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The ultraviolet spectrum of the peculiar emission-line star GG Carinae : the line identifications (*)

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Summary. — Line features as well as the continuum of GG Carinae in the ultraviolet wavelength domain were described by Brandi *et al.* (1986) on the basis of a series of high and low resolution IUE spectra for which the line identifications are detailed here. The present contribution is thus the complement to the above-mentioned paper and enables the disentangling of the complex and rich spectrum of GG Carinae.

Key words : stars : GG Carinae — lines : identification — stars : emission-line — stars : variable — ultraviolet spectra.

1. Introduction.

Recently, Gosset *et al.* (1984, 1985) have investigated the lightcurve and the radial velocity variations of the peculiar emission-line star GG Car. We refer to those two papers for the previous works on the subject. In their attempt to derive a physical model of the object, Gosset *et al.* (1984, 1985), and Brandi *et al.* (1986) pointed out several characteristics that ought to be taken into consideration :

(a) The spectrum of GG Car exhibits permitted and forbidden emission lines and P Cyg profiles for most of the Balmer lines during the whole lightcurve. These lines must originate in an envelope around the primary star and be accelerated outwards ;

(b) Since the radial velocities show different variations from line to line, as well as between absorption and emission components of the same transition, the stratified envelope is not spherically symmetric but rather is elongated ;

(c) Absorption lines of HeI as well as a second absorption component of the Balmer lines appear around phase 0.45 ($P = 31^d.020$), at the time that a possible partial occultation (« glitch ») in the lightcurve is observed ;

(d) Smaller scale variabilities are also probably present in the main trend of the light variation observed in GG Car, thus tending to confirm that at least one of

the components could be a genuine variable (Kruytbosch, 1930) ;

(e) The radial velocities analysed in terms of a binary system give a mean eccentricity of 0.3 and a mass function of order 0.02, indicating that the star we see is the most massive of the system : these values are of course to be taken with caution, as indicated in Gosset *et al.* (1985).

On the basis of low and high resolution International Ultraviolet Explorer data, Brandi *et al.* (1986) present the first study of the star in the ultraviolet wavelength domain ; they describe and analyse the UV line spectra as well as the energy distribution. The present contribution complements that of Brandi *et al.* (1986) in that it disentangles the spectrum and gives an identification for the numerous lines (absorption, emission, P Cyg) that GG Car exhibits between 1232 and 3196 Å. A gap exists between 2100 and 2300 Å because of the severe interstellar absorption « bump » centred around 2200 Å, causing the continuum to drop to a very weak value in that zone ; further, much shorter gaps may exist corresponding to réseaux marks and/or to interorder discontinuities.

2. Observations.

The present investigation is based on three high resolution International Ultraviolet Explorer images. One, image SWP 8936, concerns the short wavelength region whereas the two others, images LWR 4920 and LWR 5741, concern the long wavelength one. A detailed log of those observations can be found in table I(a) of Brandi *et al.* (1986). The fact that the

(*) Based on observations made with the International Ultraviolet Explorer.

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phases of the different spectra are respectively 0.41, 0.38, 0.46, i.e. not very different and very near or within the « glitch » along the lightcurve (Gosset *et al.*, 1984, 1985) is to be pointed out. The identifications have been achieved with the aid of the tables of Moore (1950, 1952), Kelly and Palumbo (1973) and Kelly (1979). In addition, comparisons to the published identifications of the UV spectra of α Cygni (Barbier *et al.*, 1978) and of ζ Draconis and of τ Herculis (Underhill and Adelman, 1977) were performed.

The short wavelength spectrum presents many essentially undisplaced lines of interstellar origin which permit to set a zero velocity reference scale which in fact coincides with the IUE scale. The situation is not as clear for the long wavelength spectrum: the two observed spectra are shifted one with respect to the other, and it is impossible to attribute this shift either to a bad wavelength calibration or to the difference in phase (and thus in radial velocities) of the two observations. The image LWR 5741 leads to a spectrum for which the interstellar lines are almost undisplaced in the IUE scale. The observed wavelengths reported here are taken from this image. A line is assumed to be real if it is also present in the LWR 4920 image.

The results are given in tables I and II. Table I deals with the short wavelength spectrum and the successive columns give:

Column 1: the observed vacuum wavelength.

Column 2: a flag concerning the observed line:

E: emission line;

X: affected by a réseaux mark;

S: saturated i.e. almost zero flux for an absorption line or containing saturated pixels for an emission one;

B: blended.

Column 3: the atom or the ion responsible for the line.

Column 4: the UV multiplet number.

Column 5: an asterisk indicates a zero level line.

Column 6: the laboratory vacuum wavelength of the transition.

Column 7: remarks:

CS = circumstellar ⁽¹⁾;

IS = interstellar ⁽¹⁾;

? = identification is doubtful with respect to the observed intensity or position.

⁽¹⁾ Distinction between CS and IS is explained in section 3.4 of Brandi *et al.* (1986).

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Table II deals with the long wavelength spectrum. The columns have the same meaning as for table I except that the wavelengths are air wavelengths. It is also to be pointed out that the FeII lines that are double are, of course, all blended, but this has not been indicated in the table.

In a few areas the signal to noise ratio is poor, in others the spectrum is so crowded that the continuum level is hard to determine; furthermore many absorption lines are saturated (as indicated in the Tabs.) or nearly so. These factors do not enable us to give significant intensity values in addition to the identification. Nevertheless, due to the existence of complex features (see for example Figs. 7, 12 and 13 of Brandi *et al.*, 1986), the identification was not obvious and we think that tables I and II will be of interest for subsequent analyses.

The richness of FeII lines is worth stressing: a detailed analysis of these lines will definitely be worth performing once enough atomic data and high(er) quality observations become available.

Furthermore a better phase coverage of the UV spectra would be the obvious next step in order to perform a more detailed study of the peculiar emission-line star GG Car, e.g. concerning the UV radial velocity curve.

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J.-P. Swings has read and improved this manuscript: may he find here the mark of our gratitude.

TABLE I. — *The far ultraviolet spectrum of GG Car in the range 1230-2100 Å.*

(1)	(2)	(3)	(4)	(5)	(6)	(7)							
1232.45								1291.05					
1232.80		NI 2	-		1233.036				FE 2	87		1291.594	
1236.65		NI 2	-		1233.250				TI 3	2	*	1291.622	
1239.45	S	NI 2	-		1236.799	?			FE 2	88		1293.543	CS
1239.90	S	NI 2	-		1239.832				TI 3	1,2		1294.698	
1240.45	S	MG 2	-	*	1239.925		CS		FE 2	87		1294.914	
1242.50		MG 2	-	*	1240.395		CS		TI 3	1	*	1295.883	
1247.10	S						BROAD		FE 2	86	*	1296.088	
1249.40	S	C 3	9		1247.383		BROAD		TI 3	1	*	1295.883	CS
1250.00	S	SI 2	13.05		1250.09				SI 3	4	*	1296.30	
1250.55	S	SI 2	13.05		1250.43				S 1	9	*	1296.174	IS
1253.20	S	S 2	1	*	1250.50				TI 3	1	*	1298.30	
1253.75	S	S 2	1	*	1253.79		CS		S 1	9	*	1298.90	CS
1258.90	S	S 2	1	*	1253.79		CS		SI 3	1	*	1298.659	
1259.50	S	S 2	1	*	1259.53				TI 3	1	*	1298.891	
1259.80	S	S 2	1	*	1259.53		CS		TI 3	1	*	1298.659	CS
1260.55	S	SI 2	4	*	1260.421				TI 3	1	*	1298.659	
		FE 2	9	*	1260.542				SI 3	4	*	1298.891	
		SI 2	4	*	1260.421		CS		O 1	2	*	1302.1686	
		FE 2	9	*	1260.542		CS		O 1	2	*	1302.20	S,X
		C 1	9	*	1260.736		IS		S 1	9	*	1302.337	CS
1261.55		C 1	9	*	1261.426		IS		O 1	2	*	1302.337	
1264.20	S	C 1	9	*	1261.552		IS		S 1	9	*	1302.337	
		SI 2	4		1264.737				SI 3	4	*	1302.337	
									S 1	9	*	1302.337	
									S 1	9	*	1302.337	
									SI 3	4	*	1302.337	
									S 1	9	*	1302.337	
									S 1	9	*	1302.337	
									SI 3	4	*	1302.337	
									S 1	9	*	1302.337	
									S 1	9	*	1302.337	
									SI 3	4	*	1302.337	
									S 1	9	*	1302.337	
									S 1	9	*	1302.337	
									SI 3	4	*	1302.337	
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									SI 3	4	*	1302.337	
									S 1	9	*	1302.337	
									S 1	9	*	1302.337	
									SI 3	4	*	1302.337	

TABLE I (continued).

1842.20					1913.10	S,X			
1845.00	FE 3	117	1843.502				FE 3	57	1913.622
	FE 3	117	1845.304		1914.30		FE 3	34	1914.056
	FE 3	97	1845.521		1916.40		FE 3	51	1915.083
1846.00			1846.573		1917.80	B	FE 2	96	1917.337
1848.10	FE 2	98	1848.771				FE 2	138	1918.114
1849.00	FE 3	141	1849.960		1922.10		FE 3	57	1918.284
1851.00			1851.517				FE 2	51	1922.789
1853.85	FE 3	53	1854.716		1925.25		FE 2	138	1922.797
1854.60	FE 2	65	1854.716				FE 2	123	1925.983
	AL 3	1					FE 3	57	1926.013
1856.20	AL 3	1	1854.716	CS	1928.30		FE 3	34	1926.304
1857.80			1858.026		1929.70		FE 2	-	1929.194
1859.00	AL 2	4	1859.741		1934.60		FE 3	51	1930.387
1859.30	FE 2	65	1860.055		1935.90	X	FE 2	96	1935.296
	FE 2	97	1862.311		1936.70		FE 2	96	1936.799
1861.85	AL 2	4	1862.790		1939.50		FE 3	51	1937.345
	AL 3	1			1940.70		FE 3	61	1940.018
1862.65	AL 3	1	1862.790	CS	1942.90		FE 3	79	1941.633
1864.15			1864.656		1944.80		FE 3	51	1943.481
	FE 2	126	1864.743		1950.30	X	FE 3	61	1945.342
1865.80	FE 2	126	1866.305		1952.00		FE 3	68	1951.007
	FE 3	52	1866.554		1952.80		NI 3	24	1952.540
1869.00	FE 3	52	1869.828		1871.152		FE 3	68	1953.322
1870.30	FE 3	52	1872.515		1872.65		FE 3	61	1954.223
1871.90	FE 2	-	1874.931		1962.50		FE 2	170	1963.110
	FE 2	-	1876.838		1963.80		FE 3	61	1964.260
1874.35	FE 2	65	1877.467		1971.90		FE 2	170	1964.342
1876.20	FE 2	97					FE 3	-	1972.245
1876.70	FE 2	125					FE 3	-	
1877.20			1877.989		1972.20		FE 3	-	1972.638
1878.30	FE 3	63			1974.80		-	-	
1879.45			1880.046		1975.50		FE 3	54	1976.126
1880.40	FE 2	141	1880.620		1978.15		FE 3	54	1978.417
	FE 3	62	1880.976		1979.30		FE 3	-	1980.392
1881.80	FE 3	62	1882.047		1981.40		FE 3	54	1982.076
1884.40	FE 3	96	1885.125		1993.00		FE 3	50	1994.073
1886.00			1886.61		1994.90		FE 3	50	1995.266
	FE 3	52	1886.757		1995.80		FE 3	50	1995.563
1886.70	FE 3	53	1887.197		1998.60		FE 3	50	1996.420
1887.00	FE 3	52	1887.471				FE 2	187	1999.430
1888.10	FE 2	125	1888.733				FE 2	186	1999.462
1890.00	FE 3	52	1890.669		2000.20		FE 2	122	2001.015
1891.90	FE 3	52	1892.140		2025.35	S	ZN 2	1	2026.137
1893.40	FE 2	125	1894.006		2026.15		MG 1	2	2026.477
1894.60			1895.456		2026.50		ZN 2	1	2026.137
	FE 3	34	1895.675		2029.10		MG 1	2	2026.477
1896.30	FE 2	124	1896.803		2029.10		FE 2	93	2029.834
1897.90	FE 3	83	1898.538		2032.30		FE 2	94	2033.060
1898.40	FE 2	140	1899.318		2055.00		FE 2	109	2055.927
1900.40	FE 3	96	1901.096		2055.60		CR 2	1	2056.25
1902.10	FE 3	95	1902.902		2056.10		CR 2	1	2056.25
1905.00	FE 3	-	1905.818		2061.80		CR 2	1	2062.20
1906.20	FE 3	96	1906.457		2062.10		ZN 2	1	2062.662
1907.20	FE 3	108	1907.577		2062.75		CR 2	1	2062.20
	FE 3	83	1907.741				ZN 2	1	2062.662
1910.10	FE 3	83	1910.401						
	FE 2	124	1910.669						
					2065.40		CR 2	1	2066.12
					2066.10		CR 2	1	2066.12
							FE 2	109	2066.663
					2068.00		FE 2	137	2068.576
					2090.00		FE 3	124	2090.721
							FE 3	67	2090.806
					2093.30		FE 3	77	2094.172
					2097.35		FE 3	67	2098.149
							FE 2	80	2098.176
							FE 2	120	2098.176
							FE 3	66	2098.361

TABLE II (continued).

2472.50	B				2546.20						
2473.80		FE 2	148	2473.321				NI 2	57	2547.188	
2477.70		FE 2	208	2474.766		2547.50	B	FE 2	158	2547.338	
2479.00		FE 2	161	2478.449		2548.50	B	FE 2	158	2548.589	
2481.20		FE 2	179	2478.571				FE 2	177	2549.395	
2482.80		FE 2	179	2480.115				FE 2	177	2549.461	
		FE 2	161	2482.117		2549.70		FE 2	240	2550.027	
		FE 2	207	2482.657				FE 2	158	2550.575	
2483.30		NI 2	61	2484.204		2554.50		FE 2	240	2550.683	
		FE 2	243	2484.241				FE 2	177	2555.067	
		FE 2	400	2484.441		2559.20		FE 2	177	2555.453	
2485.40		NI 2	61	2484.204		2561.00	B,S	FE 2	205	2559.774	
2488.70		FE 2	208	2484.241		2561.50	B,S	FE 2	221	2560.281	
2489.80		FE 2	161	2484.441		2561.50	B,S	FE 2	64	2562.535	
2490.40		FE 2	179	2486.343		2562.40	B	FE 2	64	2562.535	
2492.10		FE 2	207	2489.482		2562.90	B	FE 2	64	2563.477	
2496.80		FE 2	207	2490.858		2564.90	B	FE 2	64	2563.477	
2497.90		FE 2	161	2491.396		2565.90		NI 2	62	2565.923	
2499.90		FE 2	357	2493.184		2566.20		FE 2	64	2566.912	
2501.40		FE 2	207	2493.184		2567.20		FE 2	64	2566.912	
2502.70		FE 2	161	2493.262		2569.90		FE 2	145	2568.409	
2505.20		FE 2	207	2497.819		2572.25		TI 2	9	2571.036	
2510.00		NI 2	18	2498.897		2573.50		FE 2	205	2573.211	
2513.40		FE 2	285	2500.924		2573.50		FE 2	144	2574.362	
		NI 2	61	2502.393		2574.20	E ?	FE 2	144	2574.362	
				2503.560		2576.15	S	FE 2	144	2574.362	
				2506.094		2576.50		MN 2	1	2576.105	IS
				2510.871	BROAD	2576.80		FE 2	64	2577.919	
				2511.761		2577.70	E ?	FE 2	64	2577.919	
				2514.383				FE 2	64	2577.919	
				2514.627				FE 2	64	2577.919	
2515.30		TI 3	7	2516.053		2581.55	S	FE 2	64	2582.580	
2516.10		FE 2	147	2517.131		2581.70	S	FE 2	64	2582.580	
2518.20		FE 2	268	2519.046		2582.60	E	FE 2	64	2582.580	
2520.10		FE 2	268	2521.092		2584.40	S	FE 2	1 *	2585.876	
2520.80		FE 2	330	2521.816		2584.80	S	FE 2	1 *	2585.876	
2522.00						2585.90	S	FE 2	1 *	2585.876	
2524.30	B					2587.10		FE 2	1 *	2585.876	CS
2525.20	B	FE 2	159	2525.388		2590.20		FE 2	326	2587.945	
2528.50		FE 2	145	2526.294		2590.80		FE 2	64	2591.542	
2529.50		FE 2	177	2529.549		2591.70		FE 2	64	2591.542	
2532.40	E	FE 2	177	2529.549		2591.70		FE 2	318	2592.781	
2533.30	S	FE 2	159	2533.627		2592.20		FE 2	64	2593.722	
2533.30	S	FE 2	159	2534.416		2592.70		FE 2	64	2593.722	
2534.50		FE 2	177	2535.486		2593.85		FE 2	64	2593.722	
2535.80		FE 2	159	2536.803		2596.90	S	MN 2	1 *	2593.724	IS
2537.70	S	FE 2	159	2536.845		2597.10	S	FE 2	1	2598.369	
		FE 2	158	2538.799		2597.90	S	FE 2	1	2598.369	
		FE 2	158	2538.909		2598.20	S	FE 2	1 *	2599.395	
2538.60	E	FE 2	158	2538.993		2598.20	S	FE 2	1 *	2599.395	
		FE 2	158	2538.799		2599.40	S	FE 2	1 *	2599.395	
		FE 2	158	2538.909		2605.60	S	FE 2	1 *	2599.395	CS
2539.60		FE 2	177	2540.661		2605.60	S	FE 2	1 *	2605.682	IS
		FE 2	177	2541.101		2607.10	S	FE 2	1	2607.086	
2540.60		FE 2	158	2541.836		2607.10	E	FE 2	1	2607.086	
2542.50		FE 2	159	2543.377		2610.10	B,S	FE 2	173	2611.339	
		FE 2	177	2543.430		2610.40		FE 2	1	2611.873	
2544.10		FE 2	147	2544.972		2610.60		FE 2	1	2611.873	
2545.60		FE 2	177	2546.670		2611.70	E	FE 2	1	2611.873	

TABLE II (continued).

2612.40		FE 2	1	2613.820	2724.80	E,X	FE 2	62	2724.884
2612.70		FE 2	1	2613.820	2726.10		FE 2	200	2727.383
2613.90	E	FE 2	1	2613.820			FE 2	63	2727.538
2616.40		FE 2	1	2617.618	2726.30		FE 2	200	2727.383
2616.80		FE 2	1	2617.618	2727.40	E	FE 2	63	2727.538
2617.40	E	FE 2	1	2617.618	2729.50		FE 2	63	2727.538
2619.50		FE 2	1	2620.408	2729.80		FE 2	62	2730.734
		FE 2	171	2620.695	2730.40	E	FE 2	62	2730.734
2620.20		FE 2	1	2621.669	2735.70		FE 2	62	2730.734
2620.50		FE 2	1	2621.669	2736.00		FE 2	63	2736.968
2621.50	E	FE 2	1	2621.669	2736.90	E	FE 2	63	2736.968
2622.60		FE 2	171	2623.726	2738.10		FE 2	63	2736.968
2624.20		FE 2	1	2625.667	2738.40		FE 2	63	2739.546
2624.50		FE 2	1	2625.667	2739.50	E	FE 2	63	2739.546
2625.60	E	FE 2	1	2625.667	2741.80		FE 2	63	2739.546
2627.00		FE 2	1	2628.293	2742.20		FE 2	62	2743.196
2627.20		FE 2	1	2628.293	2743.20	E	FE 2	62	2743.196
2628.00	E	FE 2	1	2628.293	2745.10		FE 2	62	2743.196
2628.70	B	FE 2	171	2629.590	2745.40		FE 2	62	2746.483
		FE 2	171	2630.071			FE 2	63	2746.978
2630.00	S,X,B	FE 2	1	2631.047	2745.70		FE 2	63	2746.978
		FE 2	1	2631.322	2746.55	E	FE 2	62	2746.483
2631.30	E	FE 2	1	2631.047			FE 2	63	2746.978
		FE 2	1	2631.322	2747.60		FE 2	63	2749.178
2663.20		FE 2	263	2664.663	2747.80	S	FE 2	63	2749.178
2663.60		FE 2	263	2664.663			FE 2	63	2749.320
2664.50	E	FE 2	263	2664.663			FE 2	63	2749.485
2665.60		FE 2	263	2666.636	2748.50		FE 2	62	2749.320
2666.00		FE 2	263	2666.636	2749.10	E	FE 2	63	2749.485
2666.60	E	FE 2	263	2666.636			FE 2	63	2749.178
2683.10		FE 2	283	2684.754	2751.90		FE 2	62	2749.320
2683.60		FE 2	283	2684.754	2752.30	S,X	FE 2	63	2749.485
2691.00		FE 2	283	2692.602	2754.20		FE 2	235	2753.287
		FE 2	62	2692.834	2754.40		FE 2	62	2755.734
2691.40		FE 2	283	2692.602	2754.40		FE 2	62	2755.734
		FE 2	62	2692.834	2755.60	E,S	FE 2	62	2755.734
2692.40	E	FE 2	62	2692.834	2760.20		FE 2	62	2761.812
2702.80		FE 2	261	2703.988	2760.60		FE 2	63	2761.812
2703.40		FE 2	261	2703.988	2766.10		FE 2	235	2767.500
2703.80	E	FE 2	261	2703.988	2766.30		FE 2	235	2767.500
2705.40		FE 2	341	2706.566	2767.20	E	FE 2	235	2767.500
2707.90		FE 2	218	2709.054	2767.80		FE 2	235	2767.500
2710.80					2767.80		FE 2	63	2768.934
2713.20					2768.00		FE 2	63	2768.934
							FE 2	200	2769.153
2713.40		FE 2	63	2714.413	2768.90	E	FE 2	63	2768.934
2714.10	E	FE 2	63	2714.413	2777.80		FE 2	234	2779.299
2715.00		FE 2	63	2714.413	2778.10		FE 2	234	2779.299
2715.60		FE 2	261	2716.217	2779.00	E	FE 2	234	2779.299
		FE 2	62	2716.701	2782.30		FE 2	337	2783.410
2715.80	E	FE 2	261	2716.217	2782.50		FE 2	234	2783.410
		FE 2	62	2716.701			FE 2	295	2783.959
2723.40		FE 2	62	2724.884	2783.60	E	FE 2	234	2783.959
2723.70		NI 2	-	2724.725	2790.10		FE 2	234	2783.691
		FE 2	62	2724.884			MG 2	3	2790.771

FE 2 (201) ?

TABLE II (continued).

2792.80		FE 2	198		2793.888	
2794.20	S	MG 2	1	*	2795.523	
2795.55	S	MG 2	1	*	2795.523	CS
2796.30	E, S	MG 2	1	*	2795.523	
2797.40		MG 2	3		2797.984	
2801.30	S	MG 2	1	*	2802.697	
2802.70	S	MG 2	1	*	2802.697	CS
2803.40	E	MG 2	1	*	2802.697	
2830.30		FE 2	217		2831.561	
2834.45		FE 2	216		2835.711	
2839.30		FE 2	217		2840.649	
		FE 2	280		2840.758	
2842.10		FE 2	294		2843.485	
2851.10		MG 1	1	*	2852.127	
2852.20	S, X	MG 1	1	*	2852.127	CS
2854.50		FE 2	196		2855.689	
		FE 2	195		2856.147	
2857.10		FE 2	279		2858.340	
		FE 2	195		2858.340	
2879.20		FE 2	61		2880.757	
2879.70		FE 2	61		2880.757	
2893.90		FE 2	257		2895.071	
2921.50		FE 2	293		2922.023	
2925.00		FE 2	60		2926.586	
2925.30		FE 2	60		2926.586	
2926.40	E	FE 2	60		2926.586	
2935.70		MG 2	2		2936.501	
2943.20		FE 2	78		2944.397	
2946.30	X	FE 2	78		2947.658	
2946.50	X	FE 2	78		2947.658	
2947.30	E	FE 2	78		2947.658	
2947.80		FE 2	78		2947.658	
2952.50		FE 2	277		2949.178	
2952.50		FE 2	60		2953.774	
2953.60	E	FE 2	60		2953.774	
2963.40		FE 2	78		2964.629	
		FE 2	78		2965.035	
2964.50	E	FE 2	78		2964.629	
		FE 2	78		2965.035	
2968.80		FE 2	60		2970.517	
2969.30		FE 2	60		2970.517	
		FE 2	276		2970.682	
2970.60	E	FE 2	60		2970.517	
2983.50		FE 2	78		2984.824	
		FE 2	78		2985.545	
2984.30		FE 2	78		2985.545	
2984.90	E	FE 2	78		2984.824	
		FE 2	78		2985.545	
3001.00		FE 2	78		3002.650	
3001.30		FE 2	78		3002.650	
3002.40	E	FE 2	78		3002.650	
3106.20	E	FE 2	V68		3106.559	
3195.75	E	FE 2	V7		3196.070	