

The Struve-Sahade effect in the optical spectra of O-type binaries

I. Main-sequence systems* (Corrigendum)

N. Linder¹, G. Rauw^{1,**,}, H. Sana², M. De Becker^{1,***,} and E. Gosset^{1,**,}

¹ Institut d'Astrophysique et de Géophysique, Université de Liège, Bât. B5c, Allée du 6 Août 17, 4000 Liège, Belgium
e-mail: rauww@astro.uclg.ac.be

² European Southern Observatory, Alonso de Cordova 3107, Vitacura, Santiago 19, Chile

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This erratum corrects a few clerical errors in Table 1 (radial velocities of HD 165052) of our original paper (Linder et al. 2007, A&A, 474, 193).

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References

Linder, N., Rauw, G., Sana, H., De Becker, M., & Gosset, E. 2007, A&A, 474, 193

Table 1. Journal of the observations of HD 165052.

HJD	Φ	RV_1	RV_2
		(km s^{-1})	(km s^{-1})
–2 450 000			
1299.7250	0.23	88.3	–100.8
1300.7319	0.57	–56.5	76.9
1300.9264	0.63	–86.7	96.6
1301.9281	0.97	31.2	–24.9
1304.7434	0.93	7.4	–1.7
1304.7507	0.93	15.0	–12.5
1304.9309	0.99	40.1	–39.5
1323.8361	0.39	17.9	–25.8
1327.6014	0.66	–87.7	106.8
1327.9127	0.77	–93.7	103.3
1670.7601	0.78	–80.1	103.7
1671.7225	0.11	96.3	–111.3
1672.7016	0.44	–15.9	28.7
2335.8879	0.86	–41.0	56.9
2336.8791	0.19	97.5	–118.5
2337.8880	0.53	–54.9	70.1
2338.8808	0.87	–46.1	47.8
2339.8848	0.21	96.9	–111.2
2381.8324	0.40	6.6	–5.7
2382.8569	0.75	–79.7	103.3
2383.8577	0.09	95.4	–103.1

Notes. The first column lists the heliocentric Julian date (HJD), the second provides the phase as computed from our orbital solution (see Table 2 of Linder et al. 2007) and the third and fourth ones give the radial velocities used to compute the orbital solution respectively for the primary and the secondary stars.

* Based on observations made at the European Southern Observatory (La Silla, Chile) and at the Observatoire de Haute Provence (France).

** Research Associate FNRS, Belgium.

*** Postdoctoral Researcher FNRS, Belgium.