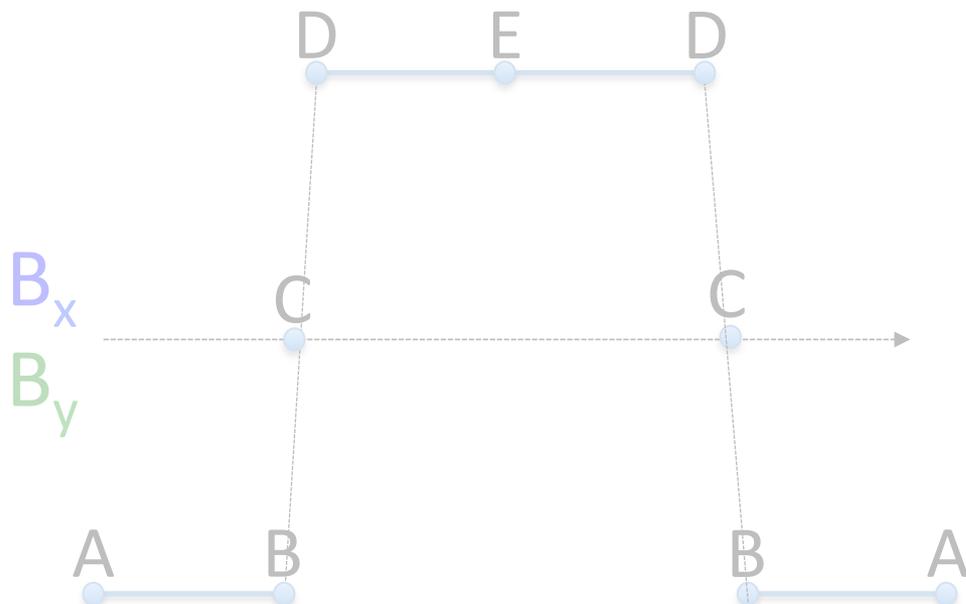
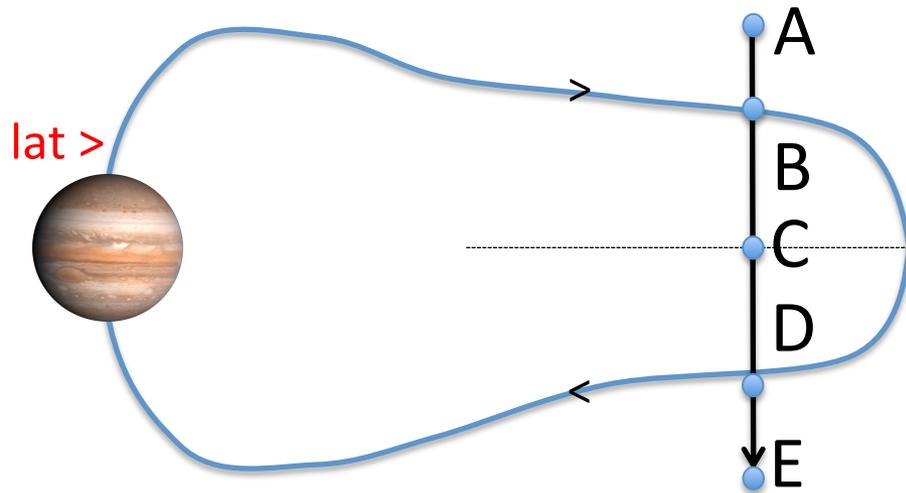
Current sheet  
flapping around  
Juno (in equatorial  
plane)

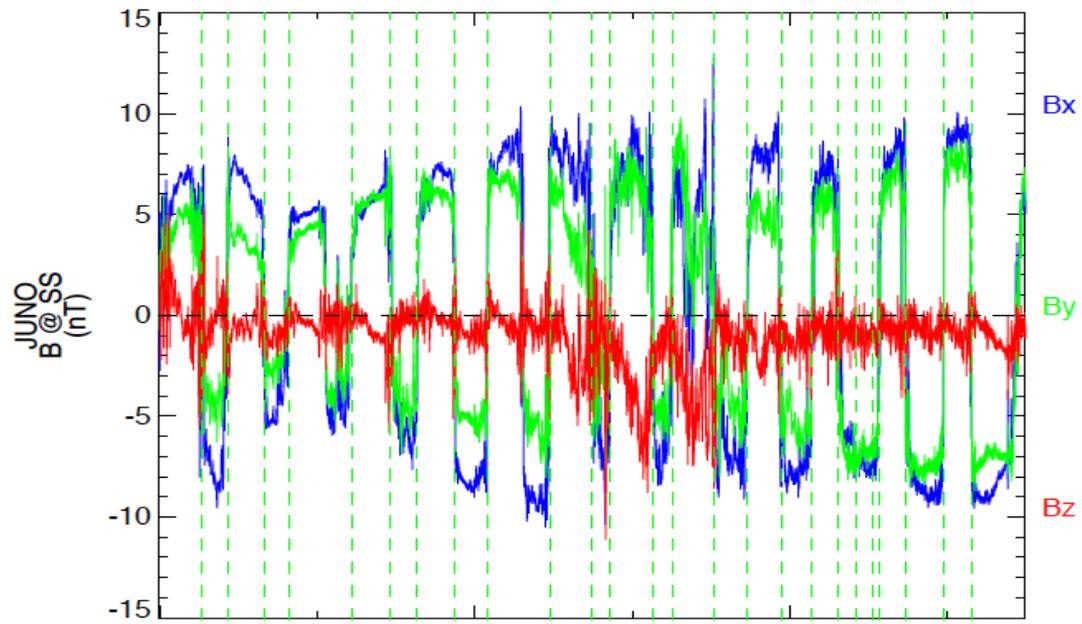


# THIN CS



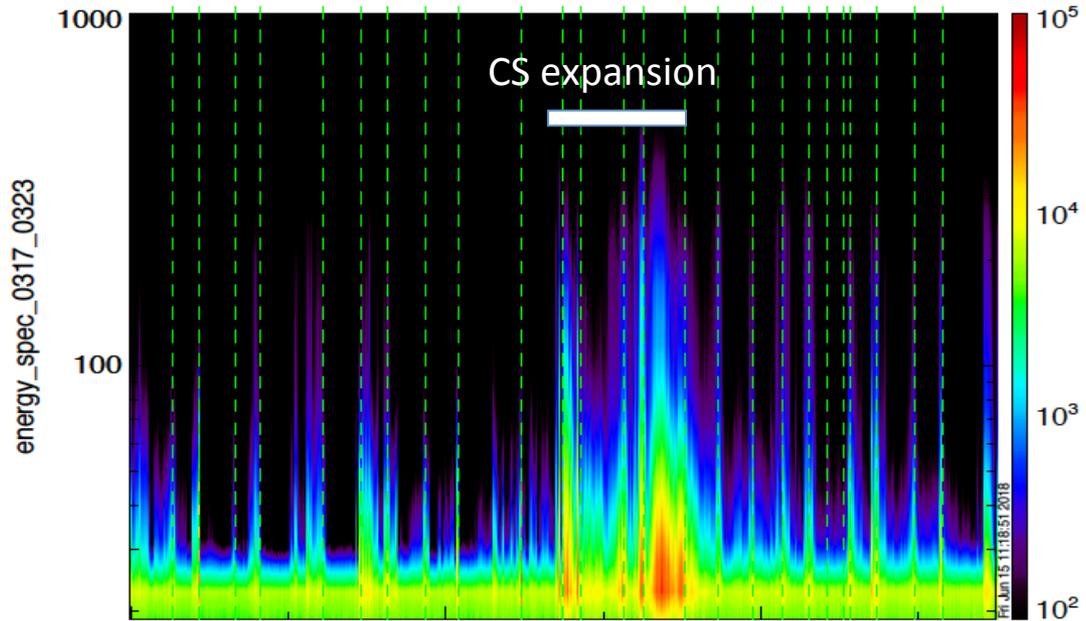
# THICK CS





**Juno orbit 05**  
PJ05 on 27 March 2017

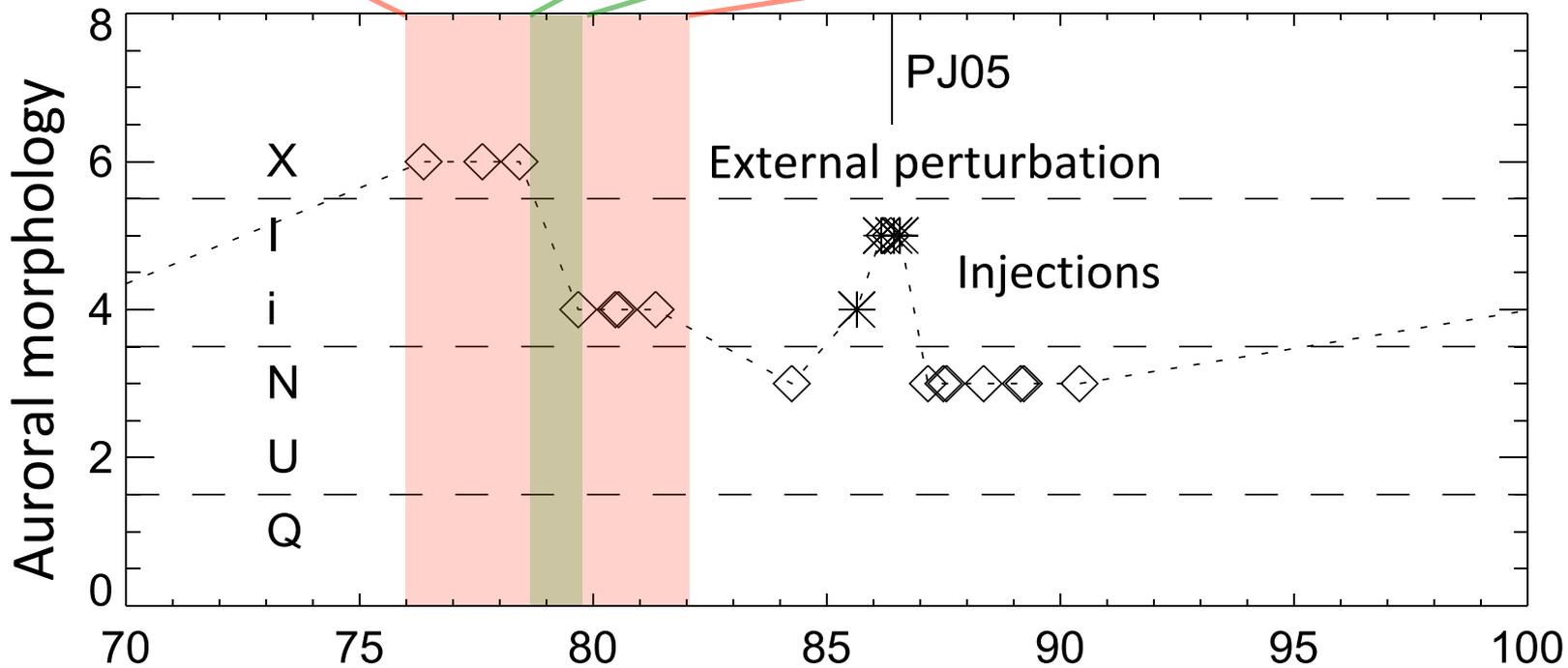
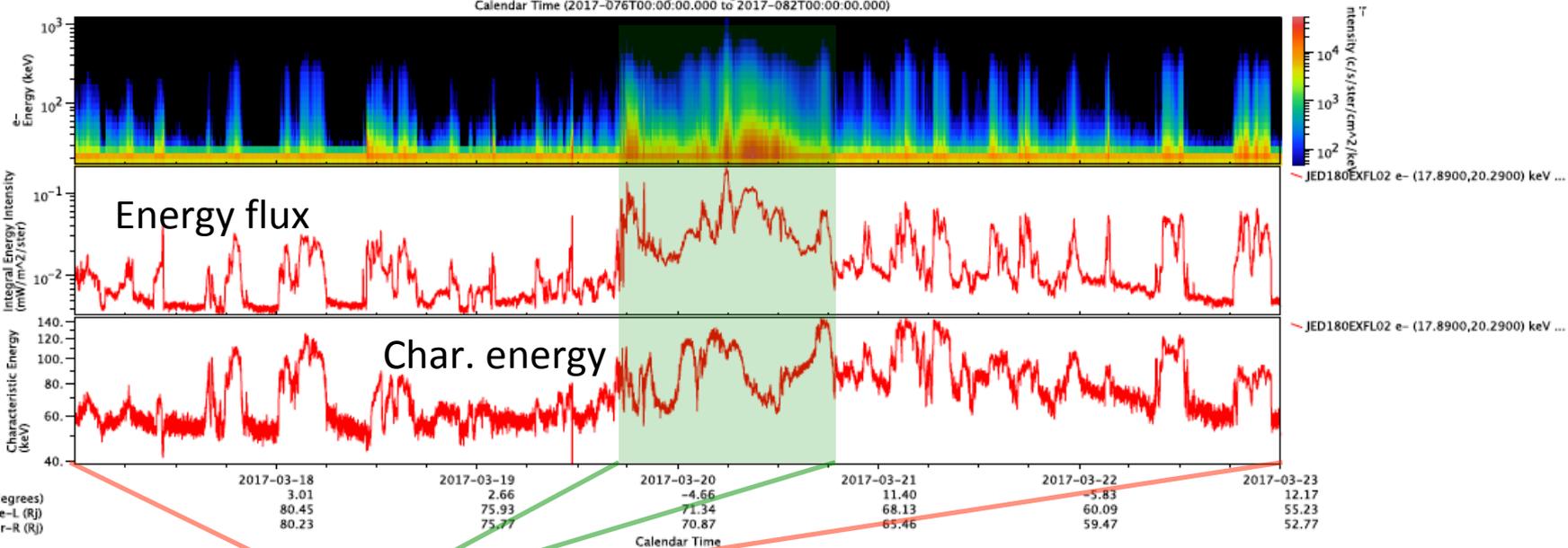
**MAG data**  
Connerney et al.



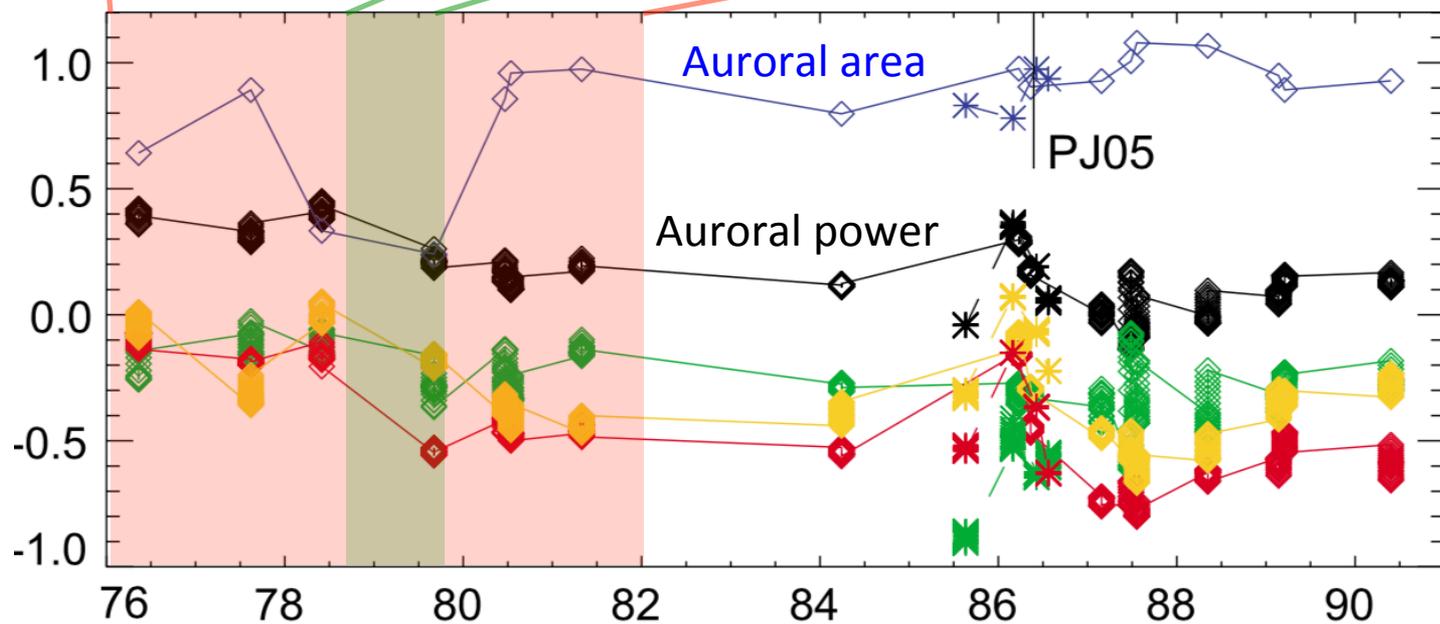
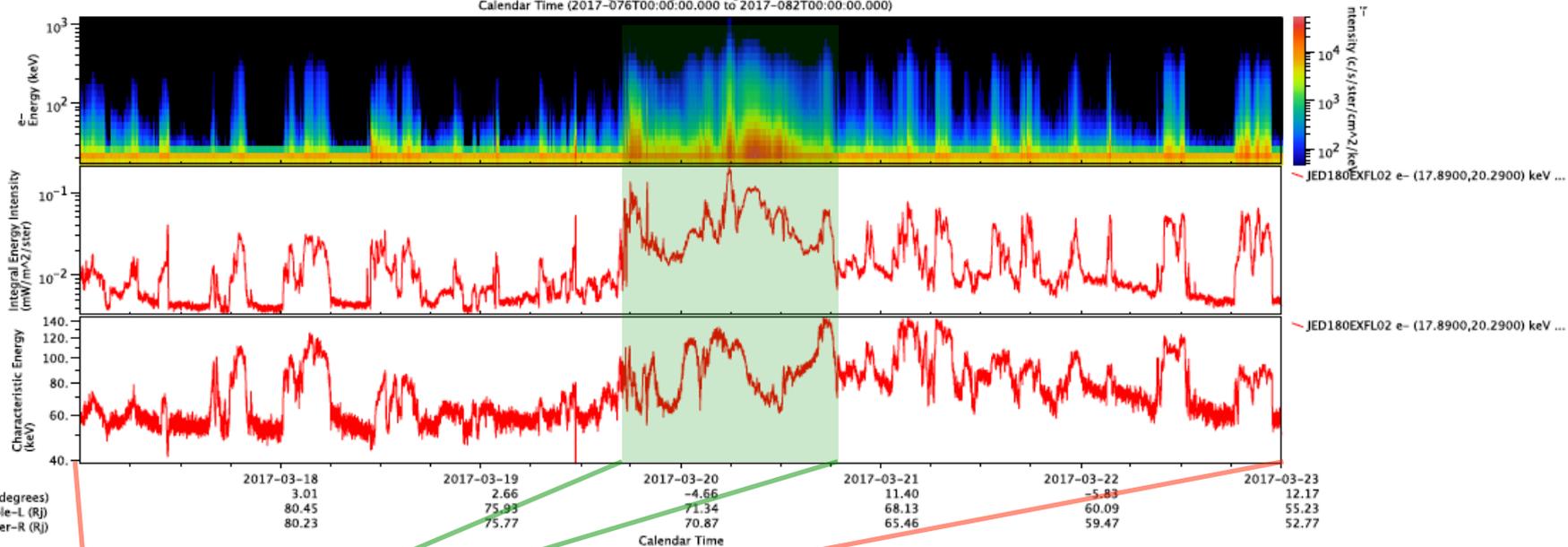
**JEDI data**  
Mauk et al.

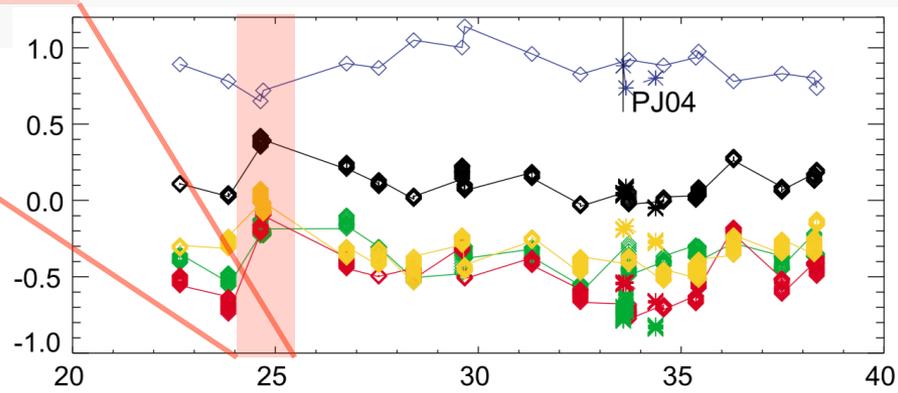
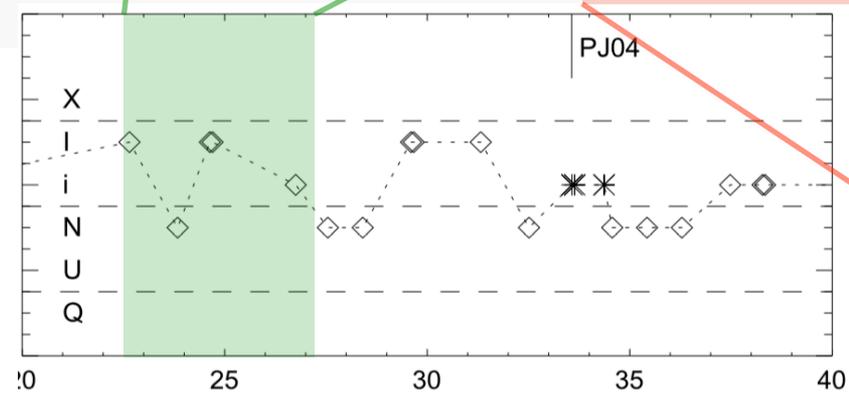
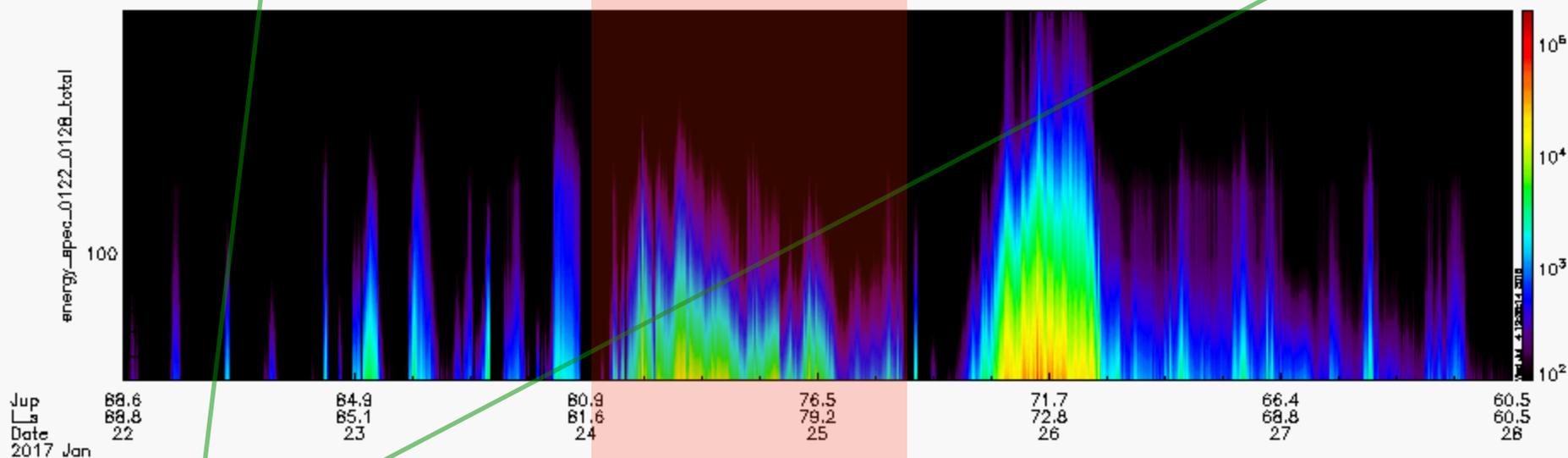
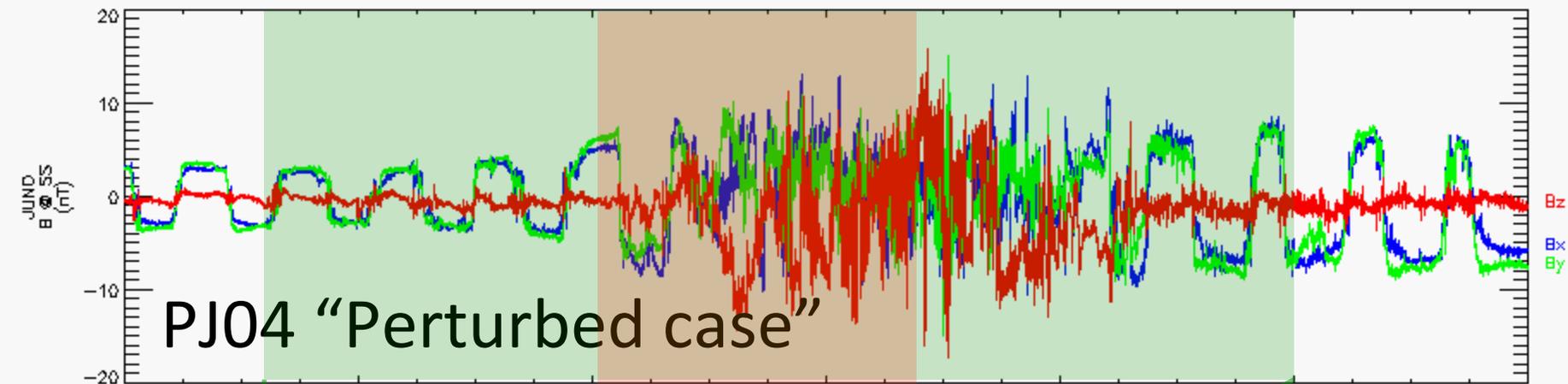
POJ05	SS	-24.7	-22.5	-19.7
POJ05	SS	-80.5	-72.3	-62.4
POJ05	SS	-4.2	-2.0	0.2
Date		17	19	21
2017 Mar				

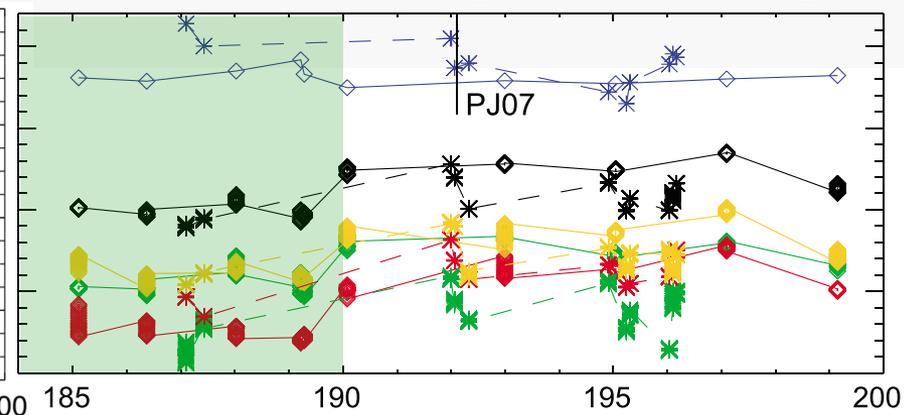
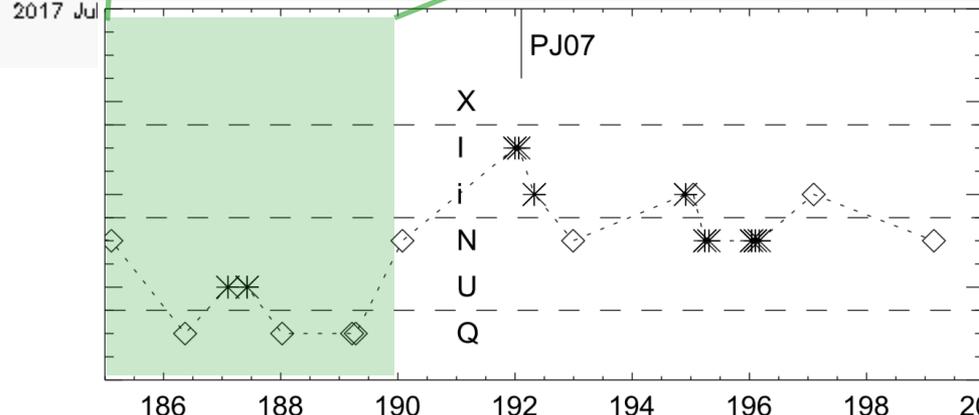
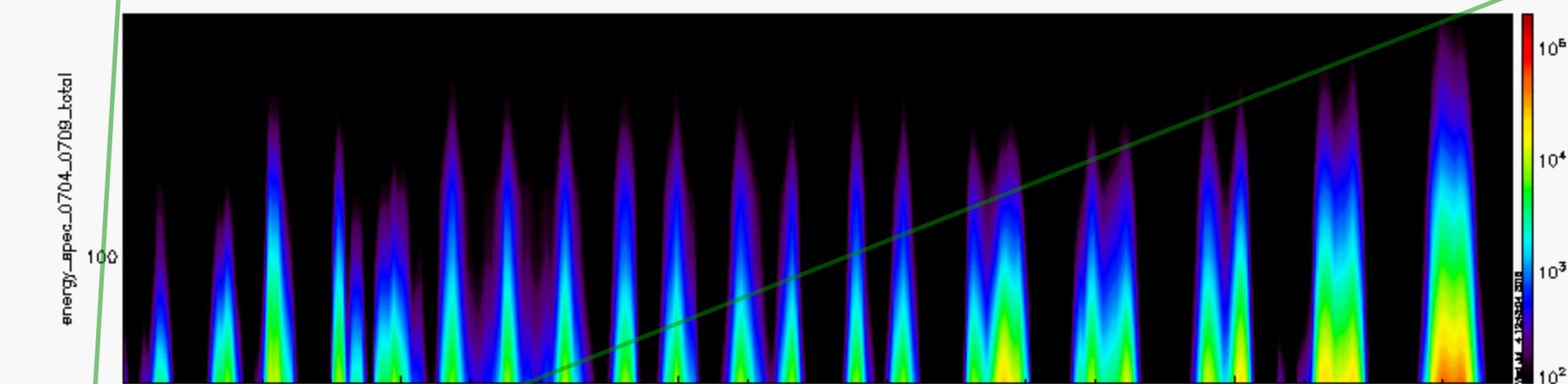
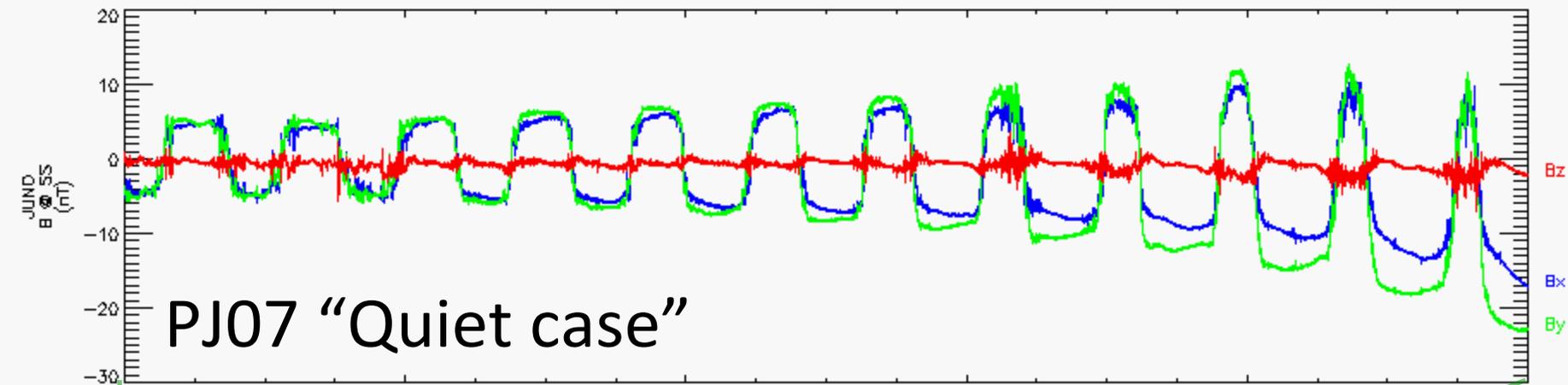
JED180EXFL02 e- (17.8900,20.2900) keV ...  
 Energy Spectrogram configuration  
 Calendar Time (2017-076T00:00:00.000 to 2017-082T00:00:00.000)



JED180EXFL02 e- (17.8900,20.2900) keV ...  
 Energy Spectrogram configuration  
 Calendar Time (2017-076T00:00:00.000 to 2017-082T00:00:00.000)







# "Conclusion" (work in progress)

1. Perturbations of B and of energetic  $\bar{e}$  [50-80R<sub>J</sub>] are highly correlated with auroral **dynamics**  $\Rightarrow$  middle to outer magnetosphere dynamics is important in driving Jovian auroral emissions.
2. Apparent anti-correlation between JEDI signatures of CS expansion and HST auroral size and bulk morphology.
3. Potential connection between CS thickening (reconnection?) and auroral injections.
4. CS expansion and plasma energisation events observed by Juno and HST are global effects.