Strongly interacting dark matter doesn't explain the DAMA signal

Maxim Laletin in collaboration with J.R. Cudell

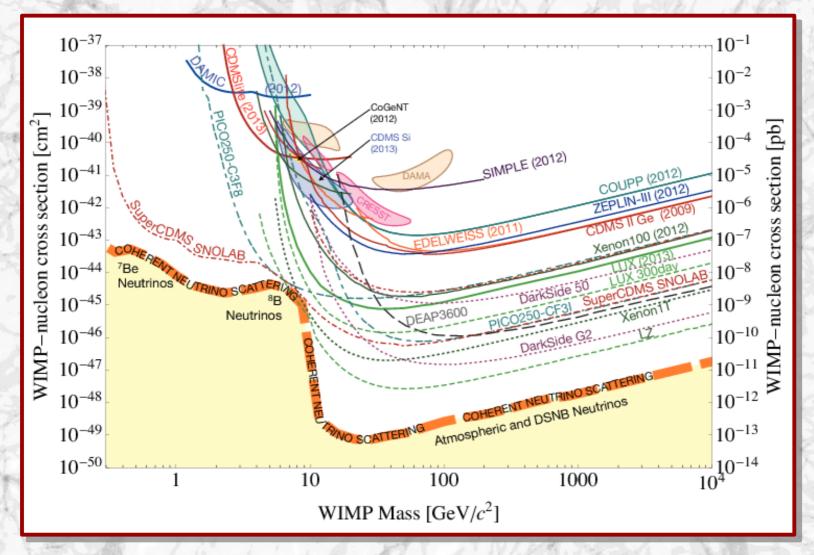
Star Institute Université de Liège



8th COSPA Meeting 10/11/2017

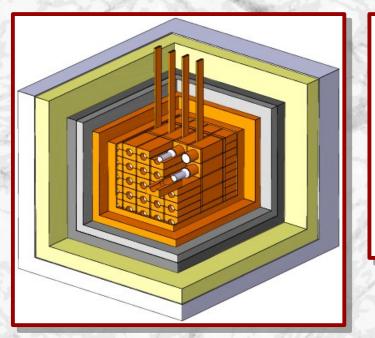


Direct detection puzzle

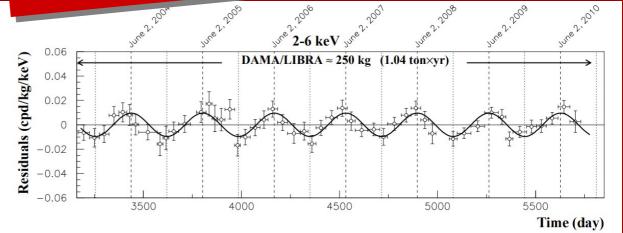


J. Cooley, Phys.Dark Univ. 4, 2014

DAMA Experiment

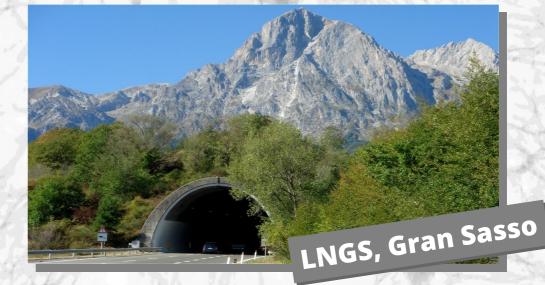


9.3 σ C.L.



Annual modulation of the signal

DAMA/LIBRA 250 kg Nal(Tl) 1.33 ton*yr ~ 10³ counts per day



Explanations

- Background neutrons
- Muon-induced neutrons
- Neutrino-induced neutrons

However, these explanations don't work

- "No role for neutrons, muons and solar neutrinos in the DAMA annual modulation results", *Bernabei et al., Eur.Phys.J., 2014*
- "Comment on "Fitting the annual modulation in DAMA with neutrons from muons and neutrinos'", *Barbeau et al., Phys.Rev.Lett., 2014*
- "Can muon-induced backgrounds explain the DAMA data?", Klinger and Kudryavtsev, J.Phys.Conf.Ser., 2016

Any room for dark matter?

Yes, but not for simple WIMPs

- Resonant Dark Matter, Yang Bai, Patrick J. Fox (Fermilab) JHEP 0911 (2009) 052
- Mirror dark matter interpretations of the DAMA, CoGeNT and CRESST-II data, R. Foot Phys. Rev. D86, 2012
- Inelastic dark matter with spin-dependent couplings to protons and large modulation fractions in DAMA, *Stefano Scopel, Kook-Hyun Yoon, JCAP 1602, 2016*

Essentially, all these models assume that DAMA detector is somehow "special", so that there is no signal in other detectors

Composite dark matter

- Significant elastic cross sections (~ 10-26 cm²)
- Tiny recoil energies (slow down before the detector)
- Produce the signal due to inelastic processes on some special component of DAMA
- Composite Dark Matter and Puzzles of Dark Matter Searches, M. Khlopov, A. Mayorov, E. Soldatov, Int.J.Mod.Phys. D19 (2010)
- Milli-interacting dark matter, Q. Wallemacq, Phys. Rev. D 88 (2013)
- Dark antiatoms explain DAMA, J.R. Cudell, Q. Wallemacq, JCAP 2015

Slowdown in the ground

DM looses energy after multiple collisions on terrestrial nuclei (mainly Si and O) and acquires the surrounding temperature (thermalizes)

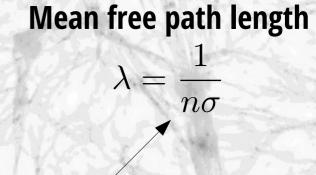
The recoil energy on nuclei is tiny

$$E_{\rm rec} = \left(\frac{M v_{\rm th}^2}{2}\right) \frac{2mM}{(m+M)^2} \approx 1\,{\rm meV}$$

Thermalization depth

$$l_{\rm th} = \frac{M\lambda}{m} \log\left(\frac{v_{\rm in}}{v_{\rm th}}\right)$$

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Number density of the medium (rock)

N

35°

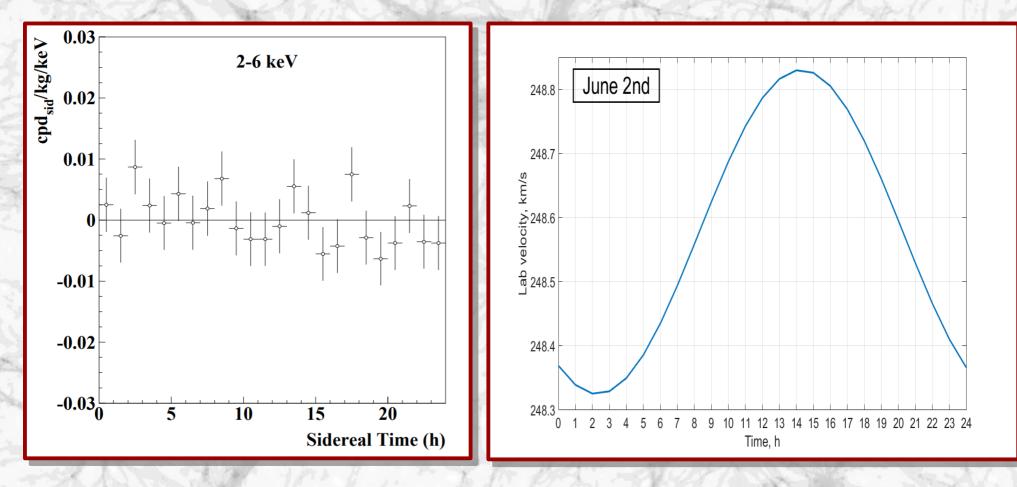
DM preferential direction at 08:00 GMST

DM preferential direction at 20:00 GMST

LNGS

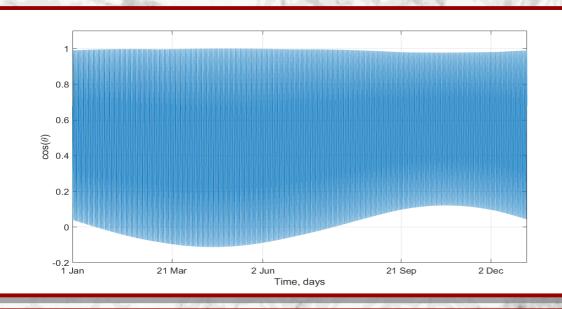
42°

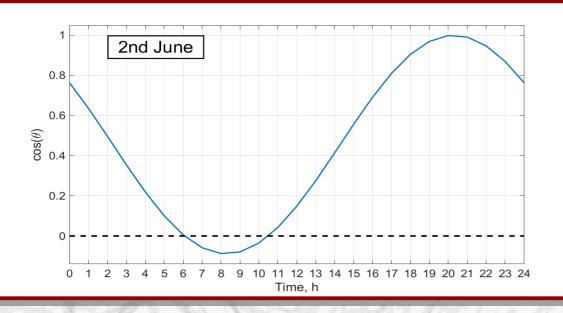
One expects diurnal modulation for SIMPs

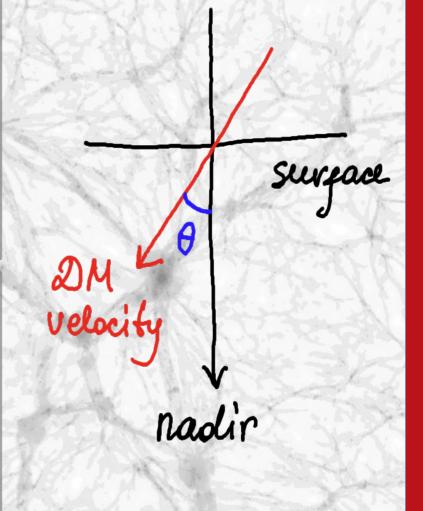


DAMA data

Expected for WIMPs







Propagation in the ground

The propagation can be described as diffusion with a drift due to gravity

$$\frac{\partial N(\vec{x},t)}{\partial t} = D\Delta N(\vec{x},t) - v_d \frac{\partial N(\vec{x},t)}{\partial z} + f(\vec{x},t)$$
Particle number
density
Diffusion coefficient
$$D = \sqrt{\frac{\pi kT}{8M}} \frac{m+M}{mn\sigma}$$
Drift velocity
(related to the diffusion coefficient
via Einstein relation)
MaD

 $v_d = \frac{MgD}{kT}$

"Flat-Earth" approximation

DM

DAMA



All the points on the surface are equal, so the only dimension that matters is Z

1D diffusion equation

Slow diffusion only takes place **below** the surface Above the surface the diffusion coefficient is 0

Density in the detector

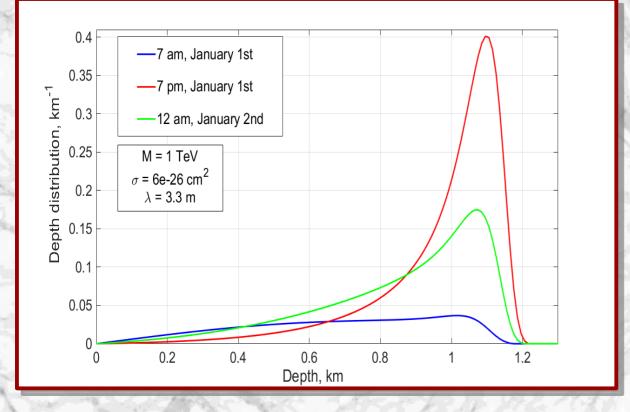
General solution of the diffusion equation in the unbounded space

$$N(t) = \frac{1}{\sqrt{4\pi D}} \int_{0}^{t} \frac{d\tau}{\sqrt{t-\tau}} \iiint_{-\infty}^{\infty} d\vec{l} \, \omega_{l}(\vec{l},\tau) \, f(\vec{l},\tau) \, G(\vec{l},\tau,t)$$
with the Green's function
$$G(\vec{l},\tau,t) = \exp\left(\frac{v_{d}(z_{det}-l_{z})}{2D} - \frac{v_{d}^{2}t}{4D}\right) \left[\exp\left(-\frac{(z_{det}-l_{z})^{2}}{4Dt}\right) - \exp\left(-\frac{(z_{det}+l_{z})^{2}}{4Dt}\right)\right]$$
Drift factor
$$f(\vec{l},\tau) = \frac{\rho_{\text{loc}} \cdot 1\%}{M} v(\vec{l}) \cos(\theta)$$
"Mirror" sink term describes the leakage of particles into the air

Velocity distribution

Dark matter thermal velocity distribution

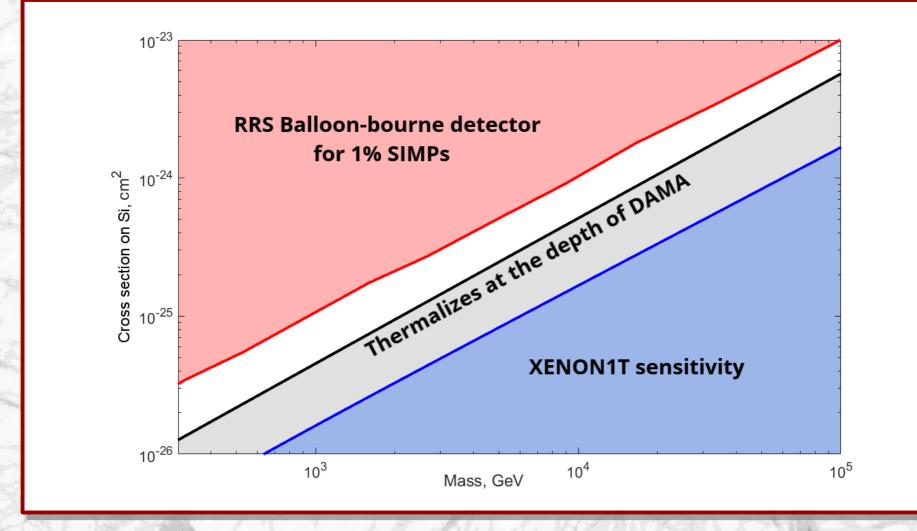
$$\omega_{v}\left(\vec{v}\right) = N \exp\left(\frac{\left(\vec{v} - v_{\text{lab}}\right)^{2}}{v_{\text{o}}^{2}}\right) \Theta\left(v_{\text{esc}} - |\vec{v} - v_{\text{lab}}|\right) \quad \text{the}$$



Can be translated into the depth distribution

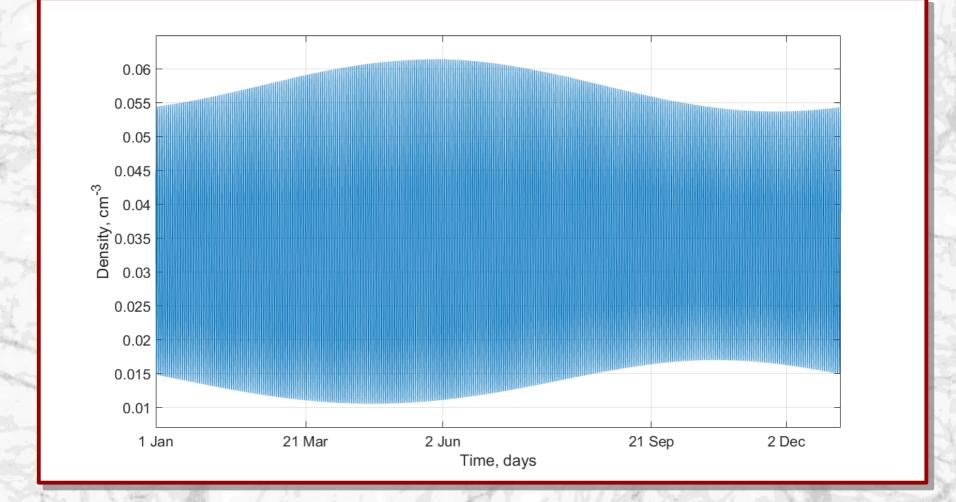
 $\omega_l(\vec{l}) = \frac{mv}{M\lambda l^2} \omega_v(\vec{l})$

Constraints

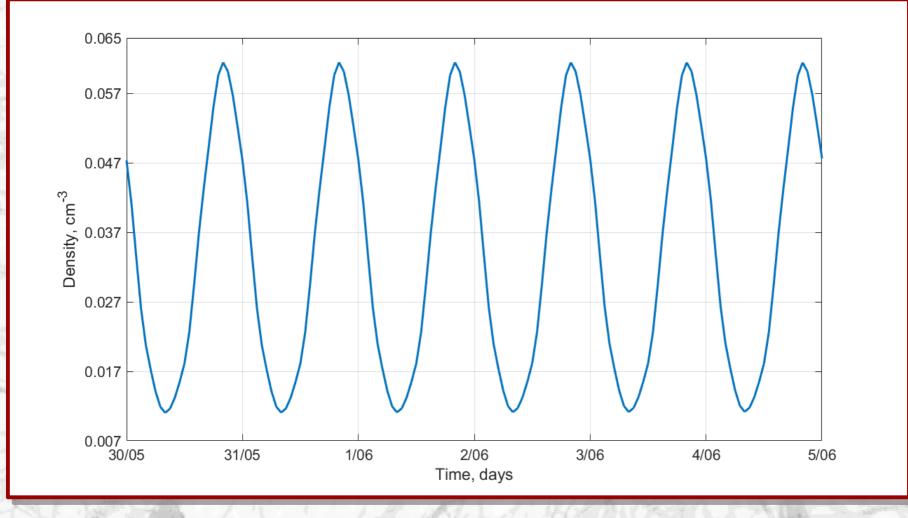


Let's study a particular set of parameters, M = 1 TeV $\sigma = 10^{-26} \text{ cm}^2$

Density inside the detector



Density inside the detector



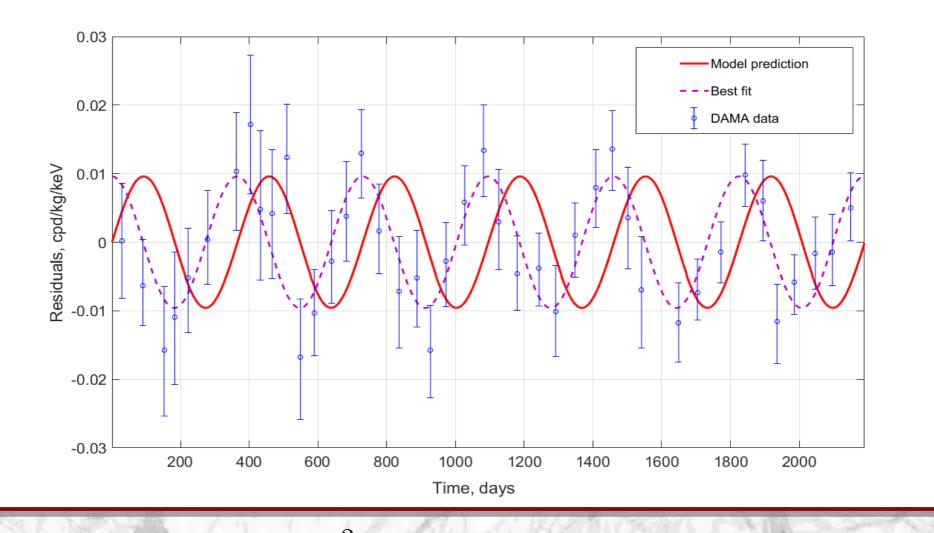
Reconstructing the signal

Inelastic cross section additional parameter that we use to normalize the signal

$$\Gamma(t) = \frac{N(t)\sigma_{\rm in}}{m_{\rm NaI}} \left(\frac{n_{\rm Tl}}{n_{\rm NaI}}\right) \sqrt{\frac{\pi kT}{8M}}$$
~ 10⁻³

The rate is than averaged over every day and the yearly average of the signal (background) is subtracted

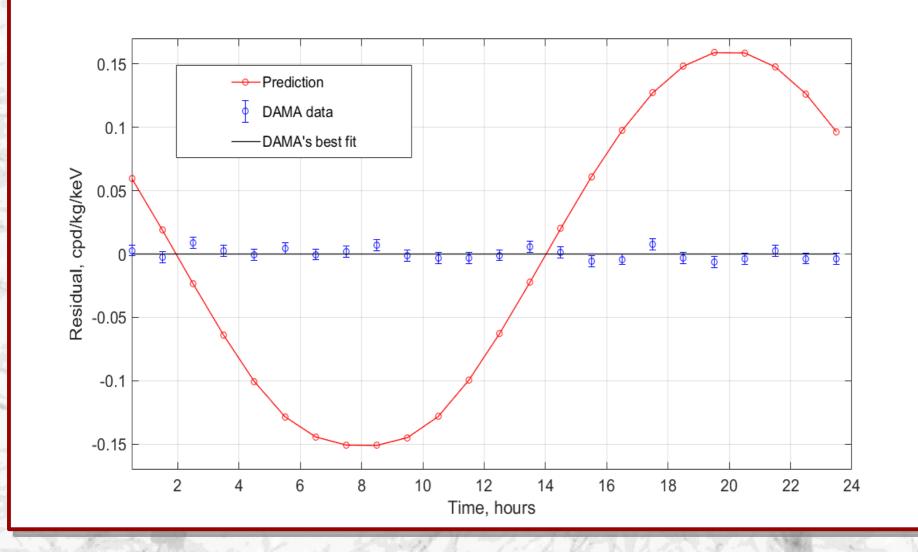
Annual modulation



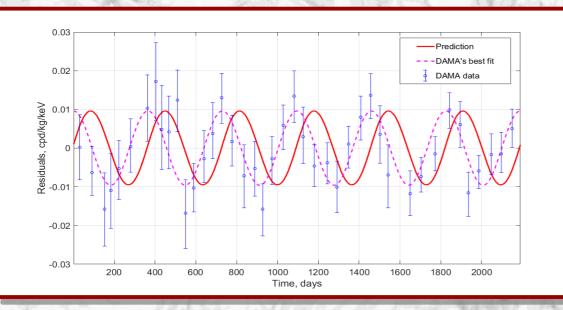
 $\frac{\chi^2}{\rm d.o.f} = 3.38$

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~ 100 days shift



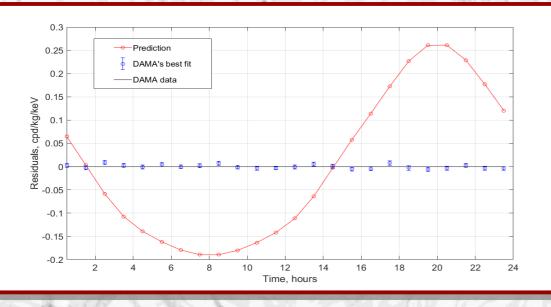
Other parameter values



M = 100 GeV $\sigma = 10^{-26} \text{ cm}^2$

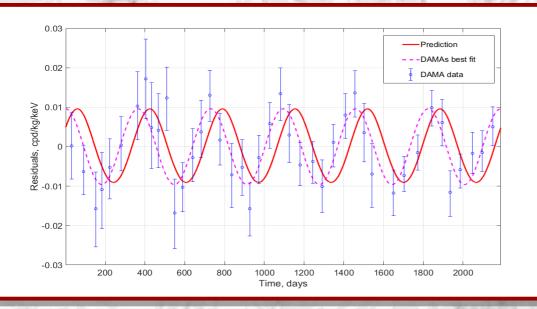
~ 90 days shift

Annual modulation



Diurnal modulation

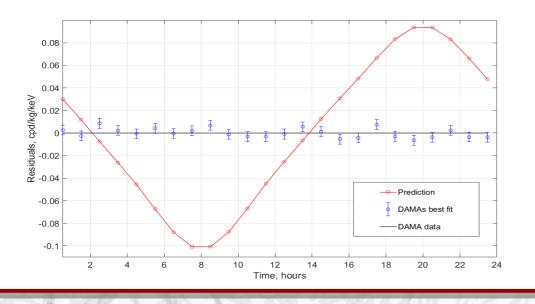
Other parameter values



M = 100 TeV σ = 6 10⁻²⁴ cm²

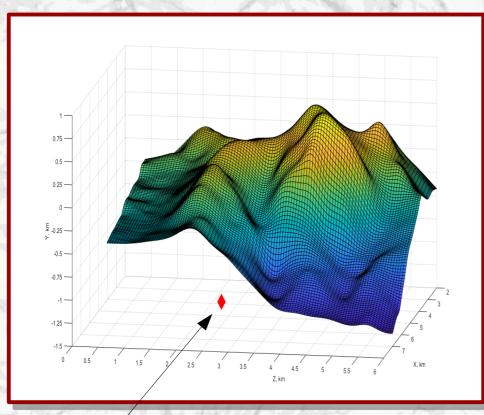
~ 60 days shift

Annual modulation



Diurnal modulation

Realistic surface



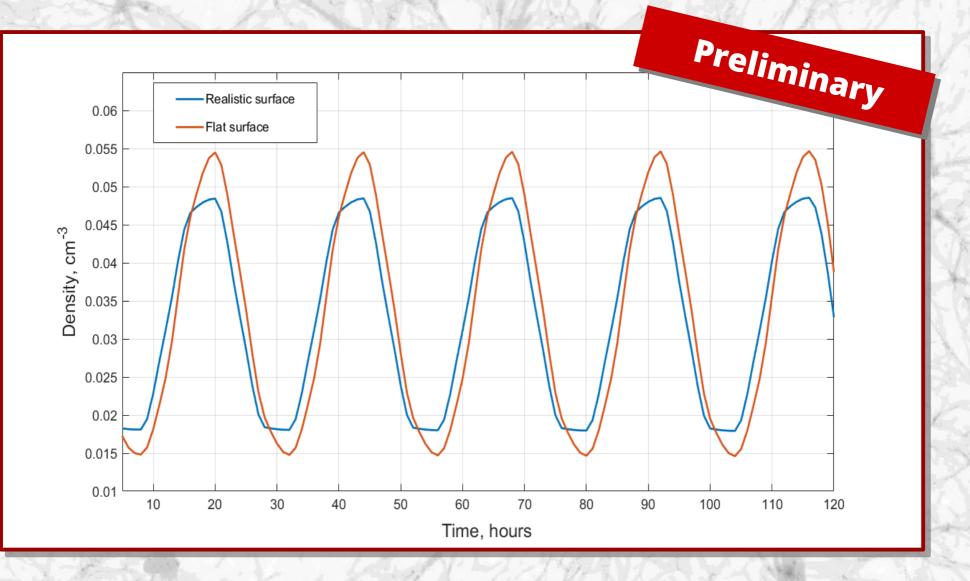


DAMA is here

Gran Sasso mountain (Google Earth)

A more detailed picture includes a complex realistic surface

Realistic surface



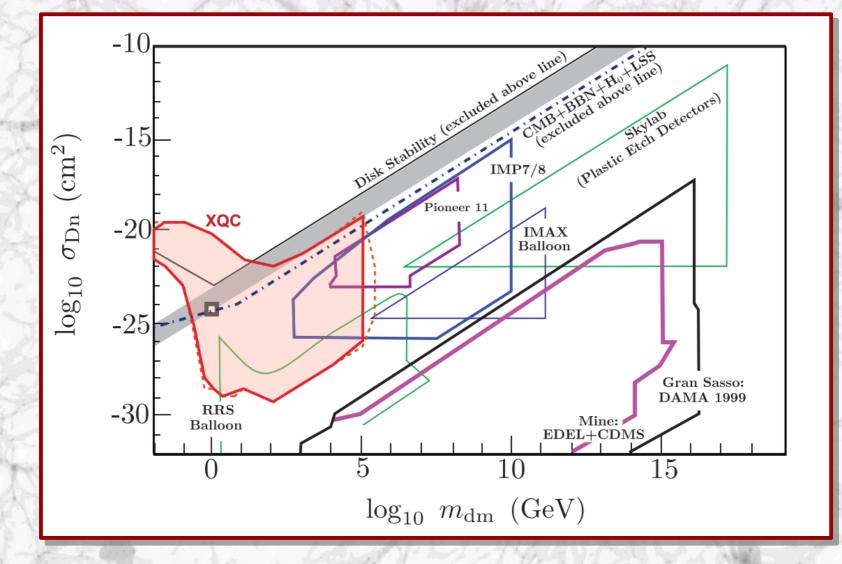
Conclusions

- The effects of propagation <u>rule out</u> some dark matter models explaining the DAMA results with strongly interacting massive particles;
- <u>Both</u> annual and diurnal modulations cannot be reproduced;
- The effects of "bouncing" particles needs to be studied, however a significant improvement of the fit is not expected;
- DAMA results will be <u>tested</u> in a series of experiments: ANAIS, COSINE-100, SABRE, COSINUS and so on.

Thank you for attention!

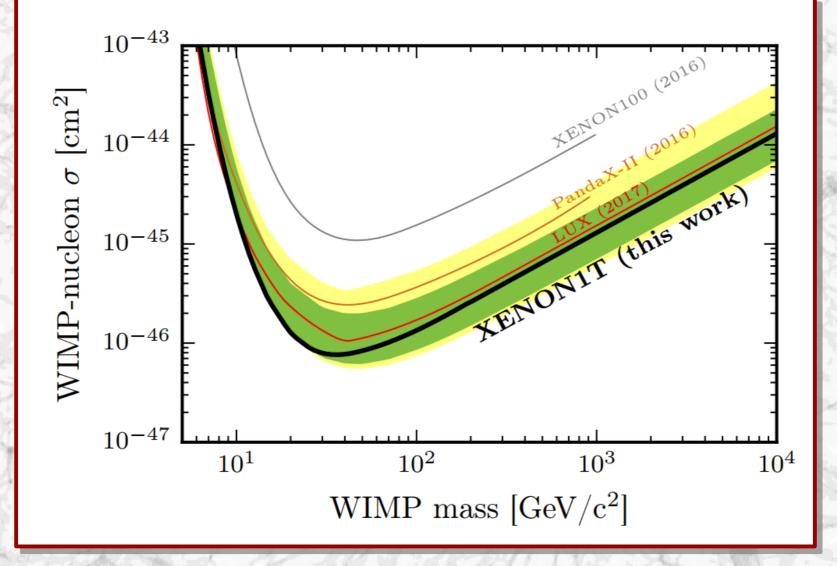
Backup

Constraints on SIMPs

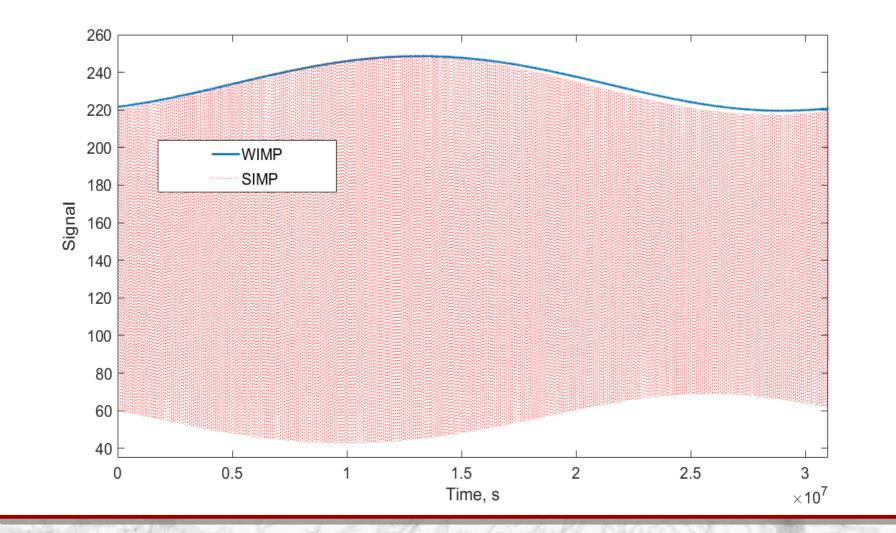


Erickcek+, Phys. Rev. D 2007, 0704.0794

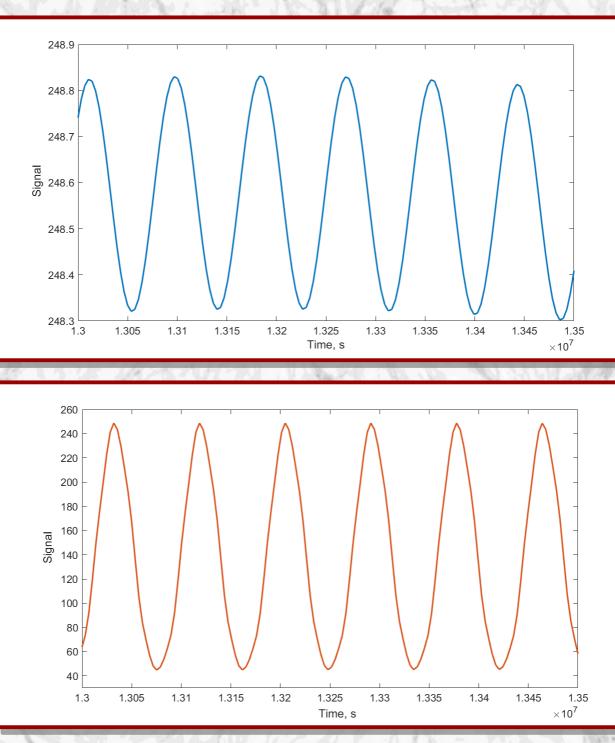
XENON1T (2017)



Signal comparison



WIMP Signal around annual maximum



SIMP Signal around annual maximum

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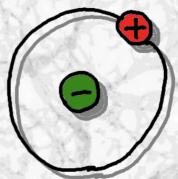
Signal reconstruction

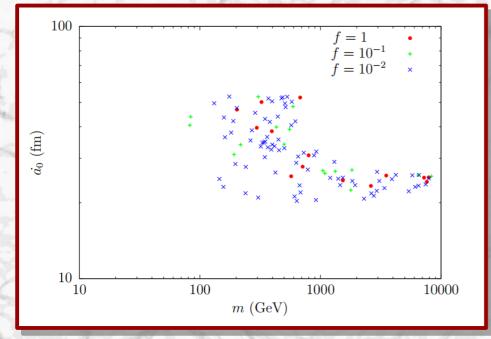
 $\Gamma_{\rm Res}^{i}(t) = \langle \Gamma(t) \rangle^{i} - \langle \langle \Gamma(t) \rangle \rangle$

Averaged over each day (annual modulation) or each hour of every day during the year (diurnal modulation) Averaged over the year

Dark antiatoms

- Dark antiproton and dark positron
- Dark photon can "convert" to ordinary photon
- Interacts strongly with "our world"
- Couple to heavy elements (Thalium in DAMA)
- Emit X-rays, which produce electron recoils







J.R. Cudell, Q. Wallemacq, JCAP 2015

"Bouncing" SIMPs

Incoming , DM

DM "fountain" Thermalizes here A THIN Trovel to the detector The Earth

However, we do not expect this to drastically change our conclusion